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**MODELISATION DE SPECTRES INFRAROUGES POUR LES GAZ A EFFECT DE SERRE A  
FORT POTENTIEL DE RECHAUFFEMENT PLANETAIRE: CF<sub>4</sub> ET SF<sub>6</sub>**

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**Résumé substantiel en français**

**Avec les publications en anglaise intégrées**

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## Introduction

L'une des tâches principales de la spectroscopie moléculaire dans le domaine infrarouge est l'identification de molécules organiques et inorganiques en phase gazeuse ou liquide caractérisées par un ensemble de paramètres spectroscopiques, des fréquences et des intensités de raies, des paramètres d'élargissement, etc. L'analyse spectrale a trouvé de nombreuses applications dans de nombreux domaines, tels que l'étude des propriétés des gaz, la biophysique, l'astrophysique ou encore la physique atmosphérique. Afin d'analyser la composition de l'atmosphère, il est souhaitable d'avoir des informations les plus complètes possibles sur une large gamme spectrale. Le développement de la spectroscopie infrarouge à transformée de Fourier (FTIR) ainsi que des technologies laser a permis de faire des progrès significatifs dans l'identification des gaz. Cependant, les spectres de molécules polyatomiques « toupies sphériques » présente des densité d'états très élevés et par conséquent une grande quantité de données à extraire et à analyser. Cette tâche peut être parfois très difficile à partir de méthodes purement expérimentales. Par exemple, afin d'étudier des spectres à haute température, les listes spectroscopiques doivent inclure une grosse quantité d'information, même pour des raies très faibles. Pour cela, il est nécessaire de pouvoir calculer un nombre important de transitions, allant de quelques millions à plusieurs milliards selon la valeur maximale de  $J$ .

Avec le développement de gros calculateurs, des progrès significatifs ont été réalisés en spectroscopie théorique en se basant sur les calculs de chimie quantique. La littérature scientifique récente démontre à quel point que les listes théoriques ont maintenant une grande place dans l'étude et l'analyse de données expérimentales ainsi que dans la modélisation d'atmosphères planétaires. Par exemple, il y a eu grand d'intérêt pour les spectres infrarouges du méthane ( $CH_4$ ) au cours des dernières décennies pour l'étude des atmosphères des planètes du système solaire et de leurs satellites [1] [2] [3]. Les paramètres de bande du méthane sont alors d'une importance capitale [4] [5] [6] [7] [8]. L'amélioration des programmes *ab initio*, des algorithmes de construction d'une surface d'énergie potentielle (PES) et des méthodes de résolution du problème nucléaire permet de calculer des niveaux (ro)vibrationnels avec une précision toujours plus grande. Les progrès récents dans les calculs variationnels à partir de PES et de surfaces de moment dipolaire (DMS) pour des molécules légères telles que l'eau, l'ozone, l'ammoniac, le méthane ou l'éthylène ont mené à l'identification de nombreuses bandes vibrationnels et conduit à la modélisation des spectres de ces molécules à différentes températures. Pour les molécules plus lourdes possédant des modes vibrationnels de plus basse énergie – ce qui est l'objet de ce travail - de tels calculs deviennent plus compliqués car ils nécessitent la prise en compte de grands nombres quantiques  $V$  et  $J$ . Pour de tels systèmes, la densité d'états est très élevée, rendant encore plus compliquée les calculs à haute température. Il est donc nécessaire d'avoir à disposition des outils efficaces pour pouvoir effectuer des calculs variationnels de qualité. Notons que les listes *ab initio* contenant plusieurs milliards de transitions sont devenues courantes pour diverses applications planétaires et atmosphériques.

*Un des objectifs de cette thèse est d'obtenir des PES et DMS ab initio très précises, permettant de calculer des niveaux rovibrationnels et de modéliser les spectres de molécules qui composent l'atmosphère terrestre. Nous nous focaliserons sur  $CF_4$  et  $SF_6$ . Une analyse des spectres expérimentaux à haute résolution de la molécule  $CF_4$  en combinant des modèles effectifs spectroscopiques et les informations obtenues à l'aide de PES et DMS ab initio pour les positions et les intensités est également effectuée*

## **Objet d'étude**

L'étude des propriétés spectroscopiques des molécules de type toupie sphérique intéresse particulièrement la communauté scientifique. Tout d'abord, cela est dû à l'utilisation de ces composés dans divers domaines industriels. En même temps, ces molécules jouent un rôle actif dans les processus se produisant dans l'atmosphère terrestre. Cette thèse propose une analyse théorique et expérimentale des gaz à effet de serre fluorés :  $CF_4$  et  $SF_6$ .

Le tétrafluorométhane  $CF_4$  (tétrafluorure de carbone, CFC-14, PFC-14, R-14) est un gaz incolore, non toxique et ininflammable appartenant au groupe des perfluorocarbures (PFC - Perfluorocarbures). Les PFC sont des gaz à effet de serre (GES) extrêmement puissants et sont considérés comme l'un des gaz traces ayant la plus longue durée de vie dans l'atmosphère. Le tétrafluorométhane est le PFC le plus abondant dans la stratosphère, où sa durée de vie, selon [9], peut atteindre 50 milles ans. Cela est dû au fait que les molécules  $CF_4$  sont les moins actives par rapport aux autres fréons et ne se désintègrent pas très lentement dans l'atmosphère, ce qui conduit à leur accumulation.  $CF_4$  est une molécule hautement absorbante et possède l'un des plus grands potentiels réchauffement mondiaux (GWP) connus [10]. Malgré une source lithosphérique naturelle [11], la principale source de  $CF_4$  est anthropique et est associée à la production d'aluminium [12].

L'hexafluorure de soufre  $SF_6$ , qui possède de hautes propriétés d'isolation électrique et d'extinction d'arc, a été largement utilisé dans de nombreux domaines de l'industrie et de la médecine [13] [14] [15]. En raison des fortes bandes d'absorption dans les fenêtres de transparence atmosphérique, le  $SF_6$  est un GES puissant, dont le GWP est 29300 fois supérieur à celui du  $CO_2$  [16].  $SF_6$  est l'un des produits chimiques gazeux les plus stables avec une durée de vie atmosphérique de 3200 ans [9]. Les isotopologues les plus abondants sont  $^{32}SF_6$  (95%),  $^{34}SF_6$  (4%) et  $^{33}SF_6$  (0,7%).

Ces gaz peuvent jouer un rôle actif dans le réchauffement climatique et sont soumis aux directives du Protocole de Kyoto de la Convention-cadre des Nations Unies sur les changements climatiques (CCNUCC) pour les espèces recommandées pour la surveillance environnementale. [17] [18] [19].

## Etat de l'art

Les émissions de GES provenant de sources anthropiques ont entraîné une augmentation significative de leur concentration dans l'atmosphère, ce qui peut provoquer des changements climatiques importants en un temps relativement court. Les GES absorbent le rayonnement dans l'infrarouge moyen et lointain dans l'atmosphère terrestre, ce qui conduit à l'effet de serre. Des bases de données spectroscopiques modernes telles que HITRAN [20] et GEISA [21], contiennent des listes « raie par raie » avec de nombreuses attributions (information sur les nombres quantiques ainsi que sur un certain nombre d'autres paramètres spectroscopiques). Cependant, pour la majorité des molécules de cinq et sept atomes, l'identification n'est connue que pour les polyades inférieures, ce qui rend difficile le calcul des spectres à différentes températures.

Ces dernières années, de nombreux travaux ont été consacrés à l'étude des spectres des molécules fluorées considérées dans cette thèse. La première étude expérimentale des spectres infrarouges de  $CF_4$  a été rapportée par Bailey et al.[22] en 1938. Les mesures correspondantes couvraient une gamme de 2 à 19 mm, qui a ensuite été étendue en 1951 par Woltz et al.[23] jusqu'à 2–33 mm. Depuis, d'autres études expérimentales ont été publiées. [24] [25], couvrant l'infrarouge lointain, qui ont ensuite été complétées par une étude théorique [26].

Au cours des quatre dernières décennies, un certain nombre d'études expérimentales et théoriques ont été consacrées ont porté sur l'analyse et la prédiction des spectres de la molécule  $SF_6$  [27]. Dans la Ref. [28] l'état  $v_3 = 2$  a été étudié en utilisant deux bandes latérales accordées indépendamment d'un laser  $CO_2$ . Les intensités des vibrations fondamentales actives dans le domaine infrarouge ont également été étudiées dans les travaux [29] et [30]. De nombreux d'efforts ont récemment été menés pour analyser les spectres  $SF_6$  à haute résolution en utilisant des hamiltoniens phénoménologiques effectifs dont les paramètres ont été empiriquement ajustés aux données expérimentales [31] [32] [33] [34] [35].

Dans ce travail, nous considérons une méthode variationnelle à partir de PES pour résoudre l'équation de Schrödinger nucléaire. Cette approche a été appliquée avec succès à de nombreuses molécules à quatre et cinq atomes. [36] [37] [38] [39], ainsi que des molécules à six atomes telles que  $C_2H_4$  [40] [41] [42].

### **Résultats de ce travail.**

5 publications scientifiques à comité de lecture ont été publiées, dans lesquelles se trouvent les principaux résultats scientifiques décrits dans ce rapport.

Les résultats inclus dans le travail de thèse ont été discutés et présentés lors de conférences russes et internationales. Les travaux ont été présentés sous forme de présentations orales et posters lors des conférences suivantes :

- XXIV Colloque international « Optique atmosphérique et océanique. Physique de l'atmosphère" 2018
- XIX Symposium international sur la spectroscopie moléculaire à haute résolution 2019

- XXVII conférence "Aérosols de Sibérie"2020

*Liste des publications de l'auteur :*

1. Michaël Rey, Andrei V. Nikitin, I.S. Chizhmakova, Vladimir G. Tyuterev Understanding global infrared opacity and hot bands of greenhouse molecules with low vibrational modes from first-principles calculations: the case of  $\text{CF}_4$  // Phys.Chem.Chem.Phys. – 2018. – V. 20. – P.21008 – 21033.  
**DOI:** 10.1039/c8cp03252a
2. M. Mattoussi, M. Rey, M. Rotger, A.V. Nikitin, I. Chizhmakova, X. Thomas, H. Aroui, S. Tashkun, V.G. Tyuterev Preliminary analysis of the interacting pentad bands ( $\nu_2 + 2\nu_4, \nu_2 + \nu_3, 4\nu_2, \nu_1 + 2\nu_2, 2\nu_1$ ) of  $\text{CF}_4$  in the 1600 – 1800  $\text{cm}^{-1}$  region // J. Quant. Spectrosc. Radiat. Transfer. – 2019. – V. 226. – P. 92 – 99  
**DOI:** 10.1016/j.jqsrt.2019.01.018
3. Chizhmakova I.S., Nikitin A.V. Potential Energy Surface of  $\text{SF}_6$  // Atmospheric and Oceanic Optics, 2019, V. 32, No. 6, P. 613–618. DOI: 10.1134/S1024856019060046
4. Nikitin, A.V., Rey, M., Chizhmakova, I.S., Tyuterev, V.G. First Full-Dimensional Potential Energy and Dipole Moment Surfaces of  $\text{SF}_6$  // J. Phys.Chem. – 2020. – V. 124. – P. 7014–7023. DOI: 10.1021/acs.jpca.0c02733A
5. Rey M. Towards a complete elucidation of the ro-vibrational band structure in the  $\text{SF}_6$  infrared spectrum from full quantum-mechanical calculations / M. Rey, I. S. Chizhmakova, A. V. Nikitin, V. G. Tyuterev / Phys. Chem. Chem. Phys. – 2021. – Vol. 23. – P. 12115–12126. – DOI: 10.1039/d0cp05727d (Scopus)

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**Contenu principal du rapport**

- L'introduction décrit la pertinence du sujet et présente les objets d'étude ainsi que l'état de l'art.

## ***Chapitre 1. Fondements théoriques et méthodes de calcul***

Le premier chapitre est consacré aux méthodes *ab initio* qui sont maintenant utilisées dans beaucoup d'études. Ce chapitre contient une description de l'équation de Schrödinger et des approximations (Section 1.1) ainsi que des principes généraux pour résoudre l'équation du mouvement des noyaux (Section 1.2).

Le choix de la méthode *ab initio* détermine en grande partie la précision des calculs des niveaux et des transitions rovibrationnelles. Afin de résoudre l'équation de Schrödinger électronique, une description de la méthode de Möller-Plesset (MP) [43] et de la méthode «coupled-cluster» (CC) [44] est donnée. Une combinaison de ces méthodes avec des bases de Dunning polarisées [45] permet d'obtenir des niveaux vibrationnels avec une précision de quelques  $\text{cm}^{-1}$ .

Le développement d'une méthode variationnelle efficace est indispensable pour pouvoir modéliser des spectres à haute résolution tout en permettant de mieux cerner l'organisation des bandes froides et chaudes sur une large gamme spectrale. Des travaux consacrés à l'étude de molécules plus légères, telles que  $\text{CH}_4$ ,  $\text{SiH}_4$  ou  $\text{GeH}_4$ , [46] [47] [48] [49] [39] démontrent à quel point les calculs théoriques récents peuvent être précis.

## ***Chapitre 2. Construction de surfaces d'énergie potentielle (PES) et de moment dipolaire (DMP) précises***

Le deuxième chapitre est consacré à la description de l'approche théorique pour le calcul des PES et le DMS de  $\text{CF}_4$  et  $\text{SF}_6$ . Afin de résoudre l'équation de Schrödinger électronique, la méthode CCSD(T) est utilisée [50] avec des bases de Dunning, ce qui permet d'obtenir les niveaux vibrationnels les plus bas avec une précision de quelques  $\text{cm}^{-1}$ . Pour une molécule polyatomique non-linéaire composée de  $N$  noyaux, la PES sera de dimension  $3N-6$ . La représentation géométrique de la surface dans l'espace multidimensionnel des variables qui déterminent la configuration des noyaux a été obtenue à partir des coupes unidimensionnelles le long de chaque coordonnée de symétrie, construites sous la forme d'une fonction polynomiale.

En utilisant l'algorithme décrit dans la Ref. [51], un ensemble de tenseurs irréductibles totalement symétriques jusqu'à un certain ordre a été construit. Chaque tenseur est représenté comme la somme des produits de puissances symétrisées  $S(R,A)_C^{p_n}$ , où  $SR$  correspond aux composantes radiales,  $SA$  correspond aux composantes angulaires,  $C$  est la représentation irréductible du groupe ponctuel;  $p_n$  est le degré de la coordonnée symétrisée  $S_n$ ,  $n = 1, 2, 3, \dots$ .

Pour chaque groupe  $\{p_1, p_2, \dots, p_n\}$ , un certain ensemble de puissances symétrisées a été fixé afin de déterminer de manière unique un jeu de constantes de force par ajustement. Ensuite, un ensemble de géométries a été déterminé pour une définition non ambiguë du champ de force.

Les propriétés de symétrie d'une molécule jouent un rôle essentiel en chimie et en physique tels que la construction du moment dipolaire, la détermination des vibrations fondamentales d'une molécule et la détermination des transitions spectroscopiques autorisées. Les sous-sections 2.1.1. et 2.2.1 donnent une description théorique des molécules  $CF_4$  et  $SF_6$  avec une présentation des groupes ponctuels auxquels appartiennent les molécules étudiées regroupant les opérations de symétrie, les éléments de matrice, les tables de caractères ainsi que la forme des vibrations normales.

Le choix d'un ensemble de points *ab initio* dans l'espace des configurations nucléaires et la représentation de la PES sous la forme d'une fonction analytique symétrisée sont donnés dans les sections 2.1.3 et 2.2.4.

La fonction potentielle a été construite à partir d'opérateurs tensoriels irréductibles

$$V(R) = \sum_p F_{ip} \prod_{i\sigma} (S_{i\sigma}^{(\Gamma_i)})^{p_{i\sigma}} \quad (1)$$

où l'indice  $i$  inclut tous les termes possibles jusqu'à une certaine puissance totale  $p_{max}$ ,  $F_{ip}$  est l'ensemble des paramètres de la PES *ab initio*, et chaque tenseur  $R_i^p$  peut être obtenu comme le produit correspondant de coordonnées radiales et angulaires symétrisées.

La molécule  $CF_4$  est décrite par quatre modes vibrationnels et la PES est donc par quatre coordonnées symétrisées selon le schéma de couplage suivant :

$$R_i^p = \left( ([SR_{A_1}]^{p_1} \times [SR_A]^{p_2})^C \times ([SR_{F_2}]^{p_3} \times [SR_{F_2}]^{p_4})^C \right)^{A_1} \quad (2)$$

où  $[SR_C]^p$  est la puissance symétrisée ;  $p = p_1 + p_2 + p_3 + p_4$  est le degré total. Dans le cas de la molécule  $CF_4$ , l'équation (2) inclut tous les paramètres jusqu'au sixième ordre, complétés par 43 paramètres angulaires principaux jusqu'au huitième ordre.

Dans le cas d'une molécule  $SF_6$  à sept atomes, trois modes de vibration radiaux  $\nu_1 - \nu_4$  se transforment selon les représentations  $A_{1g}(774.5 \text{ cm}^{-1})$ ,  $E_g(643.3 \text{ cm}^{-1})$  et  $F_{1u}(948.1 \text{ cm}^{-1})$ , et les trois modes angulaires  $\nu_4 - \nu_6$  selon les représentations  $F_{1u}(615.0 \text{ cm}^{-1})$ ,  $F_{2g}(524.0 \text{ cm}^{-1})$  et  $F_{2u}(347.7 \text{ cm}^{-1})$ .

Ainsi, l'expression (2) peut être obtenue comme suit :

$$R_i^p = \left( \left( [SR_{A_{1g}}]^{p_1} \times ([SR_{E_g}]^{p_2} \times [SR_{F_{1u}}]^{p_3})^{C_{23}} \right)^{C'} \right. \\ \left. \times \left( [SR_{F_{1u}}]^{p_4} \times ([SR_{F_{2g}}]^{p_5} \times [SR_{F_{2u}}]^{p_6})^{C_{56}} \right)^{C''} \right)^{A_{1g}} \quad (3)$$

Les points *ab initio* ont été ajustés à partir de cette représentation analytique, en tenant compte de la symétrie (2) et (3) et de la fonction de poids utilisée par Schwenke et Partridge.

$$w(E) = \frac{\tanh(-0.0005(E - E_1) + 1.002002002)}{2.002002002} \quad (4)$$

Des expressions analytiques du moment dipolaire  $\mu_\alpha^{MFF}$  ont été obtenues à la fois en coordonnées internes curvilignes et en coordonnées rectilignes normales dans le repère d'Eckart lié à la molécule (MFF). Par exemple

$$\mu_\alpha^{MFF} = \sum_i \tilde{K}_i \tilde{R}_{i\alpha}^p \quad (5)$$

où

$$\tilde{R}_{i\alpha}^p = \left( \left( [q_1^{(A_1)}]^{p_1} \times [q_2^{(E)}]^{p_2} \right)^{(C_1)} \times \left( [q_3^{(F_2)}]^{p_3} \times [q_4^{(F_2)}]^{p_4} \right)^{(C_2)} \right)^{(F_2)} \quad (6)$$

est défini par les puissances successives  $p = \{p_1, p_2, p_3, p_4\}$  de tenseurs irréductibles pour les modes une fois  $\{q_1^{(A_1)}\}$ , deux fois  $\{q_2^{(E)}\}$  et trois fois  $\{q_3^{(F_2)}\}$ ,  $\{q_4^{(F_2)}\}$  dégénérés.  $C_1$  et  $C_2$  désignent les types de symétrie tétraédrique des interactions tensorielles intermédiaires appartenant aux types  $\{A_1, A_2, E, F_1, F_2\}$ , et  $\tilde{K}_i$  sont les paramètres de la DMS.

*La section 2.4 résume le chapitre 2 et liste les principaux résultats : les résultats sont également présentés dans les articles joints à ce manuscrit.:*

*Understanding global infrared opacity and hot bands of greenhouse molecules with low vibrational modes from first-principles calculations: the case of CF<sub>4</sub>.*

*Potential Energy Surface of SF<sub>6</sub>.*

*First Full-Dimensional Potential Energy and Dipole Moment Surfaces of SF<sub>6</sub>.*

## Molécule CF<sub>4</sub>

Des PES et DMS à neuf dimensions pour le tétrafluorure de carbone ont été obtenues. Un léger raffinement empirique de la géométrie d'équilibre et de quatre paramètres quadratiques du PES, calculés au niveau de la théorie CCSD(T)/aug-cc-pVQZ-DK, a donné un écart quadratique moyen de  $\sim 0,001 \text{ cm}^{-1}$  pour les énergies de rotation et une très bonne précision pour tous les centres de bandes vibrationnelles connues expérimentalement jusqu'à  $3000 \text{ cm}^{-1}$ .

Au final, 298 paramètres sur les 330 ajustables ont été déterminés. L'ajustement de la PES a été effectué de manière fiable par la méthode des moindres carrés à partir de 19 882 points *ab initio* jusqu'à  $15\,000 \text{ cm}^{-1}$ . En effet, le rapport  $\frac{N_{param}}{N_{data}} = \frac{1}{65}$  avec un écart-type de  $0,24 \text{ cm}^{-1}$  rend cet ajustement assez fiable. Concernant le moment dipolaire calculé au niveau VTZ, 604 paramètres (sur un total de 680 paramètres au sixième ordre) ont été déterminés à partir de la même grille, avec un écart type de  $0,00001$  Debye.

## Molécule SF<sub>6</sub>

La PES *ab initio* finale de la molécule SF<sub>6</sub> a été construite à l'ordre 5 à partir de 8708 configurations nucléaires calculées au niveau CCSD(T)/CVQZ-DK. L'écart type est de 0,5 cm<sup>-1</sup> pour 4283 géométries. Sur les 260 paramètres de la DMS au quatrième ordre, 194 ont été statistiquement bien définis, l'écart type de 1,4 × 10<sup>-5</sup> u.a. . Les résultats obtenus démontrent à quel point une représentation en termes d'opérateurs tensoriels irréductibles permet d'obtenir des expressions analytiques compactes pour les PES et DMS à 9 et 15 dimensions qui seront utilisées dans la modélisation de spectres infrarouges.

## Chapitre 3. Construire des listes théoriques

Le troisième chapitre est consacré à la construction de listes *ab initio* pour les molécules (CF<sub>4</sub>) et (SF<sub>6</sub>). Les calculs quantiques incluant tous les degrés de liberté représentent une tâche extrêmement difficile et nécessitent l'utilisation de méthodes optimisées afin (i) de converger tous les calculs (ii) de diagonaliser de très grandes matrices (iii) de tenir compte de toutes les transitions nécessaires pour des calculs d'opacité fiables.

La section 3.1 décrit le code de calcul TENSOR écrit par M. Rey en FORTRAN90. Ce logiciel est organisé autour de deux axes principaux : NMMod et NMVar. Ce « package » est consacré à l'étude de molécules semi-rigides non-linéaires appartenant à un groupe ponctuel de symétrie quelconque et a récemment été étendu aux systèmes non-rigides.

- NMMod construit l'Hamiltonien de Watson et le moment dipolaire sous forme tensorielle à partir des PES et DMS *ab initio*, et d'un ensemble de coordonnées normales symétrisées {*q*}.
- NMVar calcule les solutions propres de l'équation de Schrödinger stationnaire par calcul variationnel. Un jeu de fonctions de bases dites « réduites » est introduit pour résoudre le problème rovibrationnel. La procédure pour construire un spectre moléculaire est décrite schématiquement sur la figure 1

**CF<sub>4</sub>** : La simulation globale des spectres à haute résolution de CF<sub>4</sub> dans la gamme 0-3000 cm<sup>-1</sup> est présentée dans Ref. *Understanding global infrared opacity and hot bands of greenhouse molecules with low vibrational modes from first-principles calculations: the case of CF<sub>4</sub>* [52]. Une comparaison de l'absorption théorique avec les spectres expérimentaux PNNL [53] à moyenne résolution, enregistrés à des températures de 5, 25 et 50 °C à une pression de 1 atm dans une large gamme spectrale de 570–6500 cm<sup>-1</sup> a permis de valider les calculs (voir Fig. 2). Les spectres PNNL ne sont pas attribués, donc leurs paramètres de raie ou de bande n'ont pas été extraits, mais ces données brutes fournissent une mine d'informations expérimentales précieuses pour tester les prédictions théoriques. Toutes les comparaisons dans cette section ont été faites à 296 K. On a montré qu'un quasi-continuum formé de raies très faibles pouvait apparaître dans les spectres

infrarouges de  $CF_4$  même à température ambiante en raison d'une forte augmentation de la densité des transitions. Afin d'effectuer des calculs d'opacité à température ambiante, il a fallu aller jusqu'à  $J = 80$ , et considérer plus de deux milliards de transitions.

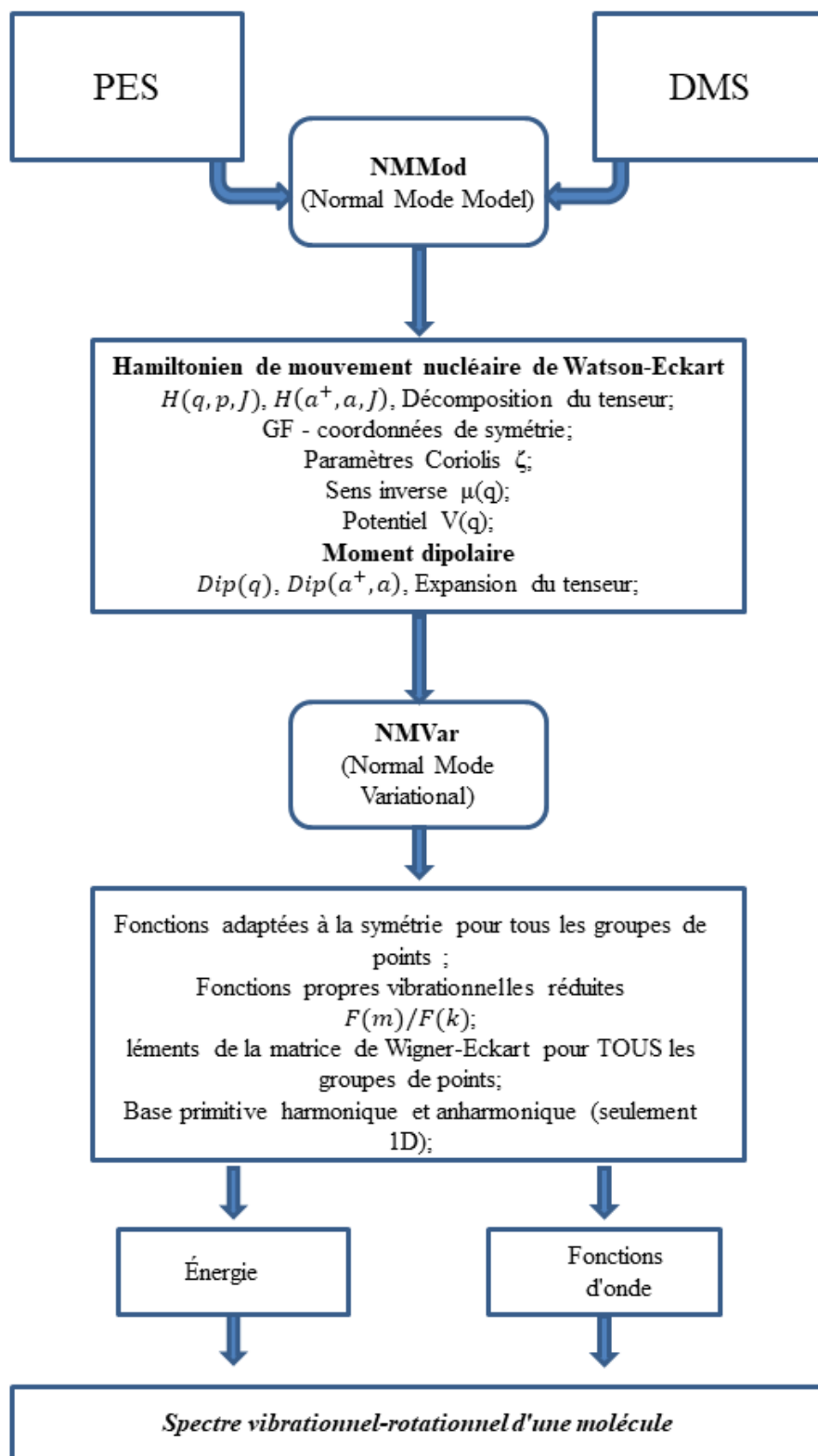


Fig. 1 Description schématique du code TENSOR

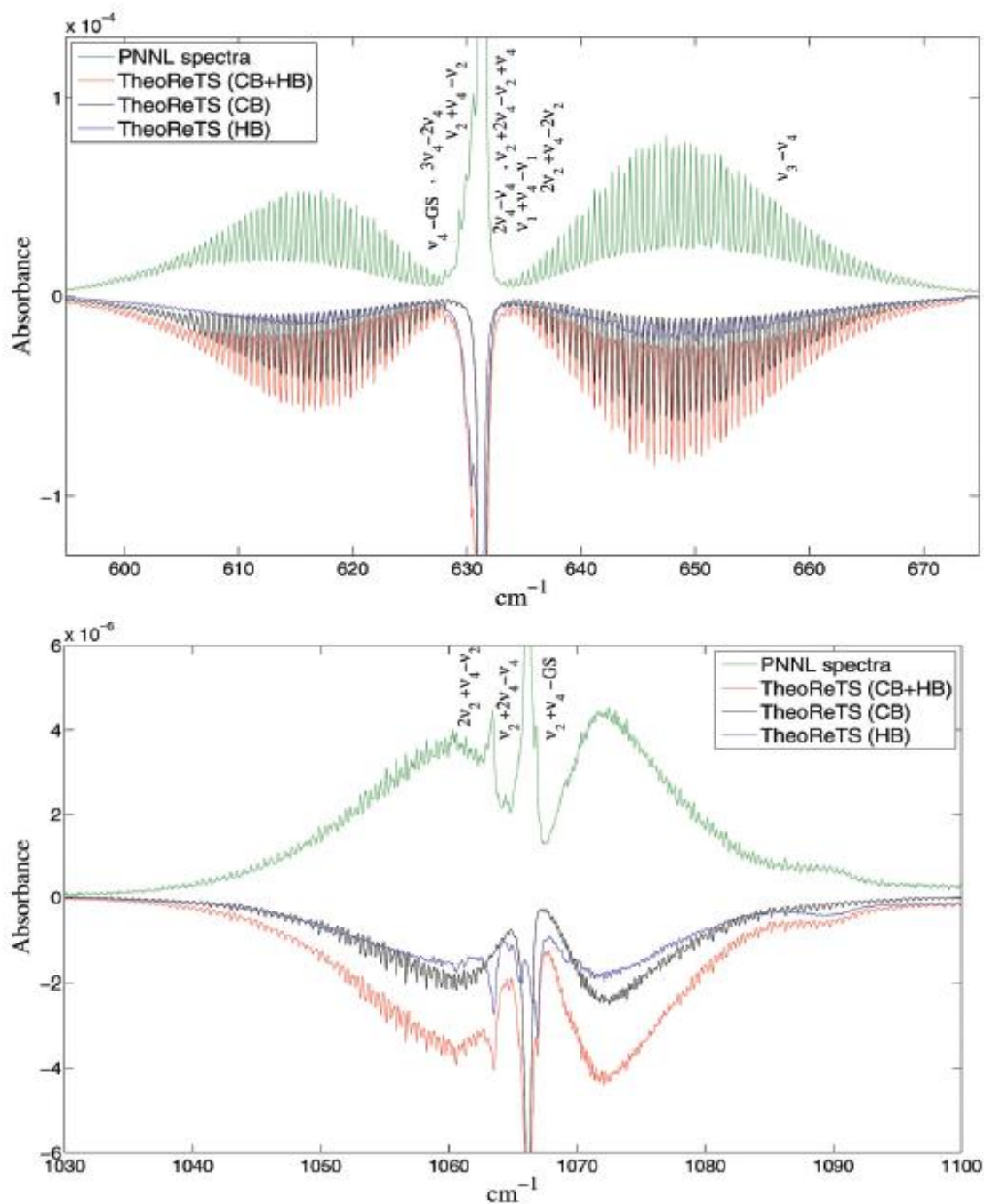


Figure 2 Un exemple de comparaison des spectres d'absorption théorique de  $\text{CF}_4$  inclus dans la base de données de Reims-Tomsk (TheoReTS) avec des enregistrements PNNL à température ambiante dans la plage de  $1850\text{--}2200\text{ cm}^{-1}$  (transitions CB=cold band, HB=hot band et CB + HB).

**$\text{SF}_6$**  : Concernant la molécule  $\text{SF}_6$ , des états rotationnels très excités ont été calculés jusqu'à  $J = 121$  dans la Réf. *Towards a complete elucidation of the ro-vibrational band structure in the  $\text{SF}_6$  infrared spectrum from full quantum-mechanical calculations* [54], ce qui a permis de prédire plus de 6 milliards de transitions attribuées à plus de 500 bandes froides et chaudes. La figure 3 donne un exemple de comparaison entre les

spectres PNNL et les calculs théoriques, en séparant les contributions provenant de bandes froides et de bandes chaudes [53]. Ces données seront prochainement incluses dans GEISA sous la forme de sections efficaces. Toutes les simulations de spectres ont été effectuées à une résolution de  $0,1 \text{ cm}^{-1}$  et une pression de 760 Torr en utilisant une longueur de trajet optique de 20 cm et un profil de Voigt  $0,06 \text{ cm}$  par atmosphère.

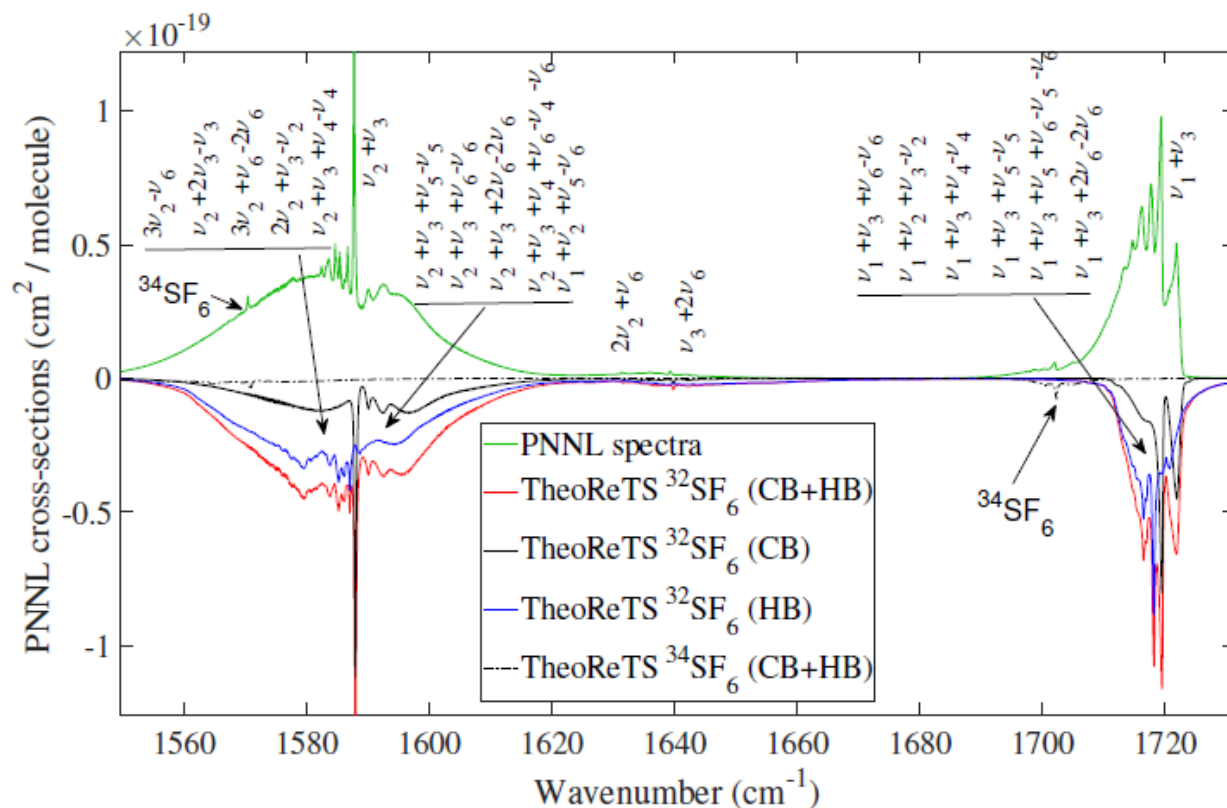


Figure 3 Comparaison des spectres d'absorption théoriques  $^{32}\text{SF}_6 + ^{34}\text{SF}_6$  inclus dans la base de données de Reims-Tomsk (TheoReTS) aux spectres PNNL expérimentaux à moyenne résolution enregistrés à température ambiante dans la gamme  $1010\text{-}1730 \text{ cm}^{-1}$  (transitions CB=cold band, HB=hot band et CB + HB).

Les intensités absolues des raies ont été calculées et des listes théoriques « globales » construites pour les 3 espèces les plus abondantes :  $^{32}\text{SF}_6$ ,  $^{33}\text{SF}_6$  et  $^{34}\text{SF}_6$ , — en utilisant nos PES et DMS. Nos spectres calculés à température ambiante reproduisent les spectres observés avec une précision jamais atteinte pour une molécule à sept atomes.

Afin d'accélérer les simulations de spectres, les données obtenues ont été divisées en deux parties :

- (a) des listes de « raies fortes » avec des niveaux d'énergie inférieurs pour simuler à différentes températures ;
- (b) des bibliothèques compressées de "superlines" composées de transitions très faibles, facilitant la modélisation du quasi-continuum. Le nombre de transitions est ainsi réduit par 1 ou 2 ordres de grandeur.

## Chapitre 4. Analyse du spectre de la molécule $CF_4$ dans la gamme 2160–2210 $cm^{-1}$

Le quatrième chapitre est consacré à l'analyse du spectre de la molécule  $CF_4$  dans la gamme 2160–2210  $cm^{-1}$ . La complexité des spectres  $CF_4$  dans la gamme 2160–2210  $cm^{-1}$  est due à la présence des transitions froides et chaudes, ainsi qu'à la présence de bandes à trois, quatre et cinq quanta. L'absorption la plus forte dans cette région est associée à la bande  $\nu_1 + \nu_3$  avec une intensité intégrée de  $\sim 5,0 \times 10^{-19} cm/molécule$ . Les bandes  $\nu_1 + \nu_2 + \nu_3 - \nu_4$  et  $\nu_1 + 2\nu_4$  sont cinq fois plus faibles que cette bande, et  $\nu_1 + \nu_3 + \nu_4 - \nu_4$ ,  $\nu_1 + \nu_2 + 2\nu_4 - \nu_2$ ,  $\nu_1 + 3\nu_4 - \nu_4$ ,  $\nu_1 + 2\nu_2 + \nu_3 - 2\nu_2$  et  $2\nu_2 + \nu_3$  est un ordre de grandeur plus faible. Afin d'effectuer une analyse précise des spectres expérimentaux dans cette gamme, une "approche mixte" a été utilisée, combinant Hamiltonien effectif par polyade obtenu à partir de la PES par transformations de contact [99] et intensités *ab initio*. Cette approche permet de caractériser de manière précise les termes de résonance et par la même occasion de lever certaines ambiguïtés dans la détermination des paramètres des modèles purement empiriques [100].

L'analyse raie par raie du spectre expérimental reste une tâche très difficile. L'Hamiltonien effectif adapté à la structure en polyade de  $CF_4$  est défini comme la somme des contributions correspondant aux différents polyades inférieures  $P_k$  ( $k = 0, 1, \dots, k$ ), où  $P_0$  est l'état vibrationnel fondamental. Le modèle utilisé est défini comme :

$$H = \sum H_k = \tilde{H}_{P_0} + \tilde{H}_{\nu_2} + \tilde{H}_{\nu_4} + \tilde{H}_{Dyad} + \tilde{H}_{\nu_2+\nu_4} + \tilde{H}_{Tetrad} + \tilde{H}_{Dyade2} + \tilde{H}_{Pentade}. \quad (7)$$

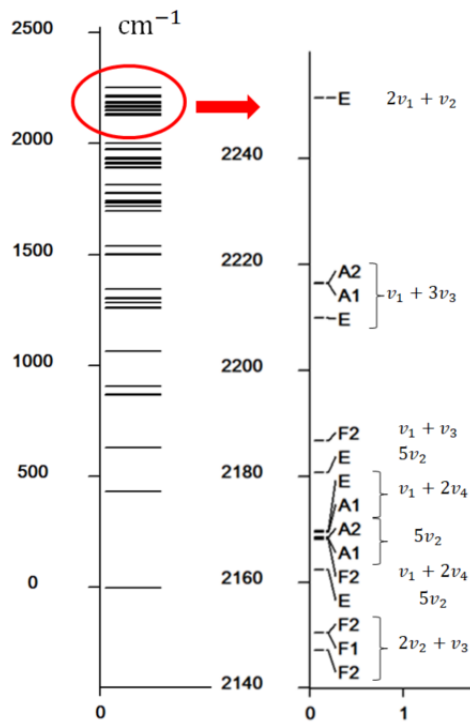


Figure 4 Schéma des niveaux vibrationnels des polyades de la molécule  $CF_4$  (à gauche) et des sous-niveaux vibrationnels (à droite) de l'hexade d'une partie du système correspondant aux bandes vibrationnelles analysées dans ce travail. La bande vibrationnelle  $2\nu_1 + \nu_2$  a un sous-niveau de symétrie vibrationnelle de

type E ; la bande  $\nu_1 + 3\nu_2$  a trois sous-niveaux ( $A_1, A_2, E$ ) ; la bande  $\nu_1 + \nu_3$  a un sous-niveau ( $F_2$ ) ; la bande  $5\nu_2$  a trois sous-niveaux ( $A_1, A_2, E$ ) ; la bande  $\nu_1 + 2\nu_4$  a trois sous-niveaux ( $A_1, F_2, E$ ) ; la bande  $2\nu_2 + \nu_3$  a dix sous-niveaux ( $F_1, F_2$ ).

Afin d'obtenir des résultats avec une précision comparable à celle expérimentale, l'Hamiltonien effectif a été développé jusqu'au 6ème ordre. La modélisation du spectre, l'attribution des transitions ainsi que le processus d'ajustement ont été effectués à l'aide des logiciels MIRS et SpectraPlot basés sur une comparaison itérative semi-automatique des spectres calculés (bâtons) et expérimentaux. Dans la Réf. *Preliminary analysis of the interacting pentad bands ( $\nu + 2\nu, \nu + \nu, 4\nu, \nu + 2\nu, 2\nu$ ) of CF in the 1600 – 1800  $\text{cm}^{-1}$  region* [106], 48 des 2188 paramètres de l'hamiltonien effectif du sixième ordre ont été ajustés sur les spectres dans la gamme 1600–1800  $\text{cm}^{-1}$ , dans cette thèse 11 autres paramètres de l'hamiltonien effectif (tableau 2) ont été ajustés sur les spectres dans la plage 2160–2210  $\text{cm}^{-1}$

Tableau 2. Valeurs des paramètres ajustés de l'hamiltonien effectif pour  $\text{CF}_4$

|    | Nomenclature tensorielle |           | Paramètres ajusté |
|----|--------------------------|-----------|-------------------|
|    | Rotation                 | Vibration |                   |
| 1  | R 0 (0, 0A1)             | 1010 1010 | -1,0529640674E+01 |
| 2  | R 1 (1, 0F1)             | 1010 1010 | -7,8977801035E-03 |
| 3  | R 2 (0, 0A1)             | 1010 1010 | -1,5039331401E-05 |
| 4  | R 1 (1, 0F1)             | 1010 1002 | 2,0455767619E-03  |
| 5  | R 0 (0, 0A1)             | 1010 1002 | 5,8729130379E-01  |
| 6  | R 1 (1, 0F1)             | 1010 1002 | -8,8290674940E-04 |
| 7  | R 0 (0, 0A1)             | 1010 0210 | 8,6692581509E-02  |
| 8  | R 1 (1, 0F1)             | 1010 0210 | -1,4030889467E-03 |
| 9  | R 1 (1, 0F1)             | 1010 0210 | -1,0127001972E-04 |
| 10 | R 0 (0, 0A1)             | 1010 0210 | 9,9141412734E-04  |
| 11 | R 1 (1, 0F1)             | 1010 0210 | -1,6383775845E-07 |

L'écart type est de 8,6243E-04, ce qui peut être considéré comme un résultat très satisfaisant.

L'ajustement de 8 paramètres du moment dipolaire effectif sur les intensités *ab initio* a permis d'obtenir les intensités pour cinq des six bandes de  $\text{CF}_4$  avec une erreur de 1,1 %. Les statistiques de cet ajustement sont présentées dans le tableau 3.

Tableau 3. Valeurs des paramètres de moment dipolaire effectif pour CF<sub>4</sub>

|   | Nomenclature tensorielle  |           | Paramètres ajusté |
|---|---------------------------|-----------|-------------------|
|   | Rotation                  | Vibration |                   |
| 1 | R 0 (0, 0A <sub>1</sub> ) | 0000 0210 | 3,0677673220E-04  |
| 2 | R 0 (0, 0A <sub>1</sub> ) | 0000 0210 | -7,0097729574E-05 |
| 3 | R 0 (0, 0A <sub>1</sub> ) | 0000 1002 | 3,8582053784E-04  |
| 4 | R 0 (0, 0A <sub>1</sub> ) | 0000 1010 | 7,4121145986E-03  |
| 5 | R 1 (1, 0F <sub>1</sub> ) | 0000 1010 | 1,3437943401E-07  |
| 6 | R 2 (0, 0A <sub>1</sub> ) | 0000 1010 | 9,8005076990E-09  |
| 7 | R 2 (2, 0E )              | 0000 1010 | 1,8003881577E-08  |
| 8 | R 2 (2, 0F <sub>2</sub> ) | 0000 1010 | -3,0670244770E-08 |

Nous observons dans la figure 5 une vue d'ensemble du spectre expérimental et du spectre calculé dans les mêmes conditions expérimentales avec le logiciel MIRS.

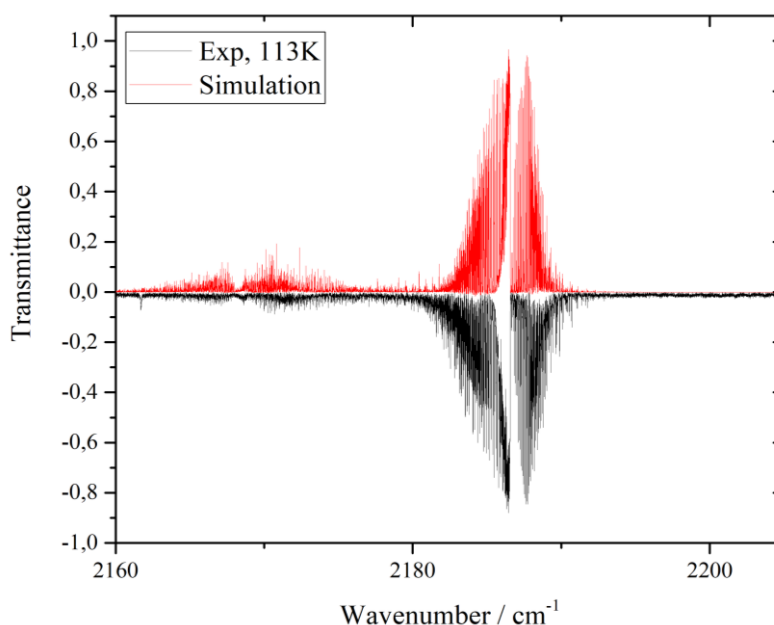


Figure 5 Comparaisons entre les spectres expérimentaux et la simulation théorique de CF<sub>4</sub> à 113K dans la gamme spectrale entre 2160 et 2205 cm<sup>-1</sup>.

*L'objectif principal de ce travail a été de construire des listes très précises pour CF<sub>4</sub> et SF<sub>6</sub> qui seront incluses dans les futures versions des bases de données HITRAN et GEISA.*

## Conclusion

Cette thèse présente les recherches menées par l'auteur jusqu'en 2022. Voici le travail qui a été effectué : une PES et une DMS à 9 dimensions ont été construites pour  $\text{CF}_4$  à l'aide de calculs *ab initio* au niveau CCSD(T) sur une grille composée de 19882 points. De manière similaire, une PES et une DMS à 15 dimensions ont été construites pour la molécule  $\text{SF}_6$  dans son état électronique fondamental.

Nous avons obtenu des listes *ab initio* très précises et les plus complètes possibles dans les gammes contenant les bandes d'absorption les plus fortes. Des calculs variationnels poussés basés sur l'introduction de bases réduites ont été effectués et ont permis la prédiction de plusieurs centaines de nouvelles bandes. Par exemple, pour  $\text{SF}_6$  les premiers spectres théoriques à haute résolution et à température ambiante dans la gamme  $300\text{--}3000\text{ cm}^{-1}$  ont été calculés pour les trois isotopologues les plus abondants ( $^{32}\text{SF}_6$ ,  $^{33}\text{SF}_6$  et  $^{34}\text{SF}_6$ ). Concernant  $\text{CF}_4$ , les positions et intensités de transition de l'hexade dans la gamme  $2160\text{--}2210\text{ cm}^{-1}$  à une température de 113 K ont été analysées. Voici les bandes en question :  $2\nu_1 + \nu_2$ ,  $\nu_1 + 3\nu_2$ ,  $\nu_1 + \nu_3$ ,  $5\nu_2$ ,  $\nu_1 + 2\nu_4$  et  $2\nu_2 + 2\nu_4$ .

Dans tous nos calculs, les intensités *ab initio* calculées permettent de rivaliser avec les meilleures mesures expérimentales. Contrairement aux modèles effectifs, des extrapolations à différentes températures sont beaucoup plus simple et plus fiable à partir de calculs variationnels. Les données théoriques sont incluses dans la base de Reims-Tomsk TheoReTS (<http://theorets.univ-reims.fr>), (<http://theorets.tsu.ru>). Ce travail servira de base lors des prochaines analyses de spectres expérimentaux à haute résolution.

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# ***Annexe***

Understanding global infrared opacity and hot bands of greenhouse molecules with low vibrational modes from first-principles calculations: the case of CF<sub>4</sub>

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# Understanding global infrared opacity and hot bands of greenhouse molecules with low vibrational modes from first-principles calculations: the case of CF<sub>4</sub><sup>†</sup>

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Fluorine containing molecules have a particularly long atmospheric lifetime and their very big estimated global warming potentials are expected to rapidly increase in the future. This work is focused on the global theoretical prediction of infrared spectra of the tetrafluoromethane molecule that is considered as a potentially powerful greenhouse gas having the largest estimated lifetime of over 50 000 years in the atmosphere. The presence of relatively low vibrational frequencies makes the Boltzmann population of the excited levels important. Consequently, the “hot bands” corresponding to transitions among excited rovibrational states contribute significantly to the CF<sub>4</sub> opacity in the infrared even at room temperature conditions but the existing laboratory data analyses are not sufficiently complete. In this work, we construct the first accurate and complete *ab initio* based line lists for CF<sub>4</sub> in the range 0–4000 cm<sup>−1</sup>, containing rovibrational bands that are the most active in absorption. An efficient basis set compression method was applied to predict more than 700 new bands and subbands *via* variational nuclear motion calculations. We show that already at room temperature a quasi-continuum of overlapping weak lines appears in the CF<sub>4</sub> infrared spectra due to the increasing density of bands and transitions. In order to converge the infrared opacity at room temperature, it was necessary to include a high rotational quantum number up to  $J = 80$  resulting in 2 billion rovibrational transitions. In order to make the cross-section simulation faster, we have partitioned our data into two parts: (a) strong & medium line lists with lower energy levels for calculation of selective absorption features that can be used at various temperatures and (b) compressed “super-line” libraries of very weak transitions contributing to the quasi-continuum modelling. Comparisons with raw previously unassigned experimental spectra showed a very good accuracy for integrated absorbance in the entire range of the reported spectra predictions. The data obtained in this work will be made available through the TheoReTS information system (<http://theorets.univ-reims.fr>, <http://theorets.tsu.ru>) that contains *ab initio* born line lists and provides a user-friendly graphical interface for a fast simulation of the CF<sub>4</sub> absorption cross-sections and radiance under various temperature conditions from 80 K to 400 K.

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## 1 Introduction

The radiative properties of relatively heavy long-lived greenhouse gases are much less studied compared with those of principal

atmospheric absorber lighter species such as water, carbon dioxide or methane. Among these, the fluorine containing molecules (CF<sub>4</sub>, C<sub>2</sub>F<sub>4</sub>, NF<sub>3</sub>, SF<sub>6</sub>, *etc.*), though being yet minor impurities, have a particularly long atmospheric lifetime and their very big estimated global warming potentials are expected to rapidly increase in the future.<sup>1</sup> Their common feature is a presence of relatively low vibrational frequencies that makes the Boltzmann population of the excited levels important. Consequently, the “hot bands” corresponding to transitions among excited rovibrational states contribute significantly to the opacity in the infrared even at room temperature conditions. This results in very crowded rovibrational spectra that are difficult to resolve line-by-line. Only a few bands of these

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molecules have been analyzed in limited wavenumber ranges and the high-resolution spectral line or band parameters in spectroscopic databases such as HITRAN<sup>2</sup> or GEISA<sup>3</sup> are much sparser than those for lighter greenhouse species. This work is focused on the global theoretical prediction of infrared spectra of the CF<sub>4</sub> molecule. Tetrafluoromethane (CF<sub>4</sub>, also denoted as Freon-14, PFC-14, halocarbon 14, carbon tetrafluoride, *etc.*) has the largest estimated lifetime of over 50 000 years in the atmosphere for freons. CF<sub>4</sub> currently present in the atmosphere is predominantly anthropogenic due to fugitive emissions from the aluminium industry and semiconductor manufacturing.<sup>4,5</sup> Together with other perfluorocarbons it is considered as a potentially powerful greenhouse gas that could play an active role in global climate change<sup>6,7</sup> and falls under the directives of the Kyoto Protocol<sup>8</sup> among species recommended for environmental monitoring. To this end the temperature dependence of its contribution to the radiation budget in the troposphere and stratosphere is required, but a significant part of this information is still lacking despite considerable efforts in laboratory experiments and analyses (see Section 2). Purely empirical spectroscopic models are very useful for spectral assignments but limited both in frequency and temperature coverage particularly for hot band contributions. Existing laboratory data are not sufficiently complete for the CF<sub>4</sub> opacity evaluations in the infrared. Recently, progress in *ab initio* calculations and variational predictions for rovibrational spectra from potential energy surfaces (PES) and dipole moment surfaces (DMS) for lighter species such as water,<sup>9,10</sup> ozone,<sup>11–13</sup> ammonia<sup>14–16</sup> methane<sup>17–24</sup> or ethylene<sup>25–28</sup> have permitted a breakthrough in resolving completeness issues and understanding the opacity distributions of these species for various temperature conditions. The new generation of *ab initio* born databases giving access to extensive line lists containing billions of transitions have become the current reference data for various planetary applications (see examples in ref. 29–31 and references therein). For heavier molecules with lower lying vibrational modes such as CF<sub>4</sub>, such predictions are even more challenging because they require bigger *V* and *J* quantum numbers to be included. At a given temperature, this results in a much larger density of spectral bands and transitions requiring efficient methods for the convergence of variational calculations. In this paper, we report the first large-scale variational predictions for CF<sub>4</sub>, including cold, hot, overtone and combination bands, which are the most active in IR absorption up to six vibrational quanta. We show that already at room-*T* a “quasi-continuum” (QC) of overlapping weak lines appears in CF<sub>4</sub> infrared spectra due to the increasing density of bands and transitions that must be accounted for in the converged opacity. A similar QC has been recently reported for methane spectra, both in theory<sup>32–34</sup> and in observations,<sup>35</sup> but at much higher temperatures of about 1000 K. The paper is structured as follows. The state-of-the art of CF<sub>4</sub> infrared spectroscopy is briefly overviewed in Section 2. Electronic structure calculations and fine tuning of the CF<sub>4</sub> PES using available experimental data are described in Section 3. Section 4 is devoted to the techniques of variational calculations, symmetrization and basis set convergence issues.

The full calculated ro-vibrational line list (Section 5) is partitioned into strong & medium intensity transitions and the quasi-continuum of individually indistinguishable weak lines. The latter ones are compressed using “super-lines” (SL) databases for an efficient modelling of absorption cross-sections. Validation *versus* raw experimental spectra is considered in Section 6. Final sets of line parameters and SL libraries are freely available *via* the TheoReTS information system<sup>29</sup> (<http://theorets.univ-reims.fr/> or <http://theorets.tsu.ru/>). Knowledge of excited ro-vibrational states of CF<sub>4</sub> is also of importance for such fundamental phenomena as clustering of high-*J* level patterns and related symmetry effects.<sup>36,37</sup>

## 2 State-of-the-art in the spectroscopy of the tetrafluoromethane

<sup>12</sup>CF<sub>4</sub> is a spherical top molecule with tetrahedral (*T<sub>d</sub>*) point group symmetry at the equilibrium geometry similar to <sup>12</sup>CH<sub>4</sub>. But in contrast to methane, the relation between the stretching ( $\nu_1(A_1)$ ,  $\nu_3(F_2)$ ) and bending ( $\nu_2(E)$ ,  $\nu_4(F_2)$ ) normal mode frequencies of CF<sub>4</sub> makes the polyad structure much less regular, *e.g.* the resonance scheme for fundamental frequencies  $\omega_1 : \omega_2 : \omega_3 : \omega_4$  of methane can be simply approximated as  $\approx 2 : 1 : 2 : 1$  while that of CF<sub>4</sub> is given by  $\approx 4 : 2 : 6 : 3$ . This more complicated polyad scheme and the high molecular symmetry give rise to numerous degeneracies and quasi-degeneracies and to vast and complicated resonance interactions due to inter-mode couplings. Because of the overlapping polyads, many sets of overtone and combination bands are not really isolated from nearby ones. For this reason, empirical effective models are more difficult to apply than for CH<sub>4</sub>, and the line-by-line analyses are much less advanced.

Another challenge concerns the modelling of hot bands (HB) that strongly contribute to the absorption, even at room temperature. CF<sub>4</sub> is heavier than methane and has two low-lying bending frequencies at 440 and 640 cm<sup>−1</sup> leading to a huge number of IR-active transitions belonging to the hot bands. An account of these hot bands using empirical effective spectroscopic models is an issue that has not been yet solved. This is why, except for very few studies, most of the analyses of cold band (CB) transitions has been carried out during the last four decades.

The first experimental study of infrared CF<sub>4</sub> spectra was reported by Bailey *et al.*<sup>38</sup> in 1938. The corresponding measurements covered the range 2–19  $\mu\text{m}$  which was then extended in 1951 by Woltz *et al.*<sup>39</sup> to 2–33  $\mu\text{m}$ . After that, other experimental studies were published<sup>40,41</sup> covering the far-infrared region, complemented later on by a theoretical study.<sup>42</sup> High-resolution spectra of the  $\nu_4$  and  $\nu_2 + \nu_4$  bands have been recorded by McDowell *et al.*<sup>43</sup> with one of the absorption features identified as belonging to the hot band  $\nu_2 + \nu_4 - \nu_2$ . McDowell *et al.*<sup>44–46</sup> also carried out the laser spectroscopy of CF<sub>4</sub>. Rotational transitions within the (0010) vibrational state were investigated using double-resonance experiments by Takami<sup>47,48</sup> and at nearly the same time both the analysis

of the  $Q$  branch of  $2\nu_1 + \nu_4$  up to  $J = 40$  and of the  $\nu_1$  were reported by Pine.<sup>49</sup>

Raman spectra, with underlying  $\nu_1 + \nu_2 - \nu_2$  hot band (HB) transitions, have been studied by Esherick *et al.*<sup>50</sup> The  $\nu_3$  band was reinvestigated using proton energy-loss spectroscopy by Maring *et al.*<sup>51</sup> followed by line width measurements of the  $R(29)$  transitions in the combination band  $\nu_2 + \nu_4$  by Hartemink and Godfried.<sup>52</sup> A simultaneous analysis of the ground state (GS) and the bands  $\nu_3$  and  $2\nu_3$  was performed by Gabard *et al.*<sup>53</sup> using the tetrahedral formalism described by Champion<sup>54</sup> and Zhilinskii *et al.*<sup>55</sup> Boujut *et al.*<sup>56</sup> have studied the energies of  $\nu_3$  and  $\nu_4$  overtones using previous work<sup>57</sup> and the initial values of the parameters from TDS<sup>58</sup> and STDS<sup>59</sup> (Tetrahedral and Spherical Top Database System). More recently, a study of the  $\nu_3/2\nu_4$  was made by Boudon *et al.*<sup>60</sup> also using the tetrahedral formalism of the Dijon group, whereas the subsequent work by Boudon *et al.*<sup>61</sup> was devoted to a simultaneous analysis of the  $\nu_1$ ,  $2\nu_1 - \nu_1$ ,  $\nu_2$ ,  $2\nu_2$  and  $3\nu_2 - \nu_2$  bands using high-resolution stimulated Raman spectroscopy. The most extensive empirical model to date was constructed very recently by Carlos *et al.*<sup>62</sup> for 17 rovibrational bands, including some HBs, where the reader can find other references to previous spectroscopic works. Note that only line positions have been fitted in the empirical model of ref. 62. For the cross-section simulations old values of two dipole moment derivatives from ref. 63 were used.<sup>60</sup>

As stated in ref. 62 “no further dipole moment refinement is possible, since no individual line intensities measurements are available, due to the spectroscopic congestion for this molecule, for which almost all observed lines result from the overlap of several transitions”. Lacking information on line-by-line Einstein radiative coefficients for most of the hundreds of bands contributing to the IR absorption is one of the major drawbacks of a purely empirical approach.

### 3 Potential and dipole moment surfaces

Several empirical and theoretical studies have been devoted to the determination of the equilibrium nuclear configuration<sup>60,62,64,65</sup> and of the harmonic  $\text{CF}_4$  frequencies.<sup>64</sup> Cubic and partially quartic force-fields have been fitted to experimental spectral parameters by Jeannotte *et al.*<sup>66</sup> and Brodersen.<sup>67</sup> Previous studies in this domain have been reviewed in detail by Wang *et al.*,<sup>64</sup> who refined *ab initio* harmonic frequencies via Van Vleck perturbation theory using experimental vibrational band centers up to  $2500\text{ cm}^{-1}$ . They have applied a “mixed approach” combining empirically adjusted harmonic frequencies and *ab initio* cubic and quartic anharmonic constants. The latter ones were computed in a similar way to that described in the *ab initio* work of Lee *et al.*<sup>68</sup> for the methane force field using the Coupled Cluster approach with all the single and double substitutions from the Hartree–Fock reference determinant augmented by a perturbative treatment of triple excitations, CCSD(T). It was also concluded<sup>64</sup> that the T1 diagnostic (which is the measure of the importance of multi-reference effects)

**Table 1** *Ab initio* and empirical values of the equilibrium C–F bond length for  $^{12}\text{CF}_4$ . An optimized value was obtained in this work (=TW), see Section 3.2

| Method/basis                | $r(\text{C–F})$ (in Angstrom) | Ref. |
|-----------------------------|-------------------------------|------|
| CVTZ-DK                     | 1.3171                        | TW   |
| CVQZ-DK                     | 1.3153                        | TW   |
| VTZ                         | 1.3192                        | TW   |
| VQZ                         | 1.3172                        | TW   |
| CVTZ                        | 1.3171                        | TW   |
| CVQZ                        | 1.3153                        | TW   |
| CVTZ-F12                    | 1.3153                        | TW   |
| CVQZ-F12                    | 1.3147                        | TW   |
| CVQZ                        | 1.3153                        | TW   |
| ACVQZ                       | 1.3161                        | TW   |
| Empirical                   | 1.31588                       | 60   |
| Empirical                   | 1.31486                       | 62   |
| Empirical                   | 1.31526                       | 64   |
| wCVTZ                       | 1.3163                        | 65   |
| wCVQZ                       | 1.3152                        | 65   |
| Exp.                        | 1.3151                        | 65   |
| Optimized value (empirical) | 1.315349                      | TW   |

was only 0.0119 for  $\text{CF}_4$  making the CCSD(T) method an appropriate choice.

#### 3.1 Electronic structure calculations

As in previous electronic structure calculations<sup>64,65,69</sup> on  $\text{CF}_4$ , we use here the CCSD(T) method,<sup>70</sup> which has been found to provide a reasonably good description of the equilibrium configuration  $r_e$  and for the fundamental frequencies. For the atomic basis we employed the standard correlation consistent polarized valence basis sets cc-pVTZ and cc-pVQZ of Dunning and co-workers,<sup>71</sup> where T and Q stand for triple and quadruple cardinal numbers. The abbreviated notations VTZ and QTZ will be used for these sets in what follows. Wang *et al.*<sup>64</sup> have shown that with the VTZ basis set all fundamental frequencies except for the  $\nu_3$  band were close to experimental values. They have extended the basis set to “VTZ + 1” for the anharmonic force field calculations, where the suffix “+ 1” denoted the addition of a single  $d$ -function to accommodate polarization effects. Accurate equilibrium structures of fluoro- and chloroderivatives of methane have been found in ref. 65 where the *ab initio* structure was optimized with the CCSD(T) all electrons being correlated and the correlation-consistent polarized weighted core-valence triple- and quadruple-zeta basis sets<sup>72</sup> (cc-pwCVTZ and cc-pwCVQZ) leading for the latter set to  $r_e = 1.3152\text{ Å}$  of  $\text{CF}_4$  quoted in Table 1.

Thanks to increasing computational resources, it is now possible to compute a full 9-dimensional PES on a large grid of geometries and not just a cubic or quartic force field near the equilibrium. To this end, we used in this work a similar technique of grid generation as described in our recent studies on methane PES<sup>73</sup> where it was possible to achieve a “spectroscopic accuracy” for *ab initio* calculations over a large range of energies. For  $\text{CF}_4$  with 42 electrons an inclusion of all fine correlations considered in the  $\text{CH}_4$  case is currently not feasible. As a compromise between accuracy and computational cost, we have chosen the VQZ basis set to compute electronic energies on 19 882 nuclear geometries. The grid of geometries

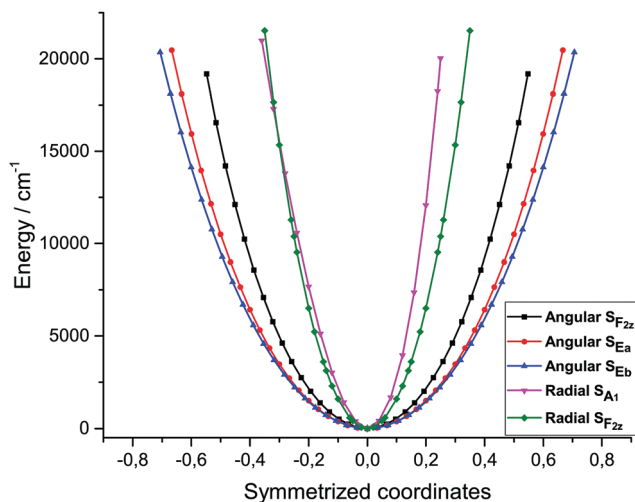


Fig. 1 Characteristic sections of  $\text{CF}_4$  *ab initio* PES along dimensionless symmetrized coordinates.

is obtained with the same method as described for the methane PES<sup>73</sup> where symmetry coordinates sample the PES at energy layers of 1000, 2500, 4000, 7000, 10 000, 12 500, and 15 000  $\text{cm}^{-1}$ . The VQZ basis reduces significantly the *ab initio* error in the  $\nu_3$  band center with respect to VTZ whereas other fundamental frequencies remain consistent with experiment (as shown in the next Section 3.2). This also improves the equilibrium geometry from  $r_e = 1.319$  Å (VTZ) to  $r_e = 1.317$  Å (VQZ) in better agreement with cc-pwCVQZ of ref. 65 that are in the range  $r_e = 1.315$ – $1.316$  Å. For completeness, we have also checked the optimization of the equilibrium geometry with core-valence correlations, including scalar relativistic Douglas-Kroll (DK)<sup>74</sup> contributions, and also with explicitly correlated F12 basis sets. This did not produce significant changes in  $r_e$  (see Table 1) whereas the corresponding ansatz would be too expensive for the full surface.

The same analytical PES representation of the 9D PES was used as for other tetrahedral molecules  $\text{CH}_4$ <sup>18</sup> and  $\text{GeH}_4$ <sup>75</sup> in terms of an expansion in the symmetrized powers  $\mathbf{p} = \{p_1, \dots, p_{4\sigma}\}$  of the  $\{S_1^{(A_1)}, S_2^{(E)}, S_{3\sigma'}^{(F_2)}, S_{4\sigma''}^{(F_2)}\}$  curvilinear coordinates expressed *via* Morse-type variables for bond stretches and cosine-type variables for angles<sup>18</sup>

$$V(\mathbf{R}) = \sum_{\{\mathbf{p}\}} \mathbf{F}_{\mathbf{p}} \prod_{i\sigma} \left( S_{i\sigma}^{(\Gamma_i)} \right)^{p_{i\sigma}}. \quad (1)$$

Here  $A_1$ ,  $E$ , and  $F_2$  denote non-degenerate, two-fold and three-fold degenerate irreducible representations (symmetry types) of the  $T_d$  point group, the indices  $\sigma$ ,  $\sigma'$  and  $\sigma''$  stand for the components of degenerate vibrations and  $\mathbf{F}_{\mathbf{p}}$  for the set of the PES parameters fitted to *ab initio* electronic energies. We have included in eqn (1) all parameters up to the sixth-order complemented by principal angular parameters up to the eighth-order. Altogether 298 parameters (of 330 adjustable ones for all four modes) were statistically well determined in the least-squares fit to 19 882 *ab initio* points up to  $E_{\text{max}} \sim 15\,000$   $\text{cm}^{-1}$ . The ratio  $N_{\text{param}}/N_{\text{data}} = 1/65$  with the RMS deviation of  $0.24$   $\text{cm}^{-1}$

Table 2 Harmonic oscillator frequencies and some band centers (in  $\text{cm}^{-1}$ ) of  $^{12}\text{CF}_4$  computed from the purely *ab initio* PES and from our empirically refined PES (TW). First dipole moment derivatives (in Debye) are quoted at the lower panel

|            | VQZ PES (TW) | Refined PES (TW) | Ref. 64 |
|------------|--------------|------------------|---------|
| $\omega_1$ | 921.16       | 920.85           | 921.57  |
| $\omega_2$ | 439.91       | 439.72           | 439.91  |
| $\omega_3$ | 1310.81      | 1305.36          | 1303.01 |
| $\omega_4$ | 638.04       | 637.18           | 637.89  |

|                       | VQZ PES (TW) | Refined PES (TW) | Exp.     |
|-----------------------|--------------|------------------|----------|
| $\nu_1(A_1)$          | 909.35       | 909.025          | 909.072  |
| $\nu_2(E)$            | 435.59       | 435.410          | 435.399  |
| $\nu_3(F_2)$          | 1289.12      | 1283.770         | 1283.664 |
| $\nu_4(F_2)$          | 631.93       | 631.086          | 631.059  |
| $2\nu_2(A_1)$         | 868.40       | 868.012          | 867.906  |
| $\nu_2 + \nu_4(F_2)$  | 1067.23      | 1066.201         | 1066.122 |
| $\nu_2 + \nu_4(F_1)$  | 1067.78      | 1066.758         | 1066.698 |
| $2\nu_4(F_2)$         | 1262.36      | 1260.593         | 1260.43  |
| $2\nu_4(A_1)$         | 1263.55      | 1261.861         | 1261.809 |
| $2\nu_4(E)$           | 1263.87      | 1262.189         | 1262.112 |
| $\nu_1 + \nu_4(F_2)$  | 1540.67      | 1539.457         | 1539.3   |
| $2\nu_1 + \nu_4(F_2)$ | 2447.04      | 2445.456         | 2445.596 |
| $2\nu_3(F_2)$         | 2572.79      | 2562.175         | 2561.912 |
| $2\nu_3(E)$           | 2580.93      | 2570.223         | 2570.013 |

|                              | VQZ DMS (TW) |   | Exp. <sup>63</sup> |
|------------------------------|--------------|---|--------------------|
| $ \partial\mu/\partial q_3 $ | 0.491        | — | 0.473              |
| $ \partial\mu/\partial q_4 $ | 0.077        | — | 0.0702             |

made the fit quite robust. This PES is given in the ESI.† Characteristic PES sections along several symmetrized coordinates are shown in Fig. 1. Harmonic frequencies and band centers up to second overtones and combination bands are collected in Table 2.

There are very few results available in the literature for a theoretical evaluation of dipole moment parameters of  $\text{CF}_4$ . To our knowledge, only two first-order derivatives  $d\mu/dq_3$ ,  $d\mu/dq_4$  computed using the density functional method have been reported by Papousek *et al.*<sup>63</sup> who also cited previous *ab initio* estimations of ref. 76. Note that our value of 0.077 D for  $|d\mu/dq_4|$  is in quite large disagreement with that obtained from the density functional theory ( $|d\mu/dq_4| = 0.0563$  D, see Table 17 of ref. 63).

No full-dimensional DMS has been published for  $\text{CF}_4$ . Accurate knowledge of this vector quantity is mandatory for the predictions of line and band intensities (see Section 5). We derived the 9D DMS components using the external electric field *ab initio* method. This is more demanding than PES as it requires six calculations for symmetry breaking nuclear displacement for each point of the DMS components. The dipole moments were computed as the derivatives of the energy with respect to the weak external uniform electric field using the finite difference scheme with the field variation of 0.0001 a.u. around the zero field strength in a two step procedure. First VTZ dipole moment calculations were carried out on the full grid of the same 19 882 nuclear geometries as for the PES.

Analytical DMS expansions of the same order as for the PES were fitted to *ab initio* values both in the curvilinear internal

and in rectilinear normal mode coordinates. The final DMS results were parametrized in the molecular fixed Eckart frame (MEF) dipole moment  $\mu_{\alpha}^{\text{MEF}}$  expansion

$$\mu_{\alpha}^{\text{MEF}} = \sum_i \tilde{K}_i \tilde{R}_{i\alpha}^p,$$

where

$$\tilde{R}_{i\alpha}^p = \left( \left( \left[ q_1^{(A_1)} \right]^{p_1} \times \left[ q_2^{(E)} \right]^{p_2} \right)^{(C_1)} \times \left( \left[ q_3^{(F_2)} \right]^{p_3} \times \left[ q_4^{(F_2)} \right]^{p_4} \right)^{(C_2)} \right)^{(F_2)}_{\alpha} \quad (2)$$

using couplings ( $\times$ ) of successive powers  $p = \{p_1, p_2, p_3, p_4\}$  of irreducible  $T_d$  tensors corresponding to one  $\{q_1^{(A_1)}\}$ , two-fold  $\{q_2^{(E)}\}$  and three-fold  $\{q_3^{(F_2)}, q_4^{(F_2)}\}$  degenerate normal mode vibrations. Here  $C_1$  and  $C_2$  denote tetrahedral symmetry species of intermediate tensor couplings belonging to  $\{A_1, A_2, E, F_1, F_2\}$  types,  $\alpha = x, y, z$  are the components of the three-fold degenerate  $F_2$  species and  $\tilde{K}_i$  are DMS parameters. We use the form (2) of the DMS because our TENSOR code, which is employed for intensity calculations, was programmed in this representation. The reader can find technical details of the formalism in ref. 77 and 78 devoted to the molecules of the same  $T_d$  symmetry. The normal-mode tensor representation is also compatible with the MIRS code<sup>79</sup> that was used for many recent analyses of methane and phosphine experimental spectra. 604 parameters (of the total 680 sixth-order ones) were statistically well determined in the fit using eqn (2) to VTZ dipole moment values on the full grid of 19 882 geometries with the RMS deviation of 0.00001 Debye.

As a next step, we computed the dipole moments with a larger atomic basis VQZ using the same scheme but on a reduced grid of 1020 geometries below  $E_{\text{max}} = 5000 \text{ cm}^{-1}$ . The dipole differences (VQZ-VTZ) were fitted by a fourth order expansion (2) using 103 adjusted parameters. This was used as a correction to the sixth-order VTZ DMS surface. The final DMS surface is provided in the ESI.<sup>†</sup>

### 3.2 Refining quadratic forces constants and geometry

The force constant matrix  $F$ , together with the geometry-based Wilson<sup>80,81</sup> matrix  $G$ , are used to determine the vibrational frequencies and the associated normal modes. The force constant matrix  $F$  directly follows from *ab initio* force field calculations but the eigenvalues of the  $GF$  matrix usually differ from the experimental measurements by some wavenumbers, typically  $1\text{--}10 \text{ cm}^{-1}$  for multi-electron molecules depending on the electronic structure calculation method. Here we start with the *ab initio* VQZ equilibrium geometry and held fixed the  $G$  matrix elements in what follows. The force constants computed at the *ab initio* CCSD(T)/VQZ level will be denoted by  $F_0$  and the corresponding vibrational frequencies by  $\Lambda_0$ . Assuming that symmetry coordinates are used, the number of independent force constants for each symmetry block is  $n_s(n_s + 1)/2$ , where  $n_s$  is the number of vibrational modes of symmetry type  $s$  such that  $\sum_s n_s = 3N - 6$ . In this section, we briefly describe a simple

method for an empirical optimization of an *ab initio* force field. This method is quite general and could be easily applied for other molecules for a fast “fine tuning” of the equilibrium geometries and quadratic force constants.

Let  $\Lambda$  be of a set a “refined” harmonic frequencies. We intend to solve the inverse harmonic eigenvalue problem written as

$$GFL = LA \quad (3)$$

to extract an optimized  $F$  matrix. At this stage, the difficulty when solving eqn (3) is twofold. First, this equation is valid only in the harmonic approximation whereas the true vibrational frequencies are always anharmonic; that is the harmonic frequencies  $\Lambda$  are not observed quantities. Second, this is a typical ill-defined problem since we have to find  $n_s(n_s + 1)/2$  unknown values from  $n_s$  data. One way to circumvent the first problem is to assume that the major part of the deviation  $\Delta$  between the anharmonic variationally computed band centers and the experimental values originates from the error in the harmonic frequencies, that is  $\Lambda \approx \Lambda_0 + \Delta$ . To regularize the ill-defined problem (second issue), we apply supplementary constraints. A simple transformation of eqn (3) permits obtaining an explicit expression for the  $F$  matrix following ref. 82

$$F = G^{-1/2} C A C^{-1} G^{-1/2} \quad (4)$$

where  $C = G^{-1/2} L$  is an arbitrary orthogonal matrix chosen to be eigenvectors of the  $(G^{-1/2} F G^{-1/2})$  matrix. For each symmetry block,  $C$  is related to the representations of the rotation group  $SO(n_s)$  that can be parametrized by means of  $n_s(n_s - 1)/2$  parameters (generators)  $\{\varphi_{ij}\}$  describing successive rotations in space. To avoid obtaining non-physical solutions of eqn (4), we consider  $F_0$  as a first guess and constrain the force constants in the  $F$  matrix to have a minimal deviation from  $F_0$ . Mathematically speaking, the physical solutions are obtained as the minimum of the functional  $\|F - F_0\|$ , where  $\|\cdot\|$  is the Euclidean norm.

For the optimization of the geometry ( $r_e$ ), we have proceeded as follows. We first compute the ground vibrational state  $J = 1$  rotational levels using the *ab initio* geometry  $r_e^{\text{ab}}$  and the force constants  $F_0$ . The variational method is fully converged and numerically exact in this case. Then we compare this to the “observed” value ( $E_{J=1} = 0.382386 \text{ cm}^{-1}$ ) computed from the STDS empirical parameters.<sup>83</sup> Assuming that the obs.–calc. error mainly originates from the rotational constant  $B_e = 6.321611/m_r r_e^2$ , we can deduce an improved  $r_e$  as an initial guess. By proceeding iteratively, we obtain simultaneously an optimal set of  $\{r_e^{\text{opt}}, F_{ij}\}$  values.

As a simple illustration, let us consider the case of the two  $F_2$ -type normal modes of  $\text{CF}_4$ . We can parametrize the  $2 \times 2$  rotation matrix  $C \in SO(2)$  by means of one angle  $\varphi$ . After optimizing  $r_e$  and solving the equation  $\partial\|F - F_0\|/\partial\varphi = 0$ , we finally obtained

$$F_0^{(F_2)} = \begin{pmatrix} 0.3168 & 0.1675 \\ & 0.2374 \end{pmatrix} \rightarrow F^{(F_2)} = \begin{pmatrix} 0.3157 & 0.1684 \\ & 0.2355 \end{pmatrix}$$

$$\Lambda_0^{(F_2)} = (638.75, 1311.22) \rightarrow \Lambda^{(F_2)} = (637.18, 1305.36)$$

$$r_e^{\text{ab}} = 1.317228 \rightarrow r_e^{\text{opt}} = 1.315349$$

where all force constants are given in a.u., the  $A$ 's in  $\text{cm}^{-1}$  and  $r_e$  in Angstroms. We give in Tables 1 and 2 the optimized geometry  $r_e$  and the harmonic oscillator frequencies  $\{\omega_i\} \equiv A$  for each irreducible representation  $\Gamma$ .

Our optimized  $r_e^{\text{opt}}$  value can be compared to those obtained from different studies (see Table 1). In comparison with the STDS database, we have estimated errors in the rotation energies of the vibration ground state errors as  $\sim 10^{-5} \text{ cm}^{-1}$  at  $J = 1$ ,  $\sim 10^{-4} \text{ cm}^{-1}$  at  $J = 20$  and  $\sim 10^{-2} \text{ cm}^{-1}$  at  $J = 50$ . From the different  $r_e$  values, we have computed the corresponding  $B_e$  values and estimated rotational errors with respect to this work (TW), taken as reference values. We obtained  $\Delta(\text{TW-ref. 64}) = 0.01 \text{ cm}^{-1}$ ,  $\Delta(\text{TW-ref. 60}) = -0.07 \text{ cm}^{-1}$ ,  $\Delta(\text{TW-ref. 62}) = 0.06 \text{ cm}^{-1}$  for  $J = 20$ , 0.06,  $-0.4$  and  $0.36 \text{ cm}^{-1}$  for  $J = 50$  and 0.13,  $-0.8$  and finally  $0.7 \text{ cm}^{-1}$  for  $J = 70$ . We clearly see, according to Table 1, that our levels would be much closer to those obtained from the pure or refined *ab initio* studies<sup>64</sup> while the two  $r_e$ 's deduced from high resolution analyses are respectively over- and underestimated and give larger discrepancies.

## 4 Theoretical background of quantum nuclear motion calculations

### 4.1 Hamiltonian model

If all vibrations are of a small amplitude and confined around a single well-defined minimum, the normal mode representation of the Hamiltonian turns out to be efficient. It generally provides an adequate description of the nuclear motions for all semirigid molecules, at least up to half way to the dissociation. The commonly used vibration-rotation Hamiltonian for a nonlinear polyatomic molecule in normal coordinates was formulated in the Eckart molecular frame in its most compact form given by Watson,<sup>84</sup> following previous studies of Wilson and Howard and Darling and Dennison. It will be simply referred to as the Eckart-Watson Hamiltonian (EWH). For a molecule with  $N$  atoms, the EWH is written in terms of dimensionless symmetry-adapted normal coordinates as

$$H^{\text{EWH}} = \frac{1}{2} \sum_{k\sigma} \omega_k p_{k\sigma}^{2(\Gamma)} + \frac{1}{2} \sum_{\alpha\beta} (J_\alpha - \pi_\alpha) \mu_{\alpha\beta} (J_\beta - \pi_\beta) + U(q) + U_W, \quad (5)$$

where  $\alpha = x, y, z$  are the Cartesian components in the Eckart frame. For a  $\text{XY}_4$  type molecule,  $k = 1, \dots, 4$  labels the four normal modes and  $\sigma$  is the component of a degenerate vibration of symmetry  $\Gamma$  ( $A_1, E, F_2$ ) in the  $T_d$  point group so that  $\dim(p_{k\sigma}) = \dim(q_{k\sigma}) = 3N - 6 = 9$ . Here  $J_\alpha$  and  $\pi_\alpha$  are dimensionless molecular frame components of the total and vibrational angular momentum, respectively.  $V(\mathbf{q})$  is the potential function converted to normal coordinates  $\mathbf{q}$  from the internal PES (1)

of Section 3.1 and  $U_W(\mathbf{q}) = -\left(\sum_{\alpha} \mu_{\alpha\alpha}(\mathbf{q})\right)/8$  is the Watson correction term.<sup>84</sup> The reciprocal inertia tensor  $\mu \equiv \mu(\mathbf{q})$  is a  $3 \times 3$  matrix related to the  $\Omega_{\alpha\beta}^i$  coefficients (see eqn (23) of ref. 85 for the definition) as<sup>86</sup>  $\mu(\mathbf{q}) = \sum_i \tilde{\Omega}^i \mathbf{m}_i^{-1} \Omega^i$ . For a

spherical top molecule like  $^{12}\text{CF}_4$  that possesses three equal principal equilibrium moments of inertia, it reads

$$\mu = I_0^{-1} \left( I_3 + I_0^{-1} \sum_{i\sigma} A_{i\sigma} q_{i\sigma} \right)^{-2} \quad (6)$$

$$I_0 = 2 \sum_i m_i a_{\alpha}^i a_{\alpha}^i, \quad \alpha = x, y, z, i = 1, \dots, N = 5.$$

where the  $a_{\alpha}^i$ 's are the Cartesian coordinates at the equilibrium configuration and  $A_{i\sigma}$  is given in ref. 86. The equilibrium rotational constant  $B_e$  for  $^{12}\text{CF}_4$  is given by  $B_e (\equiv \mu_{\alpha\alpha}^e/2) = 6.321611/m_y r_e^2 = 0.1923219 \text{ cm}^{-1}$ .

We now define a so-called reduced Hamiltonian expressed in terms of creation ( $a_{k\sigma}^{+(\Gamma)}$ ) and annihilation ( $a_{k\sigma}^{(\Gamma)}$ ) operators which are more convenient to handle the algebraic expressions involved following the method described in ref. 77 and 78. As has been shown in previous studies, this permits a drastic reduction of the number of terms in the Hamiltonian by two or three orders of magnitude, with a minimal loss of accuracy and to get rid of some artefacts inherent to high-order Taylor expansions. Following the procedure described elsewhere (see ref. 77, 85, 87 and 88 for more details), the  $p$ -order polynomial expansion of EWH in normal coordinates was reduced to a polynomial expansion of a lower order  $p'$ , such that  $p' < p$  (see Appendix B for a discussion). The reduced ( $p \rightarrow p'$ ) EWH can be written in a normal ordering form as follows:<sup>77</sup>

$$H_{\text{red}}^{\text{EWH}} = \sum_{\text{all indices}} h_{\mathbf{m},\mathbf{n}}^{n_x, n_\beta} \left( \prod_{k\sigma} \left( a_{k\sigma}^{+(\Gamma)} \right)^{m_{k\sigma}} \left( a_{k\sigma}^{(\Gamma)} \right)^{n_{k\sigma}} \right) (J_\alpha)^{n_x} (J_\beta)^{n_\beta}, \quad (7)$$

where  $n_x + n_\beta \leq 2$  and  $\{\mathbf{m}, \mathbf{n}\} = \{m_1, n_1; m_{2a}, n_{2a}, \dots\}$  is a set of  $9 \times 2$  vibrational indices. This sum-of-product Hamiltonian is thus a non-linear function of rotation-vibration operators. By defining  $\{T_i\}$  the space of operators spanned by the  $T_i$ 's where each operator  $T_i$  is associated with a specific non-linear function ( $\prod a^{+m} a^n J_\alpha J_\beta$  e.g.  $a_1^+ \equiv T_1$ ,  $a_2 \equiv T_2, \dots$ ,  $a_{3a}^+ a_{4x}^3 J_x \equiv T_j$  and so on, and  $\{h_i\}$  the set of parameters, the reduced Hamiltonian (7) can be cast into the compact vector-vector product

$$H_{\text{red}}^{\text{EWH}} = (h_1 h_2 \dots) \begin{pmatrix} T_1 \\ T_2 \\ \vdots \end{pmatrix} = \tilde{\mathbf{h}} \mathbf{T}, \quad (8)$$

where  $\tilde{X}$  is the transpose of  $X$ . As is well-known, it is highly advantageous to exploit molecular symmetry to reduce the cost of computation of the nuclear motion problem because linear combinations among some  $h_i$  parameters associated with the different components of a two or three dimensional vibrational mode are automatically accounted for in a more compact form. To obtain a "factorized" form of the Hamiltonian with linearly independent parameters, the tensor formalism proposed by Nikitin *et al.*<sup>89</sup> turns out to be very appropriate. It was initially designed for spherical top effective models and was recently extended to nuclear *ab initio* Hamiltonians

in ref. 77 and 79. In this representation, the reduced ( $p \rightarrow p'$ ) EWH can be written as

$$H_{\text{tens}}^{\text{EWH}}(p \rightarrow p') = \sum_{\text{all indices}} t_{(\text{lmn})}^{\Omega(K,I)} \left( {}^{\varepsilon} V^{(I)} \otimes R^{\Omega(K,I)} \right)^{(A_1)}, \quad (9)$$

where  $V$  and  $R$  stand for the vibrational (defined in terms of the creation/annihilation operators) and rotational (defined in terms of  $J_x, J_y$  and  $J_z$ ) tensors, respectively. At this stage the main point is to relate the non-empirical tensor parameters  $t$  to the *ab initio* parameters  $h$  in (7). This can be done in a straightforward manner by re-writing (9) as

$$H_{\text{tens}}^{\text{EWH}} = \tilde{\mathbf{t}} \mathbf{U} \mathbf{T}, \quad (10)$$

where  $\mathbf{U}$  is a  $\dim(t) \times \dim(h)$  group symmetry transformation matrix built by expanding (9) from  $T_d$  coupling (Clebsch–Gordan) coefficients. By sorting the  $\mathbf{T}$  vector in (10) in the same way as in eqn (8) and by equating to the Hamiltonians (7) and (9) expanded at the same order, the set  $\mathbf{t}$  is obtained by solving the over-determined system of equations

$$\tilde{\mathbf{U}} \mathbf{t} = \mathbf{h}, \quad (11)$$

using e.g. parallelized MKL libraries. Note that some values of  $\dim(h)$  and  $\dim(t)$  up to order 10 are given respectively in the second and third column of Table 3 of ref. 77.

Note that though the EWH – based on a locally-defined Taylor expansion – was initially designed for studying vibrational states “not far from” the bottom of the potential well, the combination of the Eckart frame/normal mode formalism with the reduced PES minimizing artefacts of the Taylor series proved to be very efficient to accurately characterize cold/hot methane<sup>28,32,34,90</sup> and ethylene<sup>28</sup> spectra, even for high wavenumber ranges.

## 4.2 Symmetry-adapted primitive basis functions

For an  $N$  atomic nonlinear molecule, the simplest way to build a multi-dimensional vibrational basis consists in forming a direct product of  $3N - 6$  primitive 1D functions as

$$|v_1 v_2 \dots v_{3N-6}\rangle \equiv \phi_{v_1}(q_1) \phi_{v_2}(q_2) \dots \phi_{v_{3N-6}}(q_{3N-6}). \quad (12)$$

Without symmetry considerations, if we assume for example that 10 functions (even more are often necessary) are taken for each degree of freedom to converge properly the energy levels, then the matrix to be diagonalized will be of the size  $D_0 \times D_0 = 10^{3N-6} \times 10^{3N-6}$ . When adding the rotational part, the size of Hamiltonian matrix will be scaled by an additional factor  $2J + 1$ . It is thus clearly seen that computing and/or storing eigensolutions from direct or even iterative eigensolvers becomes an extremely computationally demanding problem for molecules with more than four atoms.

If now we exploit the  $T_d$  symmetry where two and three dimensional degenerate irreducible representations (irreps) are involved, our primitive vibrational basis will be built as a tensor product of non-degenerate, two- and three-fold degenerate harmonic oscillator functions

$$\phi_{v,\sigma_v}^{(C_v)} = \left( \phi_{v_1}^{(l_1, A_1)} \otimes \phi_{v_2}^{(l_2, m_2, C_2)} \otimes \phi_{v_3}^{(l_3, m_3, n_3, C_3)} \otimes \phi_{v_4}^{(l_4, m_4, n_4, C_4)} \right)_{\sigma_v}^{(C_v)}, \quad (13)$$

where the indices  $l_i, m_i, n_i$  and their permutations completely characterize the symmetrized powers of the creation operators with  $v_i = l_i + m_i + n_i$  for each mode. Using the symmetry-adapted basis functions (13) (SABF) allows (i) to block-diagonalize the Hamiltonian matrix and (ii) to classify and label the eigenstates of the Hamiltonian according to the different irreps of the symmetry group.

For  $J > 0$  calculations, the direct way is to construct the Hamiltonian matrix in a primitive basis obtained as the coupling between vibrational and rotational functions

$$\Phi_{v,M,\sigma}^{(J,C)} \equiv |(v_1 v_2 v_3 v_4) C_v, J n C_r; C \sigma, M\rangle = \left( \Phi_v^{(C_v)} \otimes \Phi_M^{(J, n C_r)} \right)_{\sigma}^{(C)} \quad (14)$$

SABF are generally built as a linear combination of the primitive functions (12). Different methods can be used to construct explicitly such basis functions. For instance, Yurchenko *et al.*<sup>91</sup> proposed a numerical method implemented in TROVE for the construction of SABF by solving eigenfunctions problems for a set of reduced (not in the sense defined in Section 4.1) Hamiltonian operators. For this work, the symmetrization coefficients are computed algebraically from irreducible tensor operators. Toward this end, each vibrational SABF is obtained by applying successive creation tensor operators on the vacuum state while the rotational SABF (also called non standard basis) defined in  $O(3) \supset T_d$  are obtained by applying a similarity transformation<sup>92</sup>  $G$  on the standard basis  $|J, K, M\rangle$ ;  $K$  and  $M$  being respectively the projection onto the molecular- and lab-fixed axis. The symmetry-adapted basis set is built as follows:

$$\begin{aligned} \Phi_{v,\sigma_v}^{(C_v)} &\equiv |(v_1 v_2 v_3 v_4) C_v \sigma_v\rangle = \left( \mathcal{A}^{+(l_1, A_1)} \otimes \dots \otimes \mathcal{A}^{+(l_4, m_4, n_4, C_4)} \right)_{\sigma_v}^{(C_v)} |0\rangle, \\ \Phi_{M,\sigma_r}^{(J, n C_r)} &\equiv |J, n C_r \sigma_r, M\rangle = \sum_K^{(J)} G_{n C_r}^K |J, K, M\rangle. \end{aligned} \quad (15)$$

If the molecule is in a field-free environment, then the  $M$  quantum number can be omitted in the following. As an illustration, we give below some normalized SABF expressed as a linear combination of non-symmetrized functions

$$\begin{aligned} |(0012) F_2 x\rangle &= \frac{1}{\sqrt{3}} [|000100200\rangle + |000100020\rangle + |000100002\rangle], \\ |(2100) E, 5E; A_1\rangle &= \frac{i}{2} [|210000000\rangle (|5, 4\rangle - |5, -4\rangle)] \\ &\quad + \frac{i}{2} [|201000000\rangle (|5, -2\rangle - |5, 2\rangle)] \\ |(0110) F_1, 4A_1; F_1 x\rangle &= |010100000\rangle \left( \frac{\sqrt{5}}{4\sqrt{2}} |4, -4\rangle \right. \\ &\quad \left. + \frac{\sqrt{7}}{4} |4, 0\rangle + \frac{\sqrt{5}}{4\sqrt{2}} |4, 4\rangle \right) \\ &\quad + |001100000\rangle \left( \frac{\sqrt{5}}{4\sqrt{6}} |4, -4\rangle \right. \\ &\quad \left. + \frac{\sqrt{7}}{4\sqrt{3}} |4, 0\rangle + \frac{\sqrt{5}}{4\sqrt{6}} |4, 4\rangle \right) \end{aligned}$$

**Table 3** Vibrational band centers (in  $\text{cm}^{-1}$ ) for  $^{12}\text{CF}_4$  up to  $3000\text{ cm}^{-1}$ . Vibrational sublevels are specified and "Eig" corresponds to  $\max(\mathbf{c}^T)$  in eqn (A2) that is the maximum contribution in the wavenumber decomposition

| Band, sym., eig.                    | Band center | Band, sym., eig.                     | Band center | Band, sym., eig.                     | Band center |
|-------------------------------------|-------------|--------------------------------------|-------------|--------------------------------------|-------------|
| $\nu_2, E(0.99)$                    | 435.41      | $\nu_2 + \nu_3 + \nu_4, F_2(0.78)$   | 2345.64     | $2\nu_2 + \nu_3 + \nu_4, F_1(0.87)$  | 2779.48     |
| $\nu_4, F_2(0.98)$                  | 631.09      | $\nu_2 + \nu_3 + \nu_4, F_2(-0.77)$  | 2346.00     | $2\nu_2 + \nu_3 + \nu_4, F_2(0.83)$  | 2779.48     |
| $2\nu_2, A_1(0.97)$                 | 868.01      | $\nu_2 + \nu_3 + \nu_4, F_1(0.89)$   | 2346.12     | $2\nu_2 + \nu_3 + \nu_4, F_1(-0.88)$ | 2779.71     |
| $2\nu_2, E(0.98)$                   | 871.21      | $\nu_2 + \nu_3 + \nu_4, E(-0.89)$    | 2349.41     | $2\nu_2 + \nu_3 + \nu_4, F_2(-0.84)$ | 2779.73     |
| $\nu_1, A_1(-0.94)$                 | 909.03      | $4\nu_2 + \nu_4, F_2(0.89)$          | 2362.15     | $2\nu_2 + \nu_3 + \nu_4, A_1(-0.85)$ | 2780.06     |
| $\nu_2 + \nu_4, F_2(-0.98)$         | 1066.20     | $4\nu_2 + \nu_4, F_1(0.94)$          | 2365.52     | $2\nu_2 + \nu_3 + \nu_4, E(0.87)$    | 2783.21     |
| $\nu_2 + \nu_4, F_1(-0.98)$         | 1066.76     | $4\nu_2 + \nu_4, F_2(-0.90)$         | 2365.77     | $5\nu_2 + \nu_4, F_2(0.90)$          | 2792.38     |
| $2\nu_4, F_2(-0.96)$                | 1260.59     | $4\nu_2 + \nu_4, F_1(-0.97)$         | 2374.93     | $5\nu_2 + \nu_4, F_1(0.92)$          | 2794.00     |
| $2\nu_4, A_1(-0.98)$                | 1261.86     | $4\nu_2 + \nu_4, F_2(-0.97)$         | 2374.94     | $\nu_1 + 3\nu_4, A_1(0.84)$          | 2795.69     |
| $2\nu_4, E(-0.98)$                  | 1262.19     | $\nu_1 + 2\nu_2 + \nu_4, F_2(-0.87)$ | 2407.29     | $\nu_1 + 3\nu_4, F_2(0.83)$          | 2796.41     |
| $\nu_3, F_2(-0.94)$                 | 1283.77     | $\nu_1 + 2\nu_2 + \nu_4, F_1(-0.91)$ | 2410.78     | $\nu_1 + 3\nu_4, F_1(0.84)$          | 2796.65     |
| $3\nu_2, E(-0.96)$                  | 1301.09     | $\nu_1 + 2\nu_2 + \nu_4, F_2(0.90)$  | 2410.87     | $5\nu_2 + \nu_4, F_2(0.92)$          | 2799.46     |
| $3\nu_2, A_1(0.98)$                 | 1307.37     | $2\nu_1 + \nu_4, F_2(0.86)$          | 2445.46     | $5\nu_2 + \nu_4, F_1(0.95)$          | 2799.58     |
| $3\nu_2, A_2(0.98)$                 | 1307.40     | $4\nu_4, A_1(-0.83)$                 | 2516.33     | $\nu_1 + 3\nu_4, F_2(0.90)$          | 2799.88     |
| $\nu_1 + \nu_2, E(0.93)$            | 1344.56     | $4\nu_4, F_2(-0.86)$                 | 2516.40     | $5\nu_2 + \nu_4, F_2(-0.99)$         | 2811.81     |
| $2\nu_2 + \nu_4, F_2(0.95)$         | 1498.94     | $4\nu_4, E(0.85)$                    | 2517.09     | $5\nu_2 + \nu_4, F_1(-0.99)$         | 2811.81     |
| $2\nu_2 + \nu_4, F_1(-0.98)$        | 1502.25     | $4\nu_4, F_2(-0.87)$                 | 2518.34     | $\nu_1 + \nu_3 + \nu_4, E(-0.87)$    | 2812.24     |
| $2\nu_2 + \nu_4, F_2(-0.96)$        | 1502.34     | $4\nu_4, F_1(-0.88)$                 | 2518.85     | $\nu_1 + \nu_3 + \nu_4, F_2(-0.81)$  | 2814.92     |
| $\nu_1 + \nu_4, F_2(0.93)$          | 1539.46     | $4\nu_4, E(-0.96)$                   | 2523.73     | $\nu_1 + \nu_3 + \nu_4, F_1(-0.83)$  | 2815.26     |
| $\nu_2 + 2\nu_4, F_1(0.95)$         | 1695.62     | $4\nu_4, A_1(-0.96)$                 | 2523.75     | $\nu_1 + \nu_3 + \nu_4, A_1(0.81)$   | 2818.68     |
| $\nu_2 + 2\nu_4, F_2(-0.95)$        | 1696.05     | $2\nu_3, A_1(0.68)$                  | 2534.60     | $\nu_1 + 3\nu_2 + \nu_4, F_2(0.84)$  | 2839.64     |
| $\nu_2 + 2\nu_4, E(-0.78)$          | 1696.81     | $\nu_3 + 2\nu_4, F_2(0.60)$          | 2536.17     | $\nu_1 + 3\nu_2 + \nu_4, F_1(0.85)$  | 2840.80     |
| $\nu_2 + 2\nu_4, A_1(-0.97)$        | 1696.98     | $\nu_3 + 2\nu_4, F_1(-0.96)$         | 2536.51     | $\nu_1 + 3\nu_2 + \nu_4, F_2(0.89)$  | 2846.89     |
| $\nu_2 + 2\nu_4, E(0.78)$           | 1697.97     | $\nu_3 + 2\nu_4, E(0.82)$            | 2540.86     | $\nu_1 + 3\nu_2 + \nu_4, F_1(0.90)$  | 2846.93     |
| $\nu_2 + 2\nu_4, A_2(0.98)$         | 1698.12     | $\nu_3 + 2\nu_4, F_2(0.74)$          | 2541.09     | $2\nu_1 + \nu_2 + \nu_4, F_2(0.84)$  | 2880.83     |
| $\nu_2 + \nu_3, F_2(-0.93)$         | 1716.78     | $\nu_3 + 2\nu_4, F_1(0.86)$          | 2541.43     | $2\nu_1 + \nu_2 + \nu_4, F_1(-0.84)$ | 2881.47     |
| $\nu_2 + \nu_3, F_1(0.94)$          | 1717.19     | $\nu_3 + 2\nu_4, F_2(0.57)$          | 2544.07     | $\nu_2 + 4\nu_4, E(-0.77)$           | 2950.82     |
| $4\nu_2, A_1(0.93)$                 | 1731.52     | $\nu_3 + 2\nu_4, A_1(0.66)$          | 2547.91     | $\nu_2 + 4\nu_4, F_2(0.84)$          | 2950.98     |
| $4\nu_2, E(0.95)$                   | 1734.62     | $3\nu_2 + 2\nu_4, F_1(0.91)$         | 2560.58     | $\nu_2 + 4\nu_4, F_1(0.83)$          | 2951.37     |
| $4\nu_2, E(-0.98)$                  | 1743.93     | $3\nu_2 + 2\nu_4, F_2(0.90)$         | 2561.34     | $\nu_2 + 4\nu_4, A_2(0.84)$          | 2951.52     |
| $\nu_1 + 2\nu_2, A_1(0.89)$         | 1777.04     | $3\nu_2 + 2\nu_4, E(-0.69)$          | 2561.61     | $\nu_2 + 4\nu_4, E(0.80)$            | 2951.86     |
| $\nu_1 + 2\nu_2, E(-0.92)$          | 1780.40     | $3\nu_2 + 2\nu_4, A_1(0.92)$         | 2561.73     | $\nu_2 + 4\nu_4, A_1(-0.82)$         | 2952.13     |
| $2\nu_1, A_1(0.88)$                 | 1815.74     | $2\nu_3, F_2(0.87)$                  | 2562.17     | $\nu_2 + 4\nu_4, F_1(0.79)$          | 2952.67     |
| $3\nu_4, A_1(-0.92)$                | 1888.80     | $3\nu_2 + 2\nu_4, E(0.71)$           | 2563.79     | $\nu_2 + 4\nu_4, F_2(0.66)$          | 2952.90     |
| $3\nu_4, F_2(0.92)$                 | 1889.63     | $3\nu_2 + 2\nu_4, A_2(0.94)$         | 2563.96     | $\nu_2 + 4\nu_4, F_2(-0.66)$         | 2953.91     |
| $3\nu_4, F_1(-0.92)$                | 1889.89     | $3\nu_2 + 2\nu_4, F_2(0.92)$         | 2567.32     | $\nu_2 + 4\nu_4, F_1(0.80)$          | 2954.21     |
| $3\nu_4, F_2(0.97)$                 | 1892.98     | $3\nu_2 + 2\nu_4, F_1(0.93)$         | 2567.35     | $\nu_2 + 4\nu_4, A_1(0.94)$          | 2957.84     |
| $\nu_3 + \nu_4, E(-0.96)$           | 1909.63     | $3\nu_2 + 2\nu_4, A_1(0.96)$         | 2569.17     | $\nu_2 + 4\nu_4, A_2(0.66)$          | 2957.85     |
| $\nu_3 + \nu_4, F_2(-0.90)$         | 1912.32     | $3\nu_2 + 2\nu_4, A_2(0.95)$         | 2569.25     | $\nu_2 + 4\nu_4, A_2(-0.97)$         | 2960.14     |
| $\nu_3 + \nu_4, F_1(0.91)$          | 1912.59     | $3\nu_2 + 2\nu_4, E(0.95)$           | 2569.48     | $\nu_2 + 4\nu_4, E(0.68)$            | 2960.15     |
| $\nu_3 + \nu_4, A_1(-0.90)$         | 1916.00     | $3\nu_2 + 2\nu_4, E(-0.95)$          | 2569.55     | $\nu_2 + 2\nu_3, E(0.72)$            | 2966.82     |
| $3\nu_2 + \nu_4, F_2(0.94)$         | 1931.48     | $2\nu_3, E(-0.93)$                   | 2570.22     | $\nu_2 + \nu_3 + 2\nu_4, F_2(0.60)$  | 2968.64     |
| $3\nu_2 + \nu_4, F_1(-0.96)$        | 1932.58     | $3\nu_2 + \nu_3, F_2(-0.88)$         | 2577.97     | $\nu_2 + \nu_3 + 2\nu_4, F_1(0.75)$  | 2969.32     |
| $3\nu_2 + \nu_4, F_2(0.97)$         | 1938.41     | $3\nu_2 + \nu_3, F_1(-0.90)$         | 2578.61     | $\nu_2 + \nu_3 + 2\nu_4, F_2(-0.93)$ | 2969.61     |
| $3\nu_2 + \nu_4, F_1(0.98)$         | 1938.45     | $3\nu_2 + \nu_3, F_2(0.91)$          | 2584.64     | $\nu_2 + \nu_3 + 2\nu_4, F_1(0.59)$  | 2969.95     |
| $\nu_1 + \nu_2 + \nu_4, F_2(-0.92)$ | 1974.67     | $3\nu_2 + \nu_3, F_1(0.92)$          | 2584.67     | $\nu_2 + \nu_3 + 2\nu_4, A_2(0.81)$  | 2974.09     |
| $\nu_1 + \nu_2 + \nu_4, F_1(0.92)$  | 1975.27     | $6\nu_2, A_1(0.89)$                  | 2590.86     | $\nu_2 + \nu_3 + 2\nu_4, F_1(-0.65)$ | 2974.10     |
| $2\nu_2 + 2\nu_4, F_2(-0.92)$       | 2128.14     | $6\nu_2, E(0.90)$                    | 2593.90     | $\nu_2 + \nu_3 + 2\nu_4, F_2(-0.63)$ | 2974.20     |
| $2\nu_2 + 2\nu_4, A_1(0.92)$        | 2129.41     | $6\nu_2, E(-0.93)$                   | 2603.00     | $\nu_2 + \nu_3 + 2\nu_4, E(-0.80)$   | 2974.29     |
| $2\nu_2 + 2\nu_4, E(0.91)$          | 2129.67     | $\nu_1 + \nu_2 + 2\nu_4, F_1(-0.87)$ | 2603.31     | $\nu_2 + \nu_3 + 2\nu_4, A_1(0.79)$  | 2974.58     |
| $2\nu_2 + 2\nu_4, F_1(-0.94)$       | 2131.42     | $\nu_1 + \nu_2 + 2\nu_4, F_2(-0.87)$ | 2603.74     | $\nu_2 + \nu_3 + 2\nu_4, F_2(-0.56)$ | 2975.21     |
| $2\nu_2 + 2\nu_4, F_2(-0.93)$       | 2131.46     | $\nu_1 + \nu_2 + 2\nu_4, E(-0.70)$   | 2604.56     | $\nu_2 + \nu_3 + 2\nu_4, F_1(0.76)$  | 2975.39     |
| $2\nu_2 + 2\nu_4, E(0.87)$          | 2133.07     | $\nu_1 + \nu_2 + 2\nu_4, A_1(0.90)$  | 2604.70     | $\nu_2 + \nu_3 + 2\nu_4, F_2(-0.56)$ | 2977.03     |
| $2\nu_2 + 2\nu_4, A_2(-0.97)$       | 2133.29     | $\nu_1 + \nu_2 + 2\nu_4, E(-0.71)$   | 2605.79     | $\nu_2 + \nu_3 + 2\nu_4, F_1(-0.54)$ | 2977.61     |
| $2\nu_2 + 2\nu_4, E(-0.85)$         | 2133.54     | $\nu_1 + \nu_2 + 2\nu_4, A_2(0.91)$  | 2605.93     | $\nu_2 + \nu_3 + 2\nu_4, E(0.70)$    | 2980.25     |
| $2\nu_2 + 2\nu_4, A_1(-0.93)$       | 2133.59     | $6\nu_2, A_2(0.99)$                  | 2618.15     | $4\nu_2 + 2\nu_4, F_2(0.87)$         | 2990.92     |
| $2\nu_2 + \nu_3, F_2(0.91)$         | 2147.34     | $6\nu_2, A_1(-0.99)$                 | 2618.15     | $4\nu_2 + 2\nu_4, A_1(0.84)$         | 2992.09     |
| $2\nu_2 + \nu_3, F_1(-0.93)$        | 2150.60     | $\nu_1 + \nu_2 + \nu_3, F_2(0.85)$   | 2620.18     | $4\nu_2 + 2\nu_4, E(-0.82)$          | 2992.29     |
| $2\nu_2 + \nu_3, F_2(-0.92)$        | 2150.65     | $\nu_1 + \nu_2 + \nu_3, F_1(-0.85)$  | 2620.59     | $\nu_2 + 2\nu_3, F_1(-0.85)$         | 2993.11     |
| $5\nu_2, E(0.91)$                   | 2162.46     | $\nu_1 + 4\nu_2, A_1(0.81)$          | 2640.10     | $\nu_2 + 2\nu_3, F_2(-0.86)$         | 2993.38     |
| $\nu_1 + 2\nu_4, F_2(0.89)$         | 2168.21     | $\nu_1 + 4\nu_2, E(0.83)$            | 2643.33     | $4\nu_2 + 2\nu_4, F_1(-0.89)$        | 2994.16     |
| $5\nu_2, A_1(-0.92)$                | 2168.54     | $\nu_1 + 4\nu_2, E(0.91)$            | 2653.00     | $4\nu_2 + 2\nu_4, F_2(-0.88)$        | 2994.24     |
| $5\nu_2, A_2(0.95)$                 | 2168.65     | $2\nu_1 + 2\nu_2, A_1(0.80)$         | 2683.82     | $4\nu_2 + 2\nu_4, E(-0.75)$          | 2996.13     |
| $\nu_1 + 2\nu_4, A_1(-0.91)$        | 2169.55     | $2\nu_1 + 2\nu_2, E(0.85)$           | 2687.37     | $4\nu_2 + 2\nu_4, A_2(0.95)$         | 2996.30     |
| $\nu_1 + 2\nu_4, E(0.92)$           | 2169.84     | $3\nu_1, A_1(-0.80)$                 | 2720.18     | $4\nu_2 + 2\nu_4, E(-0.69)$          | 2997.09     |
| $5\nu_2, E(-0.97)$                  | 2180.86     | $2\nu_2 + 3\nu_4, A_1(-0.88)$        | 2755.97     | $4\nu_2 + 2\nu_4, A_1(0.85)$         | 2997.10     |
| $\nu_1 + \nu_3, F_2(0.87)$          | 2186.97     | $2\nu_2 + 3\nu_4, F_2(-0.86)$        | 2756.65     | $\nu_2 + 2\nu_3, A_1(0.92)$          | 3000.49     |
| $\nu_1 + 3\nu_2, E(0.86)$           | 2209.97     | $2\nu_2 + 3\nu_4, F_1(0.85)$         | 2756.88     | $\nu_2 + 2\nu_3, E(0.92)$            | 3001.16     |
| $\nu_1 + 3\nu_2, A_1(0.91)$         | 2216.52     | $2\nu_2 + 3\nu_4, E(0.89)$           | 2759.20     | $\nu_2 + 2\nu_3, A_2(-0.93)$         | 3001.69     |

Table 3 (continued)

| Band, sym., eig.                    | Band center | Band, sym., eig.                    | Band center | Band, sym., eig.              | Band center |
|-------------------------------------|-------------|-------------------------------------|-------------|-------------------------------|-------------|
| $\nu_1 + 3\nu_2, A_2(0.92)$         | 2216.56     | $2\nu_2 + 3\nu_4, F_2(-0.69)$       | 2759.91     | $4\nu_2 + 2\nu_4, F_2(0.92)$  | 3003.53     |
| $2\nu_1 + \nu_2, E(-0.86)$          | 2251.46     | $2\nu_2 + 3\nu_4, F_2(0.71)$        | 2760.10     | $4\nu_2 + 2\nu_4, F_1(-0.92)$ | 3003.54     |
| $\nu_2 + 3\nu_4, E(0.91)$           | 2323.84     | $2\nu_2 + 3\nu_4, F_1(0.85)$        | 2760.18     | $4\nu_2 + 2\nu_4, E(0.97)$    | 3005.65     |
| $\nu_2 + 3\nu_4, F_1(-0.77)$        | 2324.39     | $2\nu_2 + 3\nu_4, F_1(-0.87)$       | 2760.30     | $4\nu_2 + 2\nu_4, A_2(0.97)$  | 3005.95     |
| $\nu_2 + 3\nu_4, F_2(-0.69)$        | 2324.43     | $2\nu_2 + 3\nu_4, F_2(0.84)$        | 2760.43     | $4\nu_2 + 2\nu_4, E(0.97)$    | 3005.96     |
| $\nu_2 + 3\nu_4, F_2(0.69)$         | 2325.18     | $2\nu_2 + 3\nu_4, F_1(-0.96)$       | 2763.99     | $4\nu_2 + 2\nu_4, A_1(0.98)$  | 3005.97     |
| $\nu_2 + 3\nu_4, F_1(0.78)$         | 2325.22     | $2\nu_2 + 3\nu_4, F_2(-0.88)$       | 2764.68     | $4\nu_2 + \nu_3, F_2(0.85)$   | 3006.53     |
| $\nu_2 + 3\nu_4, F_2(-0.95)$        | 2327.45     | $2\nu_2 + \nu_3 + \nu_4, E(0.92)$   | 2772.75     | $4\nu_2 + \nu_3, F_1(0.87)$   | 3009.72     |
| $\nu_2 + 3\nu_4, F_1(-0.97)$        | 2329.15     | $2\nu_2 + \nu_3 + \nu_4, A_1(0.93)$ | 2775.93     | $4\nu_2 + \nu_3, F_2(-0.86)$  | 3009.80     |
| $\nu_2 + \nu_3 + \nu_4, A_1(-0.96)$ | 2342.20     | $2\nu_2 + \nu_3 + \nu_4, A_2(0.96)$ | 2776.09     | $4\nu_2 + \nu_3, F_1(-0.90)$  | 3019.11     |
| $\nu_2 + \nu_3 + \nu_4, E(-0.96)$   | 2342.65     | $2\nu_2 + \nu_3 + \nu_4, F_2(0.85)$ | 2776.09     | $4\nu_2 + \nu_3, F_2(0.90)$   | 3019.11     |
| $\nu_2 + \nu_3 + \nu_4, A_2(-0.96)$ | 2343.31     | $2\nu_2 + \nu_3 + \nu_4, F_1(0.86)$ | 2776.35     |                               |             |
| $\nu_2 + \nu_3 + \nu_4, F_1(0.88)$  | 2345.41     | $2\nu_2 + \nu_3 + \nu_4, E(0.94)$   | 2776.43     |                               |             |

A comparison of the symmetrized Cartesian and angular momentum basis sets in spherical tops is given in ref. 93.

### 4.3 Basis set compression

A full account of all symmetry properties is a necessary but unfortunately not a sufficient condition to drastically reduce the size of the basis and to make the rovibrational calculations more tractable as  $V$ ,  $J$  and/or  $N$  increase. Typically, block-diagonalization of the Hamiltonian matrix for  $XY_4$  reduces the size of the problem by one order of magnitude while it should be reduced ideally by five orders of magnitude to facilitate calculations. Many methods have been developed these past few years for reducing the size of the basis set and computing eigensolutions of a (ro)-vibrational problem.<sup>91,94–100</sup> Very recently, in a rather pedagogical paper<sup>101</sup> Carrington reviewed the main methods for computing spectra of molecules with more than four atoms using pruned-and-contracted-basis sets and direct/iterative eigensolvers.

In this work, we applied the strategy combining the basis set “pruning” and reduction–compression scheme ref. 77, 78 and 88. This approach, which permits the use of a realistic reduced anharmonic basis set for ro-vibrational calculations, has proven its efficiency for methane calculations for a large

ranges of wavenumbers and temperatures. Technical details are given in Appendix A.

To summarize, the number of vibrational functions in a  $XY_4$  molecule can be drastically reduced using the following scheme

$$\text{Eqn (12)} \rightarrow \text{Eqn (13), (A1)} \rightarrow \text{Eqn (A3)}$$

$$D_0 \sim 10^9 \xrightarrow[\text{Pruning}]{\text{Symmetry}} D_1 \sim 10^5 \xrightarrow{\text{Reduction}} D_2 \sim 10^3,$$

where  $D_i$  is the size of the Hamiltonian matrix:  $D_0$  is defined below eqn (12),  $D_1$  is defined below eqn (A1) for each symmetry block and  $D_2$  is defined above eqn (A3). Without such a compression of the basis set, full quantum calculation of five atomic systems with high  $J$  would not be feasible, even using the latest generation of computers.

170 band centers predicted from the refined PES together with their vibrational assignments, symmetry types and major basis set contributions are given in Table 3.

## 5 Line list calculation: completeness issues

Full quantum-mechanical calculations of  $CF_4$  spectra are an extremely challenging problem because highly excited basis

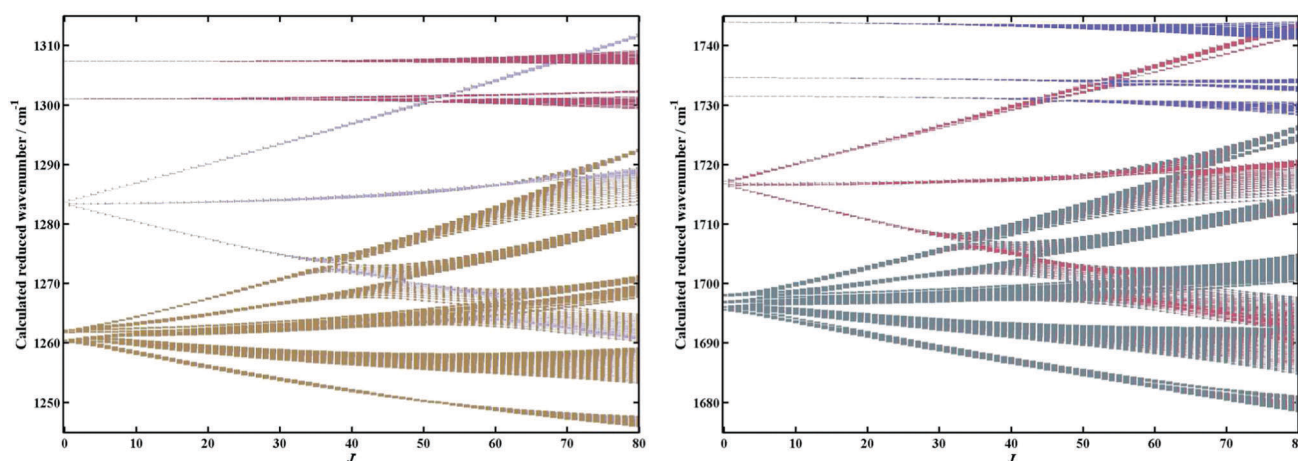


Fig. 2 Reduced energy level  $E_{vJ} - E_r^{(0)}$  for  $CF_4$  for the  $2\nu_4(F_2, A_1, E)$ ,  $\nu_3(F_2)$ ,  $3\nu_2(E, A_1, A_2)$  (left panel) and  $\nu_2 + 2\nu_4(F_1, F_2, E, A_1, E)$ ,  $\nu_2 + \nu_3(F_2)$ ,  $4\nu_2(A_1, 2E)$  (right panel).  $E_r^{(0)}$  stands for the semirigid rotor energies (see text). For each panel, the subbands are listed by increasing energy order.

states in a wide spectral range have to be involved. This requires efficient computational methods for (1) computing converged line positions – diagonalization of very large matrices being the most demanding part – as well as an efficient procedure for (2) taking into account all necessary line transitions with appropriate intensity cut-offs for converged opacities. All calculations were performed using our home-made computer code ‘TENSOR’.

### 5.1 Vibration–rotation energy levels and partition function

All details for the basis compression are given in Appendix B. To illustrate how dense the CF<sub>4</sub> energy patterns are with complex level inter-mode coupling, we plot in Fig. 2 and 3 some reduced variationally-predicted energy levels as a function of the  $J$  rotational quantum number. The term reduced means that we subtract the semirigid rotor energies  $E_r^{(0)} = 0.1902J(J+1) - 6.13 \times 10^{-8}J^2(J+1)^2$  from the total energy  $E_{vj}$ . In Fig. 2 the sublevels  $2\nu_4(A_1, E, F_2)$ ,  $\nu_3(F_2)$ ,  $3\nu_2(A_1, A_2, E)$  are displayed on the left panel and  $\nu_2 + 2\nu_4(A_1, 2E, F_1F_2)$ ,  $\nu_2 + \nu_3(F_2)$ ,  $4\nu_2(A_1, A_2, E)$  on the right panel that is “right” = “left” +  $\nu_2$ . The  $\nu_4 = 2$  levels are split to  $l_4 = 0$  and  $l_4 = 2$  in the  $O(3)$  rotation group and the  $l_4 = 2$  sublevel further splits into  $E$  and  $F_2$  according to the

$T_d$  symmetry types. The tetrahedral splitting is thus clearly seen as  $J$  increases.

Different colours are also displayed on these 3 figures, they give information on how the energy states are mixed altogether. Each eigenvector is decomposed as  $\Psi_j = c_1^j \Phi_1 + c_2^j \Phi_2 + \dots$  in the primitive basis (14), then different colours will represent a mixing coefficient  $(c_i^j)^2$ . Information in these figures is twofold. First one clearly sees the numerous overlapping vibration–rotation patterns as  $E$  and  $J$  increase making an appropriate choice of effective empirical model hazardous for the corresponding complicated band system. This point will be discussed in more detail in Section 7.1. Second, the density of state rapidly increases with  $J$  resulting in strongly congested CF<sub>4</sub> spectra even for the region below  $3000\text{ cm}^{-1}$ .

Another quantity that can be computed from energy levels is the total internal partition function expressed as a direct sum of Boltzmann factors at a given temperature

$$Q_{\text{dir}}(T) = \sum (2J+1)g_C \exp(-hcE_{vj}/kT). \quad (16)$$

where  $g_C$  are the nuclear spin statistical weights. For CF<sub>4</sub> they are given by  $g_A = 5$ ,  $g_E = 2$  and  $g_F = 3$ . At  $T = 296\text{ K}$ , we obtain

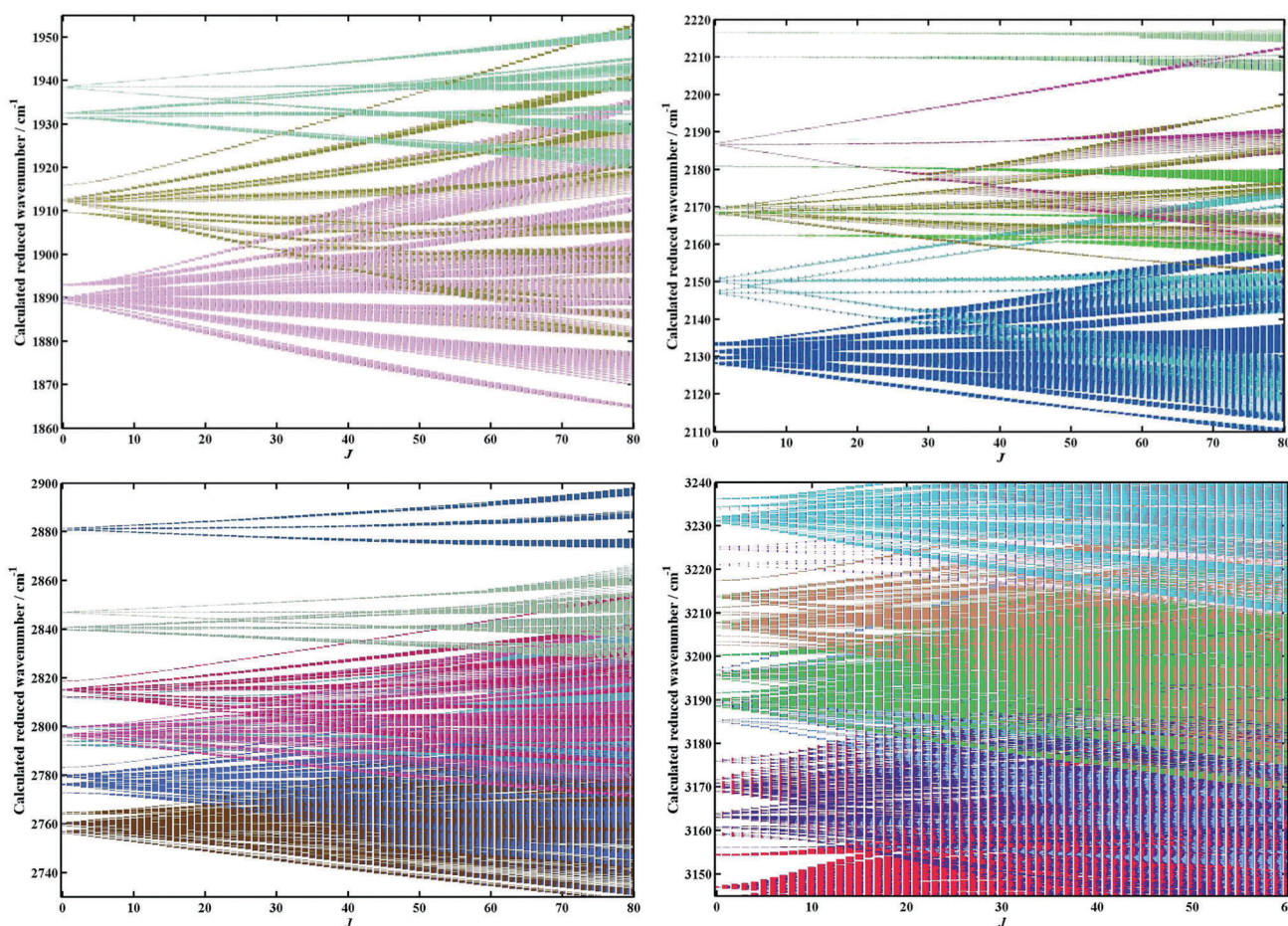


Fig. 3 Reduced energy level  $E_{vj} - E_r^{(0)}$  for CF<sub>4</sub> in the region of the  $3\nu_4$ ,  $\nu_3 + \nu_4$ ,  $3\nu_2 + \nu_4$  (top, left-hand side),  $2\nu_2 + 2\nu_4$ ,  $2\nu_2 + \nu_3$ ,  $5\nu_2$ ,  $\nu_1 + 2\nu_4$ ,  $\nu_1 + \nu_3$ ,  $\nu_1 + 3\nu_2$  (top, right-hand side),  $2\nu_2 + 3\nu_4$ ,  $2\nu_2 + \nu_3 + \nu_4$ ,  $5\nu_2 + \nu_4$ ,  $\nu_1 + 3\nu_4$ ,  $\nu_1 + \nu_3 + \nu_4$ ,  $\nu_1 + 3\nu_2 + \nu_4$ ,  $2\nu_1 + \nu_2 + \nu_4$  (bottom, left-hand side) and  $5\nu_4$ ,  $3\nu_4 + \nu_2$ ,  $2\nu_3 + \nu_4$ ,  $\nu_3 + 3\nu_4$ ,  $3\nu_2 + 3\nu_4$ , etc. (bottom, right-hand side). For each panel, the bands are listed by increasing energy order (see Table 3 for details about the vibrational sublevels).

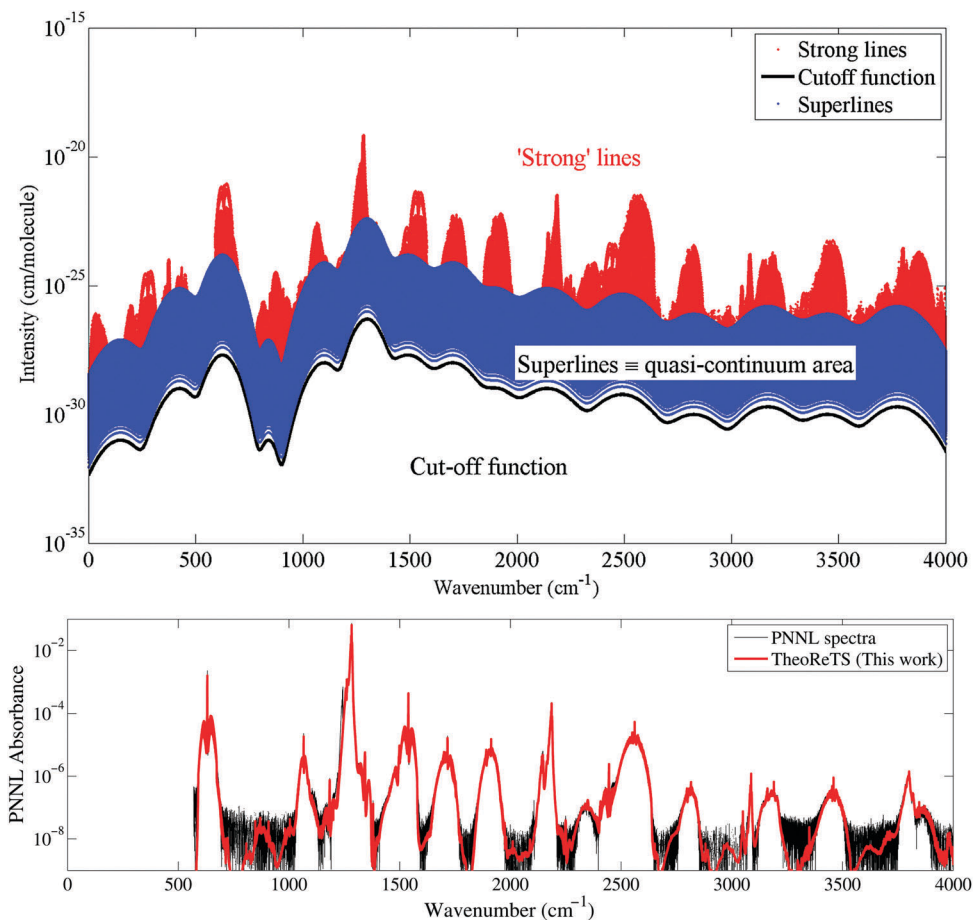


Fig. 4 Top panel: Strong line intensities (in red) and schematic representation of the quasi-continuum area (QC, in blue, treated using super-lines) of tetrafluoromethane at 296 K. A plot of the cut-off function (see Section 5.2) is also given (black line). Bottom panel: Plot of the PNNL absorbance and comparison with TheoReTS (this work).

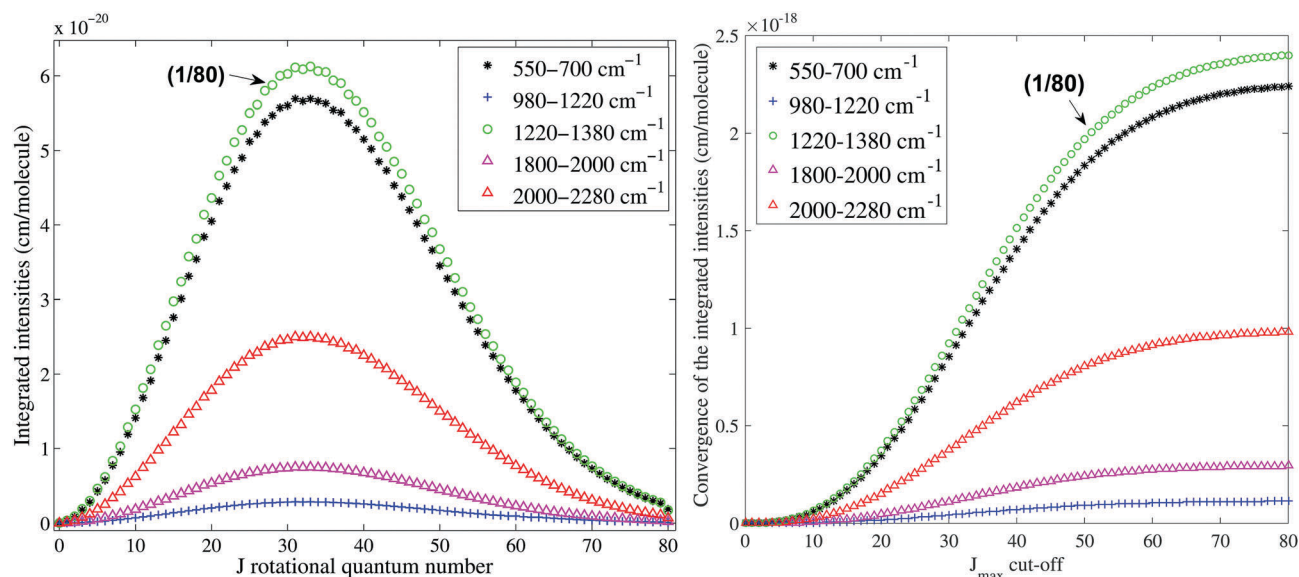


Fig. 5 Integrated intensities (in cm per molecules) for different spectral ranges (see Fig. 4) at different values of the rotational angular momentum  $J$  (left panel). Convergence of the integrated intensities (in cm per molecules) with respect to a given  $J_{\max}$  value is also shown (right panel, see Discussion in Appendix B).

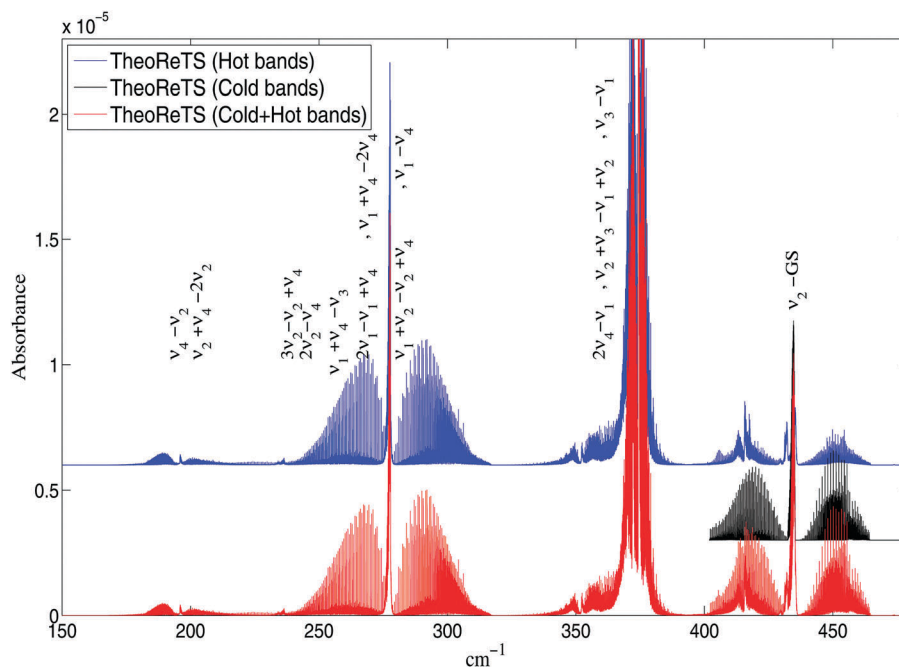


Fig. 6 Theoretical absorbance computed at 296 K with contributions of the cold and hot band transitions. The main CBs and HBs are indicated in vertical text.

$Q_{\text{dir}} = 126\,600$ . Note that at 180 K, the convergence of the partition function is better than 1% at  $J_{\text{max}} = 60$  and better than 0.1% at  $J_{\text{max}} = 71$ . At 250 K the convergence is better than 1% at  $J_{\text{max}} = 70$  and better than 0.1% at  $J_{\text{max}} = 78$  while at 296 K the convergence is better than 1% at  $J_{\text{max}} = 74$  and better than 0.1% at  $J_{\text{max}} = 80$ . For higher temperatures, the partition function is not totally converged at  $J_{\text{max}} = 80$ . Note that in the direct-product approximation  $Q_{\text{prod}}(T) \approx Q_{\text{vib}}Q_{\text{rot}}$  as usually used in effective models, the value given in HITRAN<sup>2</sup> is  $Q_{\text{prod}}(296) = 121\,270$ . The discrepancy between  $Q_{\text{prod}}$  and  $Q_{\text{dir}}$  is 4.2% at room temperature which is not negligible but this difference may be absorbed by the empirically-determined dipole moment parameters during the fit. Concerning the energy cut-off, the maximum  $E_{\text{max}}$  value used to compute the partition function is  $6700\text{ cm}^{-1}$ . The partition function is totally converged using this cut-off. Even if  $E_{\text{max}} = 5500\text{ cm}^{-1}$ , the convergence error of the partition function is less than 0.001%. Values of the partition function are provided in TheoReTS from 100 to 300 K.

## 5.2 Completeness criteria for intensities and quasi-continuum absorption

With the cut-offs  $E_{\text{low}}^{\text{max}}$  and  $I^{\text{cut}}(W)$  defined in Appendix B, the final room-temperature list is composed of more than 2 billion transitions. Such a huge number of lines is hardly manageable for practical applications using standard radiative transfer codes. We employed here the same strategy as for exhaustive high-temperature line lists, namely a separate treatment of a much smaller set of strong (sharp) absorption features and the remaining bulk of blended weak lines, resulting in quasi-continuum (QC) cross-sections.<sup>33–35</sup> The QC is not a true continuum *stricto sensu* but

rather an accumulation of very weak overlapping lines. Their individual contribution is quite negligible but the sum over a hundred million weak lines turns out to be absolutely essential for an accurate modelling of the observed opacity and of its temperature dependence.

The strong lines are selected by the condition  $I \geq \alpha I^{\text{cut}}(W)$  where  $\alpha$  is a scaling factor used for an optimal separation of the QC lines and defined to get the best ratio (number of lines)/(integrated intensity) in the file of strong lines. At 296 K we fixed  $\alpha = 8000$  for which 0.0237% of strong lines gives  $\sim 91.5\%$  of opacity in the full  $W$  range. The “strong lines” file contains 512 000 lines while the QC includes  $\sim 2.162$  billion lines. This huge amount of data was compressed using the super-line (SL) approach, as described in ref. 29 and 34. This is a quite simple tool to drastically reduce the number of lines by several orders of magnitude with almost no impact on the accuracy of the calculated absorbance. This becomes much more flexible for the user. For this work, all intensities were summed up by steps of  $0.001\text{ cm}^{-1}$  to form only one compressed SL. This final SL library now contains 3.89 million entries, the corresponding compression is thus scaled by three orders of magnitude.

## 6 Comparison with PNNL experimental spectra

Validation of our variationally computed  $\text{CF}_4$  line lists with experimental data is an important step in order to evaluate the accuracy of *ab initio* PES and DMS with respect to predicted CB and HB transitions. A possible source of validation is the comparison of theoretical absorbance with medium-resolution

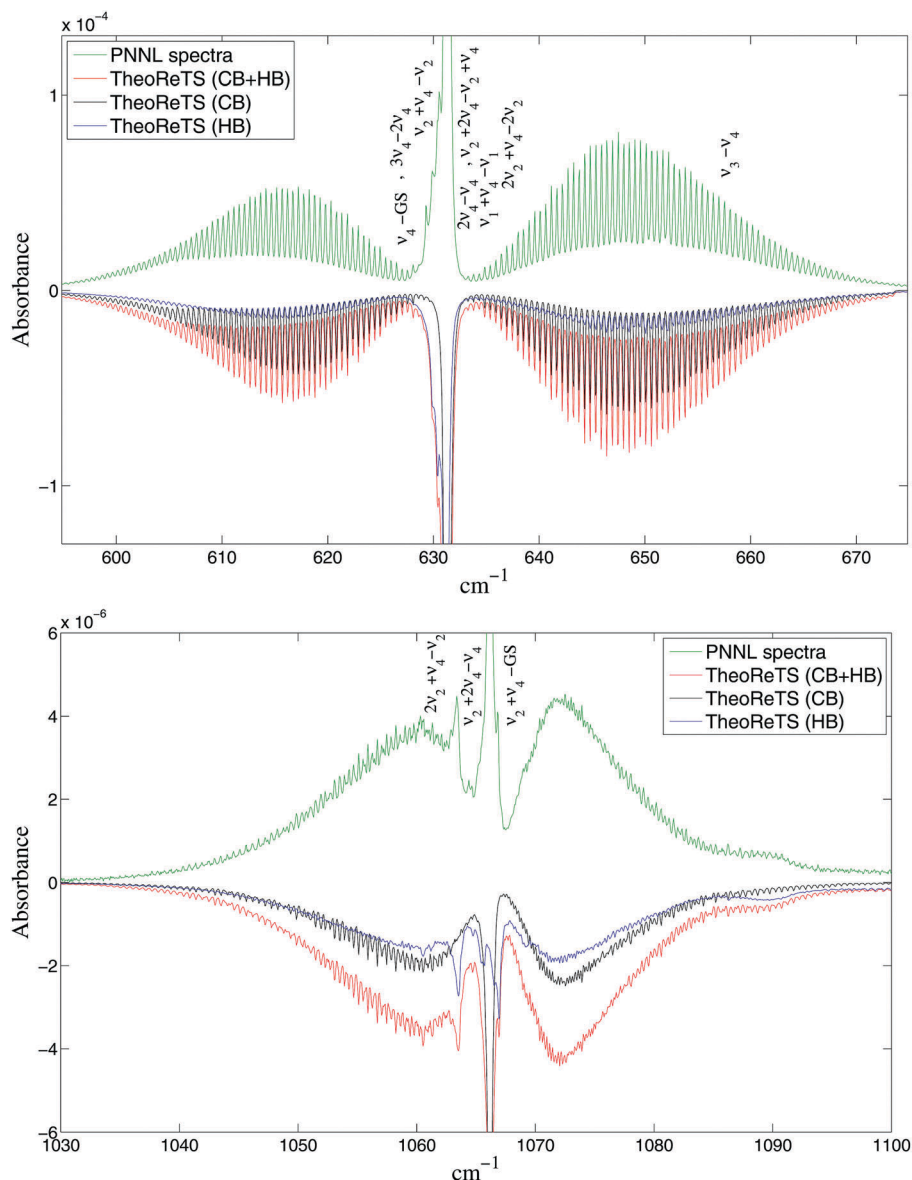


Fig. 7 Example comparison of theoretical (CB, HB and CB + HB transitions) absorption spectra of  $\text{CF}_4$  with experimental PNNL records at room temperature in the  $\nu_4$  region (top panel) and  $\nu_2 + \nu_4$  region (bottom panel) around 600–1100  $\text{cm}^{-1}$ . The main CBs and HBs are also indicated in vertical text (see also Sections 6.2 and 6.3).

experimental Fourier-transform laboratory PNNL<sup>103</sup> spectra.‡ They are recorded at temperatures of 5, 25, and 50 °C with a pressure of 1 atm over in a large spectral range of 570–6500  $\text{cm}^{-1}$ . Note that the PNNL spectra have not been assigned. Consequently, their line or band parameters have not been extracted, but these raw data provide a wealth of experimental information for validation of theoretical predictions. All comparisons of this section correspond to 296 K. The main absorption regions are considered below and in order to avoid overloading the figures, only the strongest absorption bands are specified on each one; there are many more in reality (see Section 7.1 for more details).

‡ If necessary, the PNNL absorbance spectra could be converted into cross sections (in  $\text{cm}^2$  per molecule) using the factor  $9.28697 \times 10^{-16}$ .

### 6.1 Region 150–500 $\text{cm}^{-1}$

Since PNNL spectra are not available in this region, we only give theoretical predictions of CBs and HBs, which are displayed in Fig. 6. In this region, our list contains 36 million lines essentially composed of HB transitions, except for the interval around 435  $\text{cm}^{-1}$  of the  $\nu_2$  fundamental band. The four strongest absorption intervals, in decreasing order, correspond to the HB and CB transitions belonging to  $\nu_3 - \nu_1$ ,  $\nu_1 - \nu_4$ ,  $\nu_2$  and  $\nu_2 + \nu_3 - \nu_1 - \nu_2$  with integrated intensities of  $1.48 \times 10^{-21}$ ,  $5.96 \times 10^{-22}$ ,  $2.28 \times 10^{-22}$  and  $1.44 \times 10^{-22}$   $\text{cm}^2$  per molecule, respectively, which is clearly dominated by HBs. Here and below we quote the integrated intensities of individual bands or of selective absorption features using the “strong lines” sample only. For QCs the contributions of individual band are not separable.

**Table 4** Comparison of the integrated PNNL absorbance of CF<sub>4</sub> with TheoReTS (this work) and TFMecaSDa<sup>102</sup> (<http://vamdc.icb.cnrs.fr/PHP/tfmecasda.php>)

| Region (cm <sup>-1</sup> ) | TheoReTS (TW)         | PNNL                  | Error (%) |
|----------------------------|-----------------------|-----------------------|-----------|
| 570–700                    | $3.96 \times 10^{-2}$ | $3.84 \times 10^{-2}$ | –3.0      |
| 980–1220                   | $2.00 \times 10^{-3}$ | $2.25 \times 10^{-3}$ | 12.0      |
| 1220–1380                  | 3.39                  | 3.42                  | 0.9       |
| 1380–1600                  | $2.30 \times 10^{-2}$ | $2.39 \times 10^{-2}$ | 3.8       |
| 1600–1800                  | $3.64 \times 10^{-3}$ | $3.73 \times 10^{-3}$ | 2.5       |
| 1800–2000                  | $5.18 \times 10^{-3}$ | $5.20 \times 10^{-3}$ | 0.4       |
| 2000–2280                  | $1.73 \times 10^{-2}$ | $1.75 \times 10^{-3}$ | 1.2       |
| 2280–2700                  | $2.05 \times 10^{-2}$ | $1.94 \times 10^{-2}$ | –5.5      |
| 2700–2950                  | $3.00 \times 10^{-4}$ | $2.99 \times 10^{-4}$ | –0.3      |
| 2950–3300                  | $3.77 \times 10^{-4}$ | $4.04 \times 10^{-4}$ | 6.7       |
| 3300–3950                  | $8.68 \times 10^{-4}$ | $9.07 \times 10^{-4}$ | 4.5       |

| Region (cm <sup>-1</sup> ) | TFMecaSDa <sup>102</sup> | PNNL                  | Error (%) |
|----------------------------|--------------------------|-----------------------|-----------|
| 980–1220                   | $1.42 \times 10^{-3}$    | $2.25 \times 10^{-3}$ | 74.0      |
| 1220–1380                  | 3.12                     | 3.42                  | 9.0       |
| 1380–1600                  | $1.74 \times 10^{-4}$    | $2.39 \times 10^{-2}$ | —         |

## 6.2 Region 570–700 cm<sup>-1</sup>

In this region, mainly dominated by the  $\nu_4$  fundamental band, the list contains a total of  $\sim 20.4$  million lines with the integrated intensity of  $1.40 \times 10^{-18}$  cm per molecule. The three main HBs which fall in this region are  $\nu_2 + \nu_4 - \nu_2$ ,  $2\nu_4 - \nu_4$  and  $2\nu_2 + \nu_4 - 2\nu_2$  with integrated intensities of  $3.19 \times 10^{-19}$ ,  $1.96 \times 10^{-19}$  and  $3.97 \times 10^{-20}$  cm per molecule. There also exist other bands such as  $\nu_3 - \nu_4$ ,  $\nu_2 + 2\nu_4 - \nu_2 - \nu_4$ , etc. whose contributions are respectively about 65%, 59%, etc. of  $2\nu_2 + \nu_4 - 2\nu_2$ . At 296 K, the HB contribution represents  $\sim 44\%$  of the total opacity in this region making analyses with “traditional” spectroscopic models very challenging. Fig. 7 (top) shows comparisons between theory and the observed PNNL spectra and confirms a very good agreement in the integrated absorbance of about 3% (see Table 4).

## 6.3 Region 980–1220 cm<sup>-1</sup>

In this spectral range, the list contains about 21 million lines and is dominated by the  $\nu_2 + \nu_4$ ,  $2\nu_2 + \nu_4 - \nu_2$  and  $\nu_2 + 2\nu_4 - \nu_4$  bands around 1062 cm<sup>-1</sup>. Their integrated intensities are respectively  $4.72 \times 10^{-20}$ ,  $9.10 \times 10^{-21}$  and  $1.40 \times 10^{-21}$  cm per molecule. The hot band  $2\nu_1 - \nu_4$  around 1177 cm<sup>-1</sup> has the integrated intensities of  $8.92 \times 10^{-22}$  cm per molecule. The HBs contribute at  $\sim 25\%$  of the total opacity in this region. Again, comparison with PNNL spectra in Fig. 7 (bottom panel) shows a very good agreement.

## 6.4 Region 1220–1380 cm<sup>-1</sup>

The bands of this region containing about 61 million lines in our list clearly dominate the absorption CF<sub>4</sub> spectrum in the infrared. The main absorption corresponds to the fundamental stretching band  $\nu_3$  but also, to a lesser extent, to the overtones  $2\nu_4$  and  $3\nu_2$  (resp. about 14.5% and 2.3% of the  $\nu_3$  opacity). The most active HBs falling in this region are  $\nu_2 + \nu_3 - \nu_2$ ,  $\nu_3 + \nu_4 - \nu_4$ ,  $\nu_2 + 2\nu_4 - \nu_2$ ,  $4\nu_2 - \nu_2$ ,  $2\nu_2 + \nu_3 - 2\nu_2$ ,  $3\nu_4 - \nu_4$ , etc. that contribute to 22.0, 11.5, 3.6, 3.2 and 2.3% of the  $\nu_3$  opacity, respectively.

This means that even bands with three vibrational quanta give non negligible opacity contributions, making an analysis with effective empirical Hamiltonian and effective fitted transition moments very complicated (see Section 7.1 for more discussion). The total integrated intensity at this range is  $1.91 \times 10^{-16}$  cm per molecule with  $\sim 60\%$  originating from CBs. Fig. 8 (top) displays a very dense absorption pattern of the CF<sub>4</sub> spectrum in the  $\nu_3$  region. The bottom panel shows very good agreement between the integrated absorbance and the PNNL measurements, within 0.9%, where the low-resolution theoretical spectrum simulations match nearly perfectly the observed cross-sections.

## 6.5 Region 1380–1600 cm<sup>-1</sup>

This is the region of the  $\nu_1 + \nu_4$  combination band centered around 1540 cm<sup>-1</sup> that has an integrated intensity of  $7.69 \times 10^{-19}$  cm per molecule. This band dominates  $\nu_1 + \nu_2 + \nu_4 - \nu_2$  and  $\nu_1 + 2\nu_4 - \nu_4$  by factors of 4.5 and of 6.7, respectively. The fourth strongest absorption is due to  $2\nu_2 + \nu_4$  around 1500 cm<sup>-1</sup>. The absorbance in the region is by two orders of magnitude weaker than in the  $\nu_3$  region. The HB transition represents here about 42% of the total opacity. The total number of lines in our lists is about 23.5 million. A detailed comparison with PNNL spectra is given in Fig. 9, where again we find an overall agreement with the observed cross-section, the error on the integrated absorbance being 3.8%.

## 6.6 Region 1600–1800 cm<sup>-1</sup>

In this region, the line list contains about 24 million transitions corresponding mainly to three CBs and two HBs that are assigned as  $\nu_2 + \nu_3$ ,  $\nu_2 + 2\nu_4$ ,  $4\nu_2$  and  $2\nu_2 + \nu_3 - \nu_2$ ,  $\nu_2 + \nu_3 + \nu_4 - \nu_4$ . CBs represent about 85% of the opacity in this region. Fig. 10 gives the respective contributions of the cold and hot band (top panel) as well as a detailed comparison with PNNL spectra (bottom panel) which proves again the consistency of our calculations and the validity of the *ab initio* DMS (the obs.–calc. error of  $\sim 2.5\%$  in the integrated absorbance).

## 6.7 Region 1800–2000 cm<sup>-1</sup>

This dense region (about 125 million lines in the theoretical list) is dominated by the  $\nu_3 + \nu_4$  combination band around 1913 cm<sup>-1</sup> that has an integrated intensity of  $1.40 \times 10^{-19}$  cm per molecule. The second strongest absorption is due to  $3\nu_4$  around 1890 cm<sup>-1</sup>, which is four times less important than  $\nu_3 + \nu_4$  in terms of opacity. The next five strongest absorptions (in decreasing order) are due to the  $\nu_2 + \nu_3 + \nu_4 - \nu_2$ ,  $\nu_3 + 2\nu_4 - \nu_4$ ,  $\nu_2 + 3\nu_4 - \nu_2$ ,  $4\nu_4 - \nu_4$  and  $2\nu_3 - \nu_4$  bands with integrated intensities of  $2.70 \times 10^{-20}$ ,  $1.37 \times 10^{-20}$ ,  $5.1 \times 10^{-21}$ ,  $3.69 \times 10^{-21}$  and  $2.64 \times 10^{-21}$  cm per molecule. The cold band  $3\nu_2 + \nu_4$  around 1937 cm<sup>-1</sup> is completely diluted among all the other ones. The comparison with PNNL is given in Fig. 11, the error in integrated absorbance being very small ( $\sim 0.4\%$ ).

## 6.8 Region 2000–2280 cm<sup>-1</sup>

This strongest absorption in this dense region, composed of about 168 million lines in the list, is due to  $\nu_1 + \nu_3$  with the integrated intensity of  $5.00 \times 10^{-19}$  cm per molecule. The  $\nu_1 + \nu_2 + \nu_3 - \nu_2$  and  $\nu_1 + 2\nu_4$  bands are five times weaker than this band while

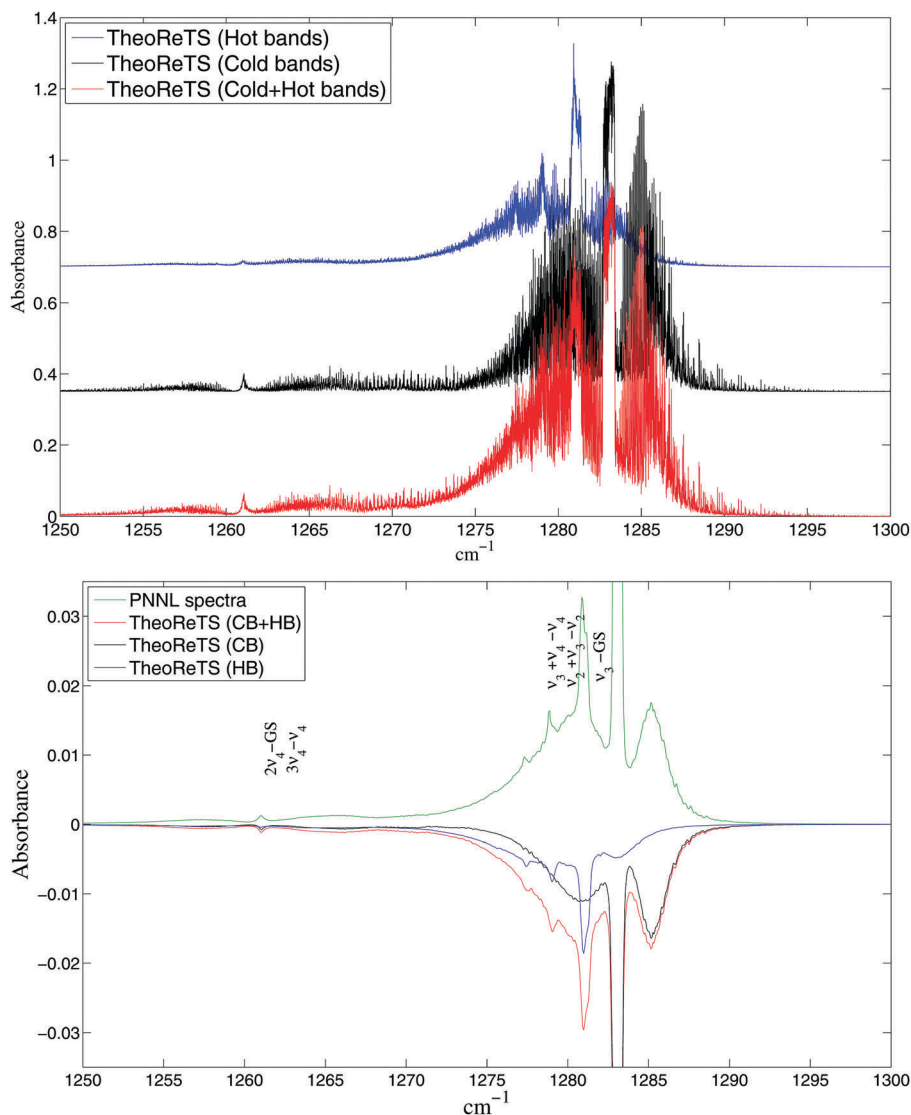


Fig. 8 Theoretical absorbance computed at 296 K with the contribution of cold and hot band transitions (top panel). Example comparison of theoretical (CB, HB and CB + HB transitions) absorption spectra of  $\text{CF}_4$  with experimental PNNL records at room temperature in the  $\nu_3$  region around  $1250\text{--}1300\text{ cm}^{-1}$  (bottom panel). The main CBs and HBs are also indicated in vertical text (see also Section 6.4).

$\nu_1 + \nu_3 + \nu_4 - \nu_4$ ,  $\nu_1 + \nu_2 + 2\nu_4 - \nu_2$ ,  $\nu_1 + 3\nu_4 - \nu_4$ ,  $\nu_1 + 2\nu_2 + \nu_3 - 2\nu_2$  and  $2\nu_2 + \nu_3$  are one order of magnitude weaker. The impressive complexity of the  $\text{CF}_4$  spectra is due to the more than 40 CBs and HBs that fall in this region, including three, four, five and even six quanta bands such as  $6\nu_2 - \nu_2$  which is by three orders of magnitude weaker than  $\nu_1 + \nu_3$ . Despite such complexity, the comparison with PNNL is very good (see Fig. 11, middle and bottom panels) and the error in integrated absorbance is only 1.2%. The HB transition represents about 44% of the total opacity in this region.

### 6.9 Region $2280\text{--}2700\text{ cm}^{-1}$

This is again a very dense and rich region composed by more than 760 million lines at 296 K. Among the strongest bands, we have the CBs  $2\nu_3$  (the strongest one, with an integrated intensity of  $4.98 \times 10^{-19}\text{ cm}^2\text{ per molecule}$ ),  $\nu_3 + 2\nu_4$ ,  $4\nu_4$  and  $3\nu_2 + 2\nu_4$  around

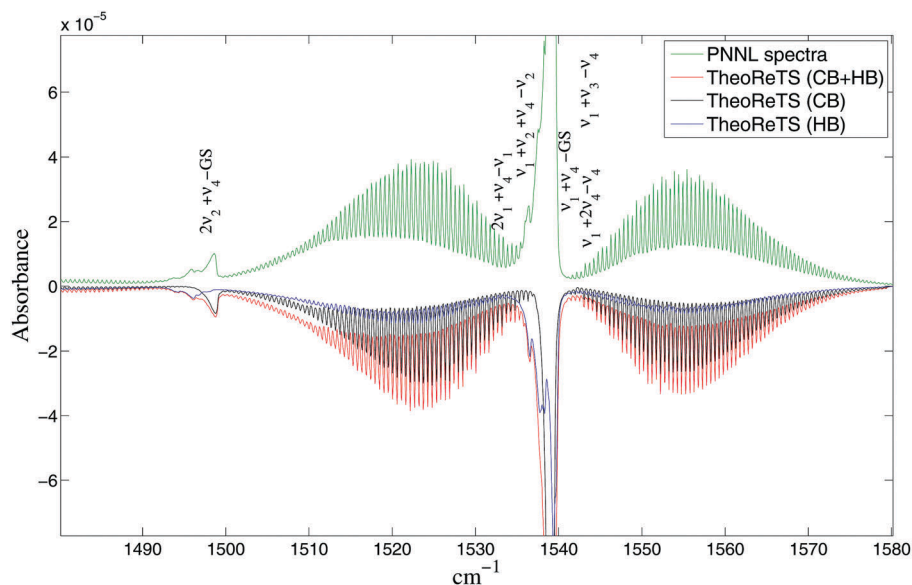
$2560\text{ cm}^{-1}$  and  $2\nu_1 + \nu_4$  around  $2445\text{ cm}^{-1}$ . Concerning the HBs, we find  $\nu_2 + 2\nu_3 - \nu_2$ ,  $2\nu_3 + \nu_4 - \nu_4$ ,  $\nu_2 + \nu_3 + 2\nu_4 - \nu_2$ ,  $\nu_3 + 3\nu_4 - \nu_4$ ,  $2\nu_2 + 2\nu_3 - 2\nu_2$ ,  $\nu_2 + 4\nu_4 - \nu_2$ , etc. contributing to 20, 10, 7, 3.6, 2.5 and 1.9% of the  $2\nu_3$  intensity. Fig. 12 (top and middle panels) shows very good agreement with the observed PNNL cross-sections.

### 6.10 Region $2700\text{--}2950\text{ cm}^{-1}$

This region contains about 165 million lines and is dominated by the two CB  $\nu_1 + \nu_3 + \nu_4$  and  $\nu_1 + 3\nu_4$  that are 70% stronger than the four HBs  $\nu_1 + \nu_2 + \nu_3 - \nu_2$ ,  $\nu_1 + \nu_3 + 2\nu_4 - \nu_4$ ,  $\nu_1 + \nu_2 + 3\nu_4 - \nu_2$  and  $\nu_1 + 4\nu_4 - \nu_4$ . Once again, Fig. 12 (bottom panel) shows a very good agreement with PNNL (error  $\sim 0.3\%$  for integrated absorbance).

### 6.11 Region $2950\text{--}3300\text{ cm}^{-1}$

This weak, but dense, region contains more than 250 million lines and is clearly dominated by three CBs, namely  $2\nu_3 + \nu_4$ ,



**Fig. 9** Example comparison of theoretical (CB, HB and CB + HB transitions) absorption spectra of CF<sub>4</sub> with experimental PNNL records at room temperature in the  $\nu_1 + \nu_4$  region around 1480–1580 cm<sup>-1</sup>. The main cold and hot bands are also indicated in vertical text (see also Section 6.5).

## 6.12 Region 3300–3950 cm<sup>-1</sup>

In this region, composed by 290 million lines, the  $\nu_1 + 2\nu_3$  band around  $3450\text{ cm}^{-1}$  and the  $3\nu_3$ ,  $2\nu_3 + 2\nu_4$ ,  $\nu_3 + 4\nu_4$  bands around  $3800\text{ cm}^{-1}$  give the strongest contributions. The first HBs are  $\nu_1 + \nu_2 + 2\nu_3 - \nu_2$ ,  $\nu_1 + \nu_2 + \nu_3 - \nu_2$  and  $\nu_2 + 3\nu_3 - \nu_2$  and are one order of magnitude weaker than the cold bands. Fig. 13 (bottom panel) confirms the relatively small error on the integrated intensities ( $\sim 4.5\%$ ) compared to PNNL. Line-by-line analyses of experimental spectra in such a complex region involving 6 quanta bands using traditional polyad models would be hardly feasible, at least at  $T = 296\text{ K}$ ; only low- $T$  spectra could help the assignments.

## 7 Discussion

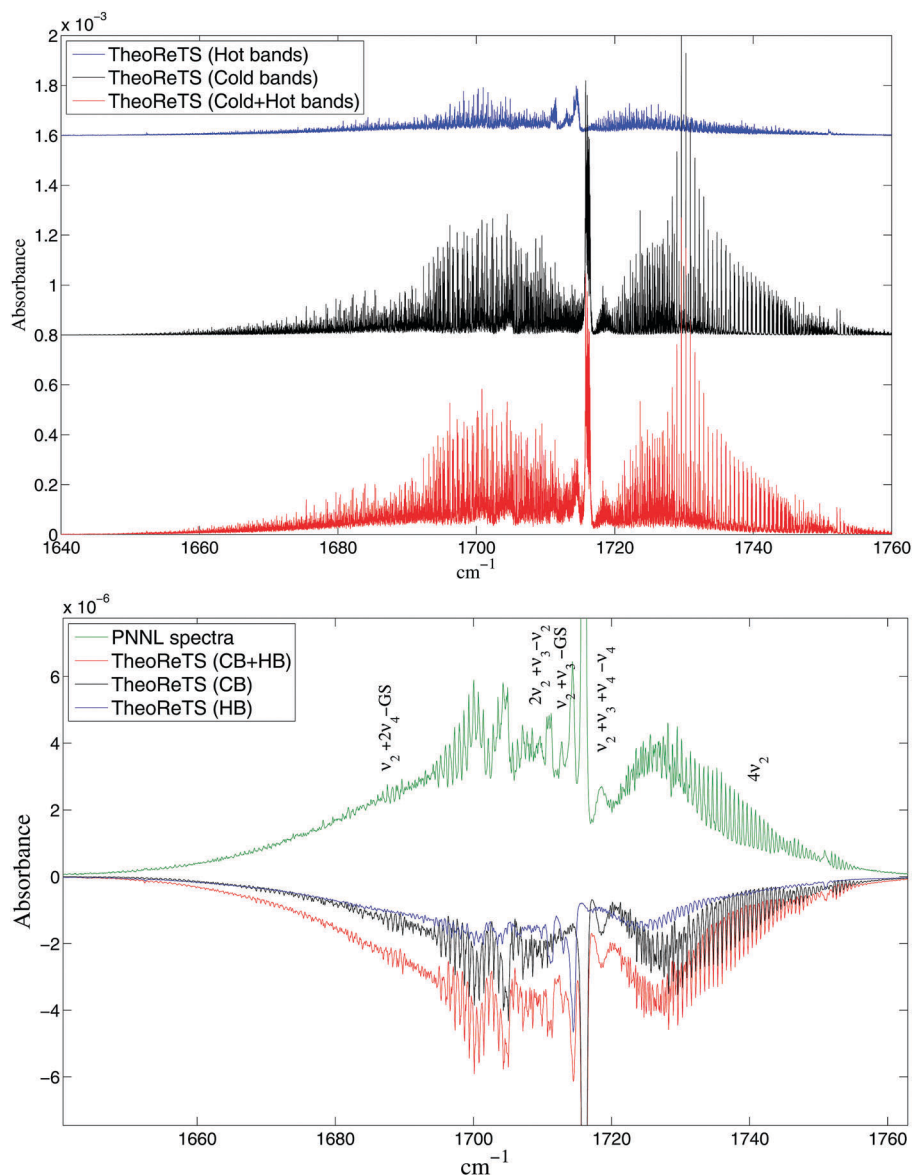
### 7.1 Global *ab initio* versus phenomenological spectroscopic models

As was mentioned in Section 2, much effort has been devoted to attempt to understand the complex polyad structure of  $\text{CF}_4$ . Analysis of ro-vibrational lines has been previously conducted using phenomenological effective spectroscopic models (ESMs) that involve empirically determined parameters fitted to experimental line positions and line intensities. This is the first of the two possible approaches. For molecules exhibiting quite regular patterns, such empirical models have proved their efficiency

these past few decades in spectroscopic data reduction and in the construction of well-known databases such as HITRAN<sup>2</sup> or GEISA.<sup>3</sup> They are particularly well adapted to provide very precise line positions, approaching experimental accuracy. However, in the case of molecules with low-frequency vibrational modes and relatively heavy “edge atoms” as for CF<sub>4</sub>, the vibrational polyad structure is not well defined and high-*J* transitions make the spectra very crowded. This makes the empirical line-by-line analyses much more complicated. Due to the huge number of overlapping lines forming dense and complex patterns composed of both CBs and HBs, the effective models could face their intrinsic limit, even though several bands were fitted simultaneously.

Very recently, a partial high-resolution analysis of  $\text{CF}_4$  spectra in the infrared has been performed<sup>62</sup> to better understanding complex absorption features. This led to the first determination of the effective parameters for 17 rovibrational bands included in a fitted effective Hamiltonian model. Their analysis was essentially focused on line positions while effective dipole parameters were fixed by manual adjustment. A part of the result of Carlos *et al.*<sup>62</sup> was used to update the Dijon University TFMcCaSDa database,<sup>102</sup> up to  $1700\text{ cm}^{-1}$ , with an intensity cut-off of  $10^{-24}\text{ cm}$  per molecule that provided the most extended line list for the  $\text{CF}_4$  in the infrared range currently computed in the framework of an ESM.

A comparison of the experimental PNNL spectra with *ab initio* born variational calculations (this work) and with the TFMCaSDa database clearly shows limitations of the effective models in terms of extrapolation (see Fig. 14). The medium-resolution absorbance simulations are quite similar in the narrow range around  $1280\text{ cm}^{-1}$  included in the EMS empirical line-by-line fits but are very different beyond this region (see Fig. 14, upper panels). Fig. 7–13 show a good agreement between our *ab initio* intensities and experiments in the whole wavenumber



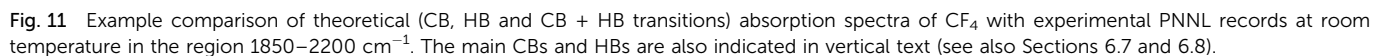
**Fig. 10** Theoretical absorbance computed at 296 K with the contribution of cold and hot band transitions (top panel). Example comparison of theoretical (CB, HB and CB + HB transitions) absorption spectra of  $\text{CF}_4$  with experimental PNNL records at room temperature in the  $\nu_2 + 2\nu_4/\nu_2 + \nu_3$  region around  $1640\text{--}1760\text{ cm}^{-1}$  (bottom panel). The main CBs and HBs are also indicated in vertical text (see also Section 6.6).

range of available PNNL cross-sections. In contrast, the EMS line list shows erratic absorbance behaviour below  $1250\text{ cm}^{-1}$  and above  $1300\text{ cm}^{-1}$  in a qualitative disagreement with the PNNL experiment (Fig. 14, right-hand panels). The deviations in hot bands are particularly large (Fig. 14, left-hand upper panel) leading to quite different temperature dependence of the opacity. The reasons for these disagreements could be the following.

- First, very few experimental assignments are available in the literature for the HB transitions and for the corresponding excited vibrational states. CB and HB transitions are often blended together, consequently it was not possible to measure their individual line intensities.<sup>62</sup> If the line strength information on HBs was not included in the EMS, their contributions to the observed cross-sections could be “absorbed” by effective

dipole moments of CBs. For example, in the  $1220\text{--}1380\text{ cm}^{-1}$  region the integrated PNNL absorbance using the TFMCaSDa database<sup>102</sup> is higher than ours (2.49 vs. 2.24) while it is inverted for HBs (0.62 vs. 1.16).

- As only partial information on energy levels was obtainable from empirically assigned lines, the ESM calculations are not complete and include extrapolations beyond the range of observed data that are not always reliable. This was the case of energy patterns in the region  $2080\text{--}2160\text{ cm}^{-1}$  given in Fig. 12 of ref. 62. In this range where no experimental assignments were included, the extrapolated level exhibited drastically different behaviour from the *ab initio* levels (Fig. 2) at large  $J$ -values that may explain the disagreement of ESM with observations in Fig. 14. This is also corroborated by comparing Fig. 2 of ref. 60



with our Fig. 2 (top panel) or Fig. 8 of ref. 62 with our Fig. 2 (bottom panel). In one case, the three vibrational sublevels  $3\nu_2(A_{1,A_2},E)$  were missing in ref. 60 and in the other case the three vibrational sublevels  $4\nu_2(A_{1,A_2},E)$  were missing in ref. 62.

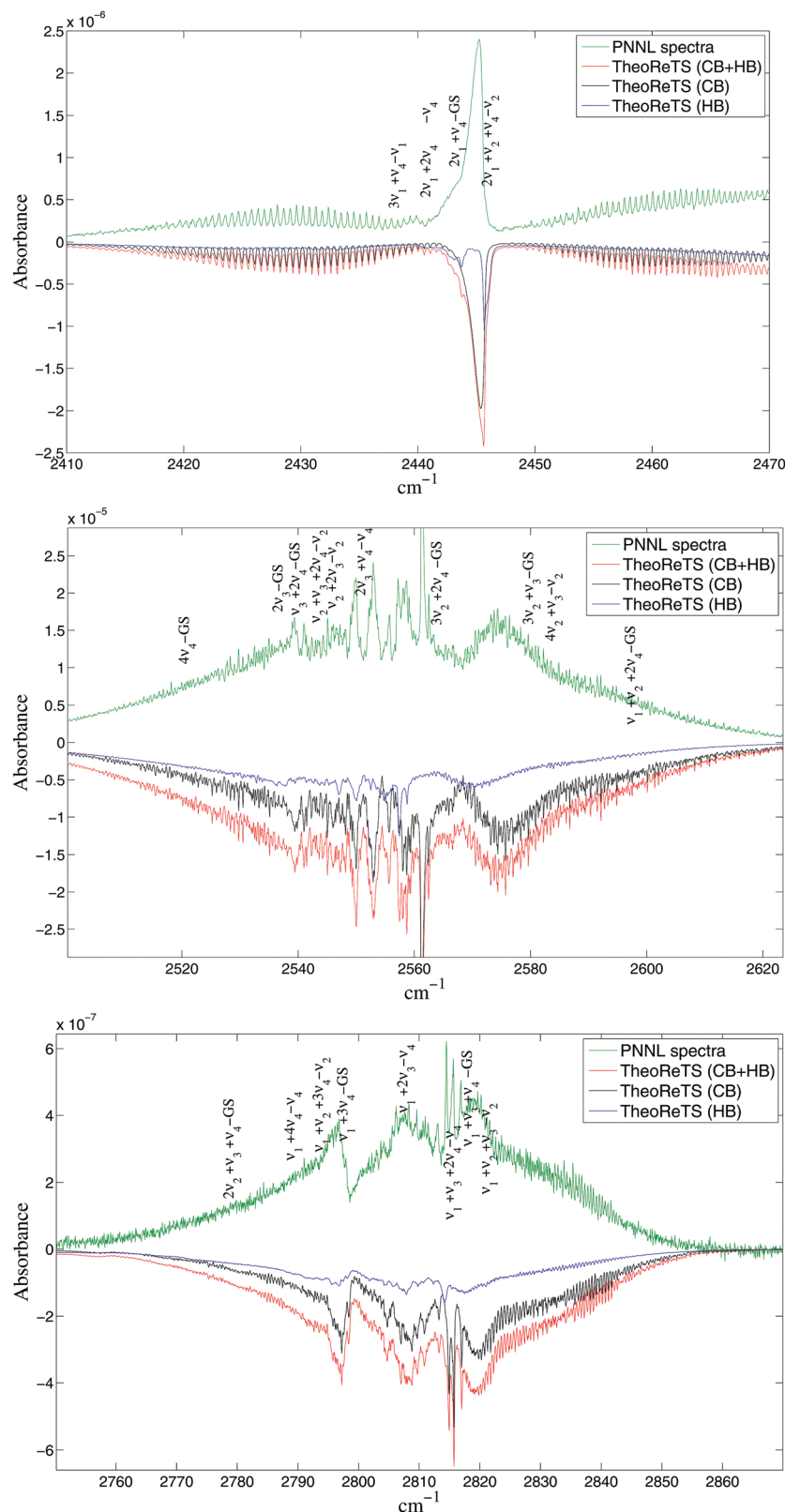


Fig. 12 Example comparison of theoretical (CB, HB and CB + HB transitions) absorption spectra of  $\text{CF}_4$  with experimental PNNL records at room temperature in the region  $2500\text{--}2870\text{ cm}^{-1}$ . The main CBs and HBs are also indicated in vertical text (see also Sections 6.9 and 6.10).

Note that one of the two sublevels  $\nu_2 + 2\nu_4(A_1)$  should read  $\nu_2 + 2\nu_4(A_2)$  in Fig. 8 of ref. 62.

• The term “global” in some previous ESM analyses has to be taken with caution because many CBs and HBs falling in the

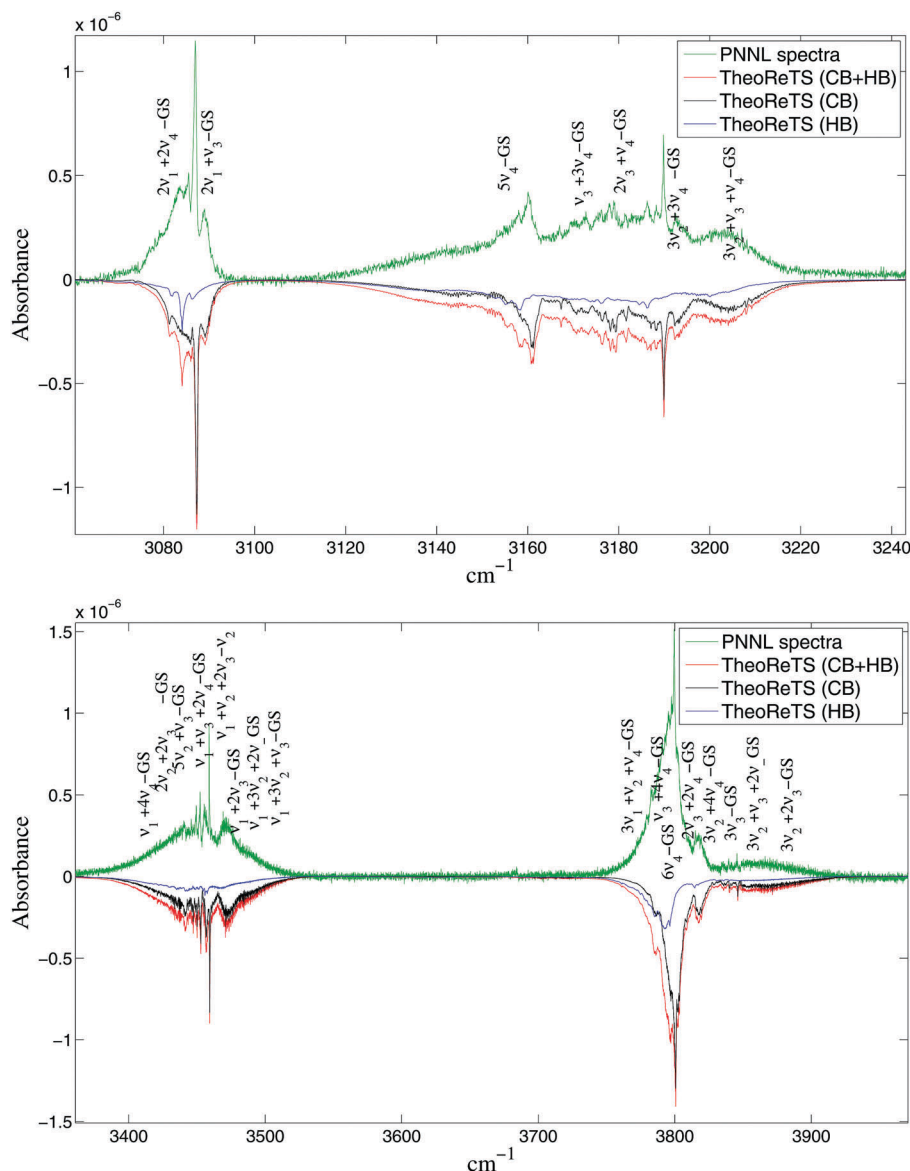


Fig. 13 Sample comparison of theoretical (CB, HB and CB + HB transitions) absorption spectra of  $\text{CF}_4$  with experimental PNNL records at room temperature in the region  $3400\text{--}3950\text{ cm}^{-1}$ . The main CBs and HBs are also indicated in vertical text (see also Section 6.12).

same spectral range have been omitted and many resonance coupling parameters had to be neglected. Modelling of “dark states” (which are not directly observable, but perturb active bands) is one of the major obstacles for a purely empirical approach. Variational *ab initio* based calculations include an entire set of bands contributing to the opacity and account for all interactions among them. For example, 17 rovibrational bands have been included in the last ESM analysis<sup>62</sup> up to  $2200\text{ cm}^{-1}$ ; in our variational calculations, we predict at least 160 rovibrational bands of the type  $\sum_i \alpha_i \nu_i - \sum_j \beta_j \nu_j$  in the same region. Up to  $4000\text{ cm}^{-1}$ , almost 800 vibrational states need to be included, leading to  $\sim 300$  “strong” transitions of the type  $\sum_i \alpha_i \nu_i - \sum_j \beta_j \nu_j$  and many more in the QC absorption.

Though *ab initio* methods do not achieve high-resolution accuracy for individual line positions, the comparison with

experimental PNNL spectra (Fig. 7–13) clearly show that our calculations provide more complete and reliable intensities both for CBs and HBs than the available ESM line lists in a large wavenumber range.

## 7.2 Hints for further improving line-by-line analyses

$\text{CF}_4$  belongs to the category of molecules where a complete line-by-line analysis of the overall rotation–vibration spectrum in the infrared will take many years or even decades, similar to the long history of methane studies (see for example the review by Brown *et al.*<sup>104</sup>) but with a much larger number of lines. One of the well known issues is that empirical ESM fits behave as mathematically ill-defined problems (see ref. 105 and 106) because the effective Hamiltonian matrices and coupling parameters are not unambiguously determined from experimental levels.

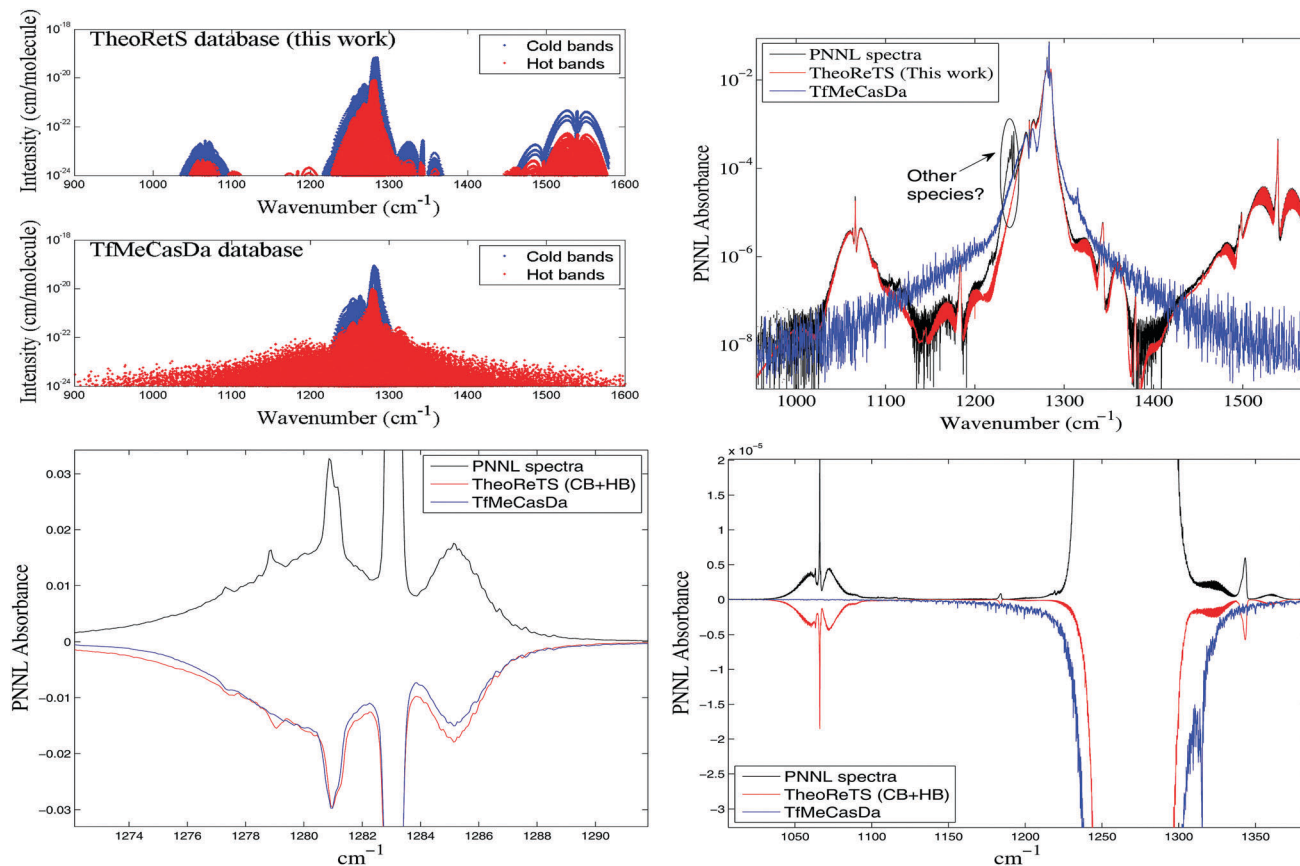


Fig. 14 Comparison of the TheoReTS and TFMeCasDa<sup>102</sup> (<http://vamdc.icb.cnrs.fr/PHP/tfmecasda.php>, downloaded on 20/05/2018) line intensities (top panel, left-hand side) in the range 900–1600 cm<sup>-1</sup> and detailed comparison of the simulated absorption spectra of CF<sub>4</sub> with experimental PNNL records at room temperature.

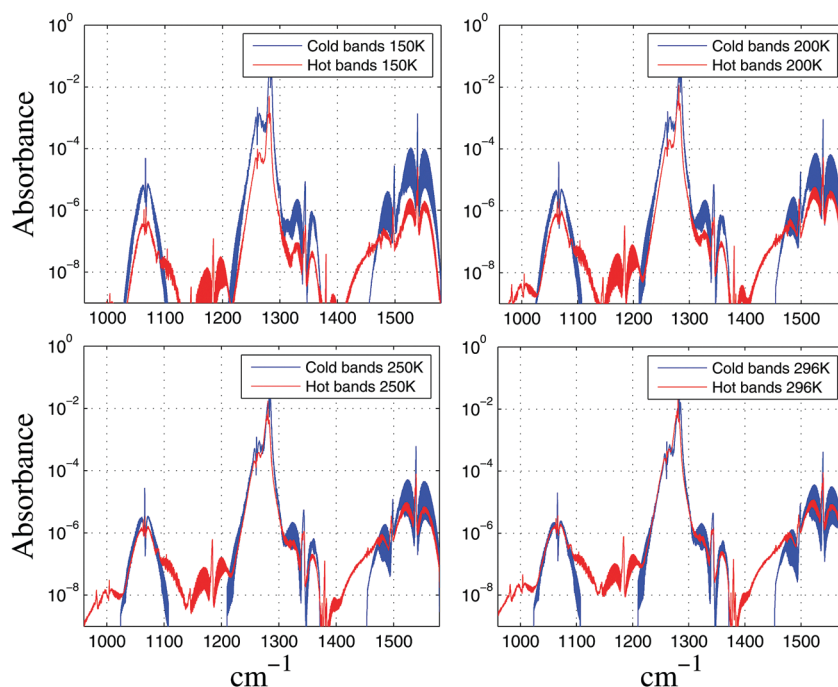


Fig. 15 Temperature effect on the absorption spectra of CF<sub>4</sub> in the range 1000–1500 cm<sup>-1</sup> and impact on the HB contributions from 150 to 296 K.

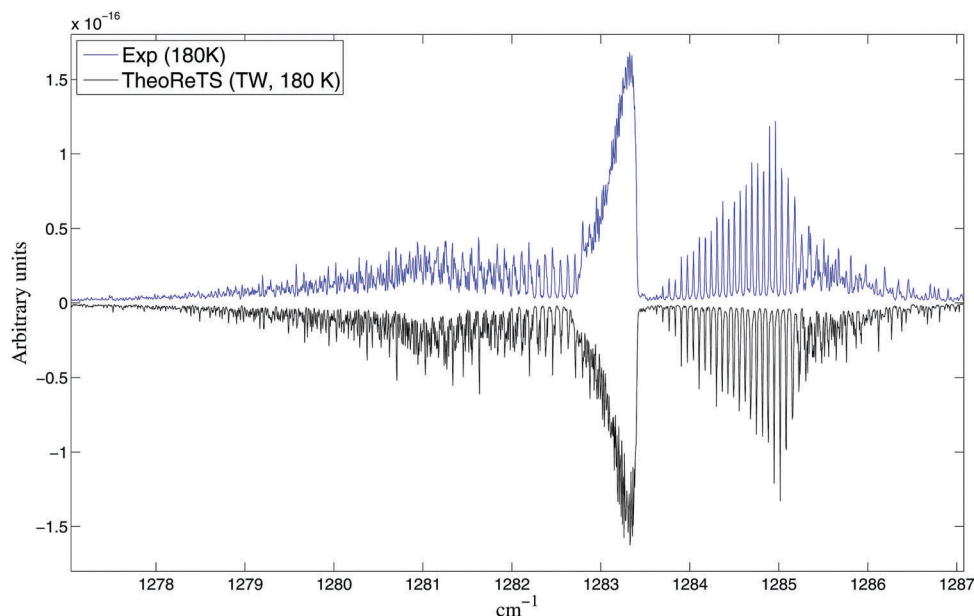


Fig. 16 Comparison of the simulated spectra (this work) of  $\text{CF}_4$  with experiment at 180 K, in the  $\nu_3$  region.

This can lead to false resonance mixing of wavefunctions and consequently erratic intensity borrowing effects among coupled strong and weak bands.

In order to make the analyses of high-resolution spectra more robust, the *ab initio* effective spectroscopic models adapted to successive spectral ranges can be accurately constructed from the potential energy surface by the high-order contact transformation method.<sup>64,107</sup> For the sets of states contributing to a considered range (including “dark” states) this will provide initial physically meaningful coupling parameters and wavefunctions, which were poorly determined in a purely empirical approach. Diagonal parameters, including band centers and rotational constants can be optimized by fine tuning to experimental spectra, while band intensities can be evaluated using *ab initio* DMS. Such a strategy has been recently proven efficient for  $\text{CH}_4$ <sup>107</sup> and an analysis of room-temperature  $\text{CF}_4$  Fourier transform spectra in the region 1600–1800  $\text{cm}^{-1}$  is currently in progress using this method.

It will be also useful to record cooled  $\text{CF}_4$  spectra down to  $T = 100$  or 150 K, in order to simplify CB analyses by removing completely or partially HB transitions. Fig. 15 shows the temperature effect when cooling spectra at 150, 200, 250 K. We can see that HB contributions are reduced by one order of magnitude from  $T = 296$  to  $T = 150$  K. Thus recording low- $T$  spectra in a large range in the future would be of great help for progress in the line-by-line assignment procedure. Fig. 16 gives a comparison of the experimental and calculated spectra at 180 K in the  $\nu_3$  interval.

## 8 Conclusion

In this work, we have shown that an efficient variational method for nuclear motion calculations permitted reliable predictions

of rotationally resolved spectra and a full understanding of the hot band patterns in the wide IR spectral range for a five-atomic molecule with relatively heavy edge F atoms. In previous studies<sup>23,34,75,90,108</sup> such quality of theoretical IR spectra has been achieved for lighter molecules only with hydrogen edge atoms like  $\text{CH}_4$ ,  $\text{SiH}_4$ , or  $\text{GeH}_4$ . A big difference is that  $\text{CF}_4$  has much lower vibrational modes and much higher  $J$ -levels are populated making the HB spectral patterns significantly more complicated. We show that the strong absorption features are superimposed with a kind of quasi-continuum already at room temperature conditions, similarly to that observed<sup>35</sup> and calculated<sup>32–34</sup> in hot methane spectra at about 1000 K.

Minor empirical refinement of the equilibrium geometry and of force constants of the PES combined with purely *ab initio* DMS (provided as ESI†) yielded a correct global description of the positions and shapes of the overtone and combination bands up to five or six vibrational quanta validated by detailed comparisons with experimental PNNL spectra.

This was used to construct the first accurate and complete *ab initio* based line lists for  $\text{CF}_4$  in the range 0–4000  $\text{cm}^{-1}$ , containing absorption bands that are the most active in IR. In order to converge the IR opacity at room temperature, it was necessary to include a high rotational quantum number up to  $J = 80$  resulting in 2 billion rovibrational transitions. In practical terms, these data can be used for a reliable modelling of the full IR opacity of this long-lived greenhouse molecule under various temperature conditions.

The data obtained in this work will be made available through the TheoReTS<sup>29</sup> information system (<http://theorets.univ-reims.fr>, <http://theorets.tsu.ru>) that contains *ab initio* born line lists and provides a user-friendly graphical interface for a fast simulation of the absorption cross-sections and radiance.

In order to make the cross-section simulation faster, we have partitioned our data in two parts: (a) strong + medium line lists

with lower energy levels for calculation of selective absorption features that can be used at various temperatures and (b) compressed “super-line” libraries of very weak transitions contributing to the quasi-continuum modelling. The latter ones will be available in TheoReTS, from  $T = 100$  K to 300 K by steps of 20 K and also for 400 K. Extrapolations for higher temperatures can be produced by request if necessary for some particular applications.

We also propose solutions (see Section 7.2) to combine completeness and line position accuracy for future *ab initio* guided high-resolution spectra analyses. Further progress in the description absorption/emission for fluoro-containing greenhouse molecule in a wider range and for applications under various conditions could benefit from a collaboration of spectroscopists, theoretical chemists, and specialists in atmospheric modelling.

## Conflicts of interest

There are no conflicts to declare.

## Appendix A: construction of a “pruned” vibrational basis set

- First, a “pruned basis” is defined by selecting only a limited set of relevant primitive vibrational functions as

$$F_k^\Gamma(n) \Leftrightarrow \sum_{i=1}^4 \kappa_i v_i \leq n, \quad (\text{A1})$$

with  $v_i = 0, \dots, n$ , for a symmetry block  $\Gamma (\equiv C_v)$ . Here  $\kappa_i$  are certain weight coefficients chosen to select and optimize an appropriate number of stretching and bending basis functions for each mode. The associated vibrational subspace will be denoted as  $\mathcal{F}_p^\Gamma(n)$  such that  $\dim(\mathcal{F}_p^\Gamma) = D_1^\Gamma$  that is the basis in the block  $\Gamma$  will contain  $D_1^\Gamma$  vibrational functions. Here the “ $p$ ” index stands for a pruned subspace. This pruning scheme amounts to discarding “unnecessary” functions having a very small impact on calculations or probing non-physical regions of the PES. For example, the matrix element  $\langle 9,0,10,0 | a_1^3 a_4 | 6,0,11,0 \rangle$  computed in the unconstrained basis (12) couples vibrational states located at 20 900 and 19 500  $\text{cm}^{-1}$  for  $^{12}\text{CF}_4$ , which is quite useless for practical calculations. Setting  $n = 11$ , such matrix elements would be not considered from the pruning scheme (A1) with all  $\kappa_i = 1$ . In the case of  $T_d$  symmetry, the number of pruned basis functions in the  $F$  ( $F_1$  or  $F_2$ ) symmetry blocks is  $D_1^F \sim 10^5$  ( $\approx 100$  GB of computer memory against hundreds of TBytes from eqn (12)).

- Pruning decreases the number of basis functions but is still not sufficient to make rovibrational calculations feasible, especially for the very large values of the rotational angular momentum involved in hot calculations and/or heavy molecules. Typically, energy levels have to be computed up to  $J \approx 80$  to predict  $\text{CF}_4$  room temperature spectra.

In the next step, we apply our two-step procedure to reduce the length  $D_1^\Gamma$ . First, the Hamiltonian (9) is partitioned as

$H = H_v + H_{vr}$  and the eigenvalue equation  $H_v \Psi_v^\Gamma = E_v^\Gamma \Psi_v^\Gamma$  is variationally solved with a primitive  $F_k^\Gamma(n)$  and the vibrational eigensolutions  $\{E_v^\Gamma(n), \Psi_v^\Gamma(n)\}$  are stored. Here  $n$  is chosen to properly converge all necessary energy levels. The resulting anharmonic vibrational eigenfunctions  $\Psi_v$  for a given block  $\Gamma$  read

$$\Psi_v^\Gamma(n) = \sum_{\mathcal{F}(n)} c_i^\Gamma \Phi_{iv}^\Gamma. \quad (\text{A2})$$

Let us assume now that there exist a basis set  $F_{\kappa_2}^\Gamma(m) \in \mathcal{F}_r^\Gamma(m < n)$  of size  $\dim(\mathcal{F}_r^\Gamma) = D_2^\Gamma$  with  $D_2^\Gamma \ll D_1^\Gamma$  such that the  $J > 0$  vibration-rotation calculations are feasible. This amounts to project the anharmonic solutions (A2) onto the reduced subspace  $\mathcal{F}_r^\Gamma$  where the “ $r$ ” index stands for the reduced subspace. The resulting set of anharmonic vibrational reduced functions  $\{\Psi_v^{(n \rightarrow m)}\}$  are given by

$$\mathcal{G} \mathcal{P}^{(m)} \Psi_v^\Gamma(n) = \Psi_v^{\Gamma(n \rightarrow m)} = \sum_{\mathcal{F}_r(m)} c_i'^\Gamma \Phi_{iv}^\Gamma \quad (\text{A3})$$

and will be used for further  $J > 0$  calculations. Here  $\mathcal{P}^{(m)}$  is a projector onto the  $\mathcal{F}_r(m)$  subspace and  $\mathcal{G}$  is the Gram-Schmidt transformation.

Note that in the range of the observed energy levels the new expansion coefficients  $c_i'^\Gamma$  are quite close to the initial  $c_i^\Gamma$  coefficients so that we can write  $H_v \Psi_v^{\Gamma(n \rightarrow m)} \approx E_v^{\Gamma(n)} \Psi_v^{\Gamma(n \rightarrow m)}$ . Finally the size  $D_2^\Gamma$  of the reduced basis is of the order  $\sim 10^2$ – $10^3$  (only few Mb of memory for the  $J = 0$  problem) and the largest matrices to be diagonalized will be of size  $\sim 10^5$  for high  $J$  values. Thus the full problem has been scaled by several orders of magnitude.

## Appendix B: cut-off criteria in variational spectra calculations

### Hamiltonian expansion and vibrational basis set

Initially, the Eckart-Watson nuclear Hamiltonian was expanded up to the total powers  $p = (m_{k\sigma} + n_{k\sigma})$  in the tensor representation of eqn (7) and (8) with a subsequent reduction to  $p' = 7$  as described in ref. 77. In the case of methane, it has been shown that such ( $p = 14 \rightarrow p' = 7$ ) Hamiltonian reduction provides good convergence in energies up to five-quanta vibrational overtones and combination states.<sup>88</sup> Also, agreement within  $\sim 0.1 \text{ cm}^{-1}$  was obtained with respect to independent variational calculations using the exact kinetic energy operator and curvilinear internal coordinates.<sup>109</sup> The corresponding tensor expansion (9) contains 23 548 rovibrational terms.

The vibrational basis  $F_k(n = 14)$  in eqn (A1) was used with the weighting coefficient vector  $\kappa = (1.05; 1; 1.18; 1)$  to select an appropriate number of bending/stretching functions. The larger matrix to be considered in this step is of size  $D_1^{F_2} \sim 55\,000$ . For  $J > 0$  calculations, these functions have been further reduced to eqn (A3) using  $m = 7$  and the  $\kappa_1$ 's chosen below.

To ensure the completeness of our line lists at least up to  $T = 296$  K, a global set of variational vibration-rotation energy levels was computed up to  $J = 80$ . Let us recall that our recent hot methane calculations were practically converged at  $J = 67$

for  $T > 2000$  K. Much higher  $J$ -values required for  $\text{CF}_4$  made calculations very demanding though the wavenumber range considered in this work will not be as extended as that of methane. Thus a heavy molecule such as  $\text{CF}_4$  requires the same amount of computational effort at room temperature as lighter systems such as methane for very hot temperatures.

Since the sizes of the matrices become larger as  $J$  increases, we have adopted a strategy similar to that of ref. 32, 34, 78 and 87 which consists of partitioning the rovibrational problem as follows: calculations were performed using the weighting coefficient vector  $\kappa_1 = (1.38; 1; 1.38; 1)$  up to  $J = 60$  and  $\kappa_1 = (1.3; 1; 1.5; 1.1)$  from  $J = 61$  to up to 80. For the first case, the larger vibrational basis is of size  $D_2^{F_2} \sim 500$  while for the second case, we have  $D_2^{F_2} \sim 390$ . Combining also with the parameter %VSS defined in ref. 88, the size of the rovibrational blocks ( $J, C = F$ ),  $F = F_1$  or  $F_2$ , to be diagonalized does not exceed 100 000 such that it can be easily managed with a standard modern PC.

### Cut-off for intensities

For the present study, because of the increasing number of vibrational excitations involved in overtones and combination bands, we concluded that it was reasonable to limit the line list to the spectral range  $[0-4000] \text{ cm}^{-1}$ . Beyond  $4000 \text{ cm}^{-1}$ , the line intensities are very weak and not clearly observable in existing experimental spectra (such as PNNL cross-section libraries<sup>103</sup>) due to the decreasing signal/noise ratio. Because the  $\text{CF}_4$  room- $T$  calculations are almost as demanding as the  $\text{CH}_4$  high- $T$  calculations, we have adopted a strategy as that of ref. 34 producing full theoretical spectra. Completeness of line lists is essential for appropriate atmospheric retrievals. At least three criteria are required to properly converge opacity calculations (or integrated intensities in other words). The first one is to choose a  $J_{\text{max}}$  that can be evaluated from the temperature dependence of the partition function. Here the value  $J_{\text{max}} = 80$  turns out quite reasonable to converge  $Q(296)$  and also the integrated intensities for various spectral ranges (see Fig. 5). For example, in the range  $550-700 \text{ cm}^{-1}$ , the integrated intensity up to  $J = 80$  is  $4.010 \text{ cm per molecule}$ . At  $J_{\text{max}} = 75$ , this value is  $3.970 \text{ cm per molecule}$  and at  $J_{\text{max}} = 79$ , we have  $4.005 \text{ cm per molecule}$ . So we have convergence errors of 1.03% and 0.13% at  $J_{\text{max}} = 75$  and  $J_{\text{max}} = 79$ , respectively. Similarly, in the  $980-1220 \text{ cm}^{-1}$  region, we have errors of 0.43% and 0.07% if  $J_{\text{max}} = 75$  and 79, respectively. Higher  $J$  values will not bring additional opacity at 296 K. The second one is to choose the maximum lower state energy cut-off that will govern the total number of HB transitions. After some trial tests, we have chosen  $E_{\text{low}}^{\text{max}} = [2700 + 0.19(J + 1)] \text{ cm}^{-1}$ . This value permitted to account for all HBs that produce a significant contribution to the integrated intensities in the range from  $T = 80 \text{ K}$  to  $T = 400 \text{ K}$ . Finally last but not least is to choose the intensity cut-off  $I_{\text{cut}}^{\text{max}}$ . This value is crucial and will control the total number of lines in the construction of the final line lists. Following ref. 34, we introduce wavenumber-dependent intensity cut-off functions  $I^{\text{cut}}(W)$  by splitting the whole wavenumber range  $W = 0-4000 \text{ cm}^{-1}$  in fifteen intervals  $R_k$  ( $k = 1, \dots, 15$ ), each one roughly corresponding to the most

abundant band systems (see Fig. 4). To define  $I^{\text{cut}}(W)$  for each interval, the function should mimic the shape of the absorption spectrum so that we can write  $I^{\text{cut}}(W) = \sum_k I_k^{\text{cut}}(W)$  (see Fig. 4, black curve). Appropriate analytical forms are given by Gaussian functions as  $I_k^{\text{cut}}(W) = I_k^{(m)} \exp(-2(W - W_k^{(m)})/A_k^2)$  where  $W_k^{(m)}$  is a wavenumber corresponding to the maximum absorption in every  $R_k$ , and  $I_k^{(m)}$ ,  $A_k$  are adjustable parameters so that the integrated intensity in each  $R_k$  converge. The  $I_k^{(m)}$ 's vary from  $10^{-31}$  to  $5 \times 10^{-27} \text{ cm per molecule}$ .

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Preliminary analysis of the interacting pentad bands ( $\nu_2 + 2\nu_4$ ,  $\nu_2 + \nu_3$ ,  $4\nu_2$ ,  $\nu_1 + 2\nu_2$ ,  $2\nu_1$ ) of  
CF<sub>4</sub> in the 1600 – 1800 cm<sup>-1</sup> region

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# Preliminary analysis of the interacting pentad bands ( $\nu_2 + 2\nu_4$ , $\nu_2 + \nu_3$ , $4\nu_2$ , $\nu_1 + 2\nu_2$ , $2\nu_1$ ) of $\text{CF}_4$ in the 1600 – 1800 $\text{cm}^{-1}$ region

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## ABSTRACT

The Fourier transform spectrum of  $\text{CF}_4$  in the 1600–1800  $\text{cm}^{-1}$  was recorded in Reims by using a White-type multi-pass cell to provide a path length of 8.262 m. In the present work, all spectrum analyses and fits were realized using the MIRS software based on tetrahedral tensorial formalism. By combining non-empirical contact transformation Hamiltonians for line positions and *ab initio* ro-vibrational normal-mode predictions for line intensities, we are able to achieve a simultaneous fit of effective Hamiltonian and dipole moment parameters of several cold and hot bands of  $\text{CF}_4$ . Hamiltonian operator was expanded up to the sixth order for the ground state and for the  $\{\nu_2 + 2\nu_4, \nu_2 + \nu_3, 4\nu_2, \nu_1 + 2\nu_2, 2\nu_1\}$  pentad. 1831 line positions were fitted to RMS of  $1.5 \times 10^{-3} \text{ cm}^{-1}$ . The standard deviation for line intensities for the cold bands  $\{\nu_2 + 2\nu_4, \nu_2 + \nu_3, 4\nu_2, \nu_1 + 2\nu_2, 2\nu_1\}$  is of 1.1% and of 8% and 5% for the hot band transitions  $\{2\nu_2 + 2\nu_4 - \nu_2, 2\nu_2 + \nu_3 - \nu_2\}$  and  $\{\nu_2 + 3\nu_4 - \nu_4, \nu_2 + \nu_3 + \nu_4 - \nu_4\}$ , respectively.

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## 1. Introduction

Tetrafluoromethane  $\text{CF}_4$  appears to be of vast industrial, environmental and economical interest. For these reasons, the ro-vibrational spectroscopy of this molecule was a subject of several theoretical and experimental studies.  $\text{CF}_4$  belongs to the group of perfluorocarbons (PFCs) which designates the chemicals composed of carbon and fluorine only. These species predominantly  $\text{CF}_4$  are extremely stable, inoffensive for the stratospheric ozone layer but powerful greenhouse gases (GHGs). This molecule is found among the longest-lived atmospheric trace gases because of its great chemical stability with an estimated lifetime of more than 50,000 years [1,2]. In addition, it seems to be the most abundant PFC in the stratosphere. Infrared high spectral resolution solar occultation spectrometer lead to an abundance of  $70.45 \pm 3.40$  ppt [3]. Because of this atmospheric persistence,  $\text{CF}_4$  is regulated by the Kyoto Protocol of the United Nations Framework Convention on Climate Change (UNFCCC) [4–6] as a substance whose atmospheric concentration should be supervised and reduced.

Besides its natural origin [7–9], the second source of  $\text{CF}_4$  is linked to aluminium production and to the manufacturing of semiconductor [10]. The spectroscopic databases like HITRAN [11,12] or GEISA [13,14] only contain a limited range of data for  $\text{CF}_4$ , without intensity analysis. The spectroscopy of  $\text{CF}_4$  which is essential for quantitative measurement in the atmosphere still remains relatively poorly studied because of the complexity of analyses. Some detailed high-resolution studies exist, however they mostly concern line positions. This was the case of the analysis of the  $\nu_3$  fundamental with the  $2\nu_3$  overtone [15] and the interaction of this stretching state with the  $2\nu_4$  overtone [16]. Recent work was already dedicated to the analysis of some rovibrational bands of this molecule including some hot bands around 1283  $\text{cm}^{-1}$  [17], the data being included in the Dijon database [18]. First accurate and complete *ab initio* based line lists for  $\text{CF}_4$  up to 400 K in the range 0–4000  $\text{cm}^{-1}$  have been recently published by the Reims–Tomsk groups [19], resulting in 2 billion rovibrational transitions. In Ref. [19], although the variationally-predicted were not computed at the spectroscopic accuracy, the error on line intensities was only of few % allowing a correct global description of the overtones and combination bands up to 6 vibrational quanta validated by direct comparisons with experimental PNNL spectra.

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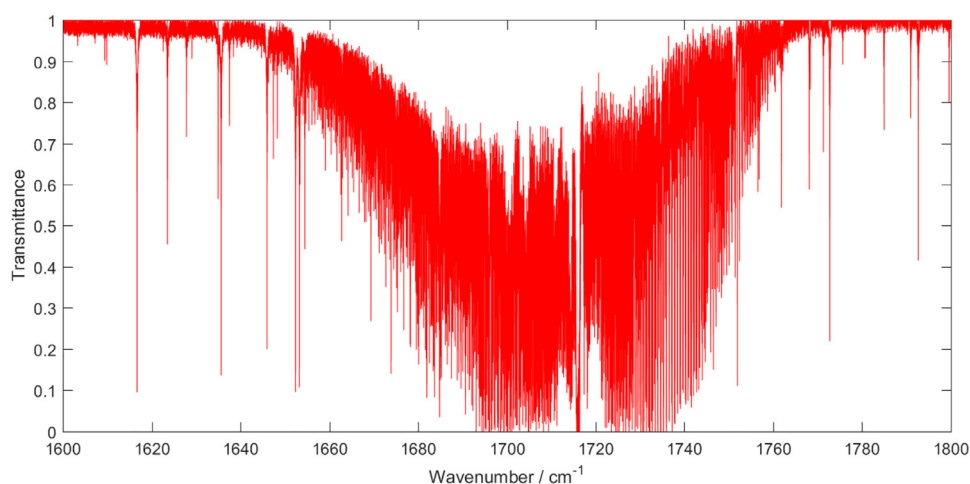


Fig. 1. Transmittance spectrum of  $\text{CF}_4$ , recorded at 294 K at a pressure of 9.93 Torr in the spectral range between 1600 and 1800  $\text{cm}^{-1}$ .

In this work, we report the analysis of rovibrational interacting pentad bands ( $\nu_2 + 2\nu_4$ ,  $\nu_2 + \nu_3$ ,  $4\nu_2$ ,  $\nu_1 + 2\nu_2$ ,  $2\nu_1$ ) of this molecule, in order to obtain a reliable simulation of  $\text{CF}_4$  atmospheric absorption in the region between 1600–1800  $\text{cm}^{-1}$ , including the most significant hot bands. The study of this system of bands requires the investigation of various vibrational fundamentals, combinations and harmonics bands. For the first time, high resolution analysis was made by combining spectroscopic effective models and information provided by accurate *ab initio* potential energy (PES) and dipole moment (DMS) surfaces both for line positions and line intensities. A non-empirical effective Hamiltonian was built from a PES by applying high-order contact-transformations (CT) [20,21]. For the line intensities, some effective dipole moment parameters have been adjusted directly on the *ab initio* intensities [19]. Section 2 describes the experimental data used, while Section 3 reviews some basic principles concerning the theoretical model for tetrahedral molecules used for our analysis and presents the obtained results.

## 2. Experimental details

Spectra measurements were carried out with the Fourier Transform Spectrometer (FTS) of GSMA laboratory [22]. The  $\text{CF}_4$  spectrum was recorded at room temperature over the range 1600–1800  $\text{cm}^{-1}$ , with a spectral resolution of 0.003  $\text{cm}^{-1}$  (maximum optical path length of 2.99 m). This experimental setup was equipped with a globar source, a  $\text{CaF}_2$  beam splitter and a  $\text{HgCdTe}$  detector. A White-type multi-pass cell with a basis length of 2 m was used to provide a path length of 8.262 m. The pressure in the cell was maintained at 9.93 Torr. The spectrum was obtained by co-adding and averaging 289 one-sided interferometer scans in order to achieve a sufficient signal-to-noise ratio. The diaphragm diameter was set to 4.5 mm. Wavenumber calibration was carried out using residual of ten water lines with intensities in the range  $2 \times 10^{-19}$  –  $2 \times 10^{-20}$  cm/molecule.

The detailed experimental conditions are listed in Table 1, and an overview of transmittance spectrum of the ( $\nu_2 + 2\nu_4$ ,  $\nu_2 + \nu_3$ ,  $4\nu_2$ ,  $\nu_1 + 2\nu_2$ ,  $2\nu_1$ ) pentad is plotted in Fig. 1.

## 3. Spectra analysis and modeling

### 3.1. Theoretical methods

The aim of the present work is the detailed analysis of the pentad system composed of five bands, namely  $\nu_2 + 2\nu_4$ ,  $\nu_2 + \nu_3$ ,  $4\nu_2$ ,  $\nu_1 + 2\nu_2$  and  $2\nu_1$  in the spectral range 1600–1800  $\text{cm}^{-1}$  using

Table 1

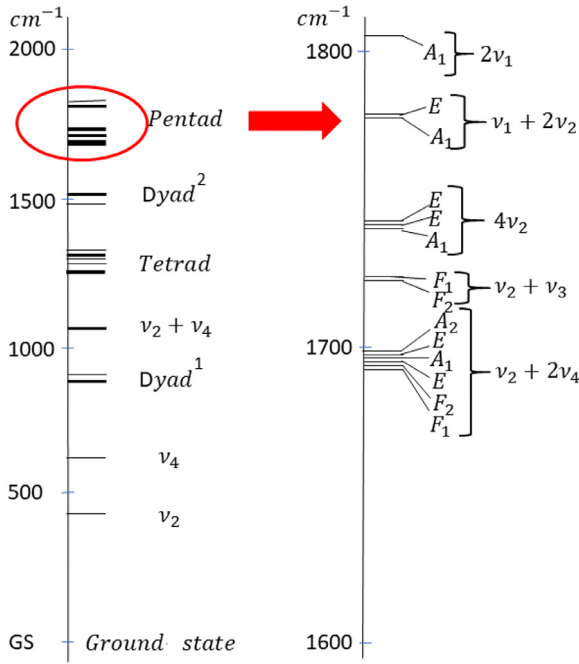
Experimental details.

| Spectral range ( $\text{cm}^{-1}$ ) | 1600–1800        |
|-------------------------------------|------------------|
| Resolution ( $\text{cm}^{-1}$ )     | 0.003            |
| Source                              | Globar           |
| Separator                           | $\text{CaF}_2$   |
| Detector                            | $\text{HgCdTe}$  |
| Cell                                | White multi-pass |
| Optical Path (m)                    | 8.262            |
| Temperature (K)                     | 294              |
| Pressure (Torr)                     | 9.93             |
| Aperture diameter (mm)              | 4.5              |

a specific spectroscopic model (see also section 4.2). The absorption in this range is dominated by  $\nu_2 + \nu_3$  with the band center near 1717  $\text{cm}^{-1}$ . The complexity of  $\text{CF}_4$  vibrational states is outlined in Fig. 2 where the scale on the right panel represents sub-level with their symmetry type of the pentad split into 14 different components. The  $\nu_2 + 2\nu_4$  vibrational band has 6 vibrational sub-levels of  $A_1$ ,  $A_2$ ,  $2E$ ,  $F_1$  and  $F_2$  symmetry types, the  $\nu_2 + \nu_3$  band has 2 sublevels ( $F_1$ ,  $F_2$ ),  $4\nu_2$  has 3 sublevels of symmetry  $A_1$  and  $2E$ ,  $\nu_1 + 2\nu_2$  has 2 sublevels ( $A_1$ ,  $E$ ) and  $2\nu_1$  has one upper state vibrational level of  $A_1$  symmetry.

The infrared tetrafluoromethane spectra are known to be very complex due to the high tetrahedral symmetry of  $\text{CF}_4$  which causes degeneracies and quasi-degeneracies of the vibration modes and large and complicated resonance interactions because of the inter-modes coupling. Note that such a line-by-line analysis from a judicious combination of effective and *ab initio* models – both for line positions and line intensities – was not yet available in this spectral interval. In order to extend rovibrational analyses of experimental spectra, we used a tensorial formalism with vibrational extrapolation [23,24] including supplementary information derived from *ab initio* PES/DMS and first-principles calculations. These *ab initio* variational line lists are available in the TheoReTS information system [25]. Significant recent advance in *ab initio* PES, DMS and in rotationally resolved variational spectra calculations reported by Rey et al. [19] permitted identification of the most important bands in this range. But, many weak transitions remained not yet assigned.

The rovibrational model for line-by-line analysis is built by admitting that the vibrational energy levels of  $\text{CF}_4$  can be gathered in a series of polyads. In the framework of this theory, an effective Hamiltonian (EH) for all the considered polyads was constructed from high-order contact transformations (CT) to eliminate inter-



**Fig. 2.** Scheme of vibrational level patterns of the  $\text{CF}_4$  polyads (left side) and vibrational sublevels of the part of the pentad system corresponding to ro-vibrational bands analyzed in this work. At the right hand side panel, symmetry type ( $T_d$  irreducible representations) of vibration sublevels, vibrational ranking numbers within the pentad and the corresponding vibrational bands are given. The first dyad is composed of  $[2v_4, v_1]$ , the tetrad is composed of  $[2v_4, v_3, 3v_2, v_1 + v_4]$  and the second dyad is composed of  $[v_1 + v_2, 2v_2 + v_4]$ .

polyad interaction terms in the rovibrational Hamiltonian (transform completely the Hamiltonian to a block-diagonal form) [20,21]. This approach allows a fairly precise characterization of resonance coupling terms to avoid the ambiguity of purely empirical EH models [26,27]. The effective Hamiltonian operator  $\tilde{H}$  adapted to the polyad structure of the  $\text{CF}_4$  is constructed as the sum of contributions appropriate to the various polyads designed by  $P_k$  ( $k=0, 1, \dots, k, \dots$ ),  $P_0$  being the ground state (GS). This effective Hamiltonian is expressed as (see Fig. 2)

$$\begin{aligned} \tilde{H} &= \sum \tilde{H}_{\{P_k\}} = \tilde{H}_{\{P_0\}} + \tilde{H}_{\{P_1\}} + \dots + \tilde{H}_{\{P_k\}} + \dots \\ &= \tilde{H}_{\{P_0\}} + \tilde{H}_{\{v_2\}} + \tilde{H}_{\{v_4\}} + \tilde{H}_{\{\text{Dyad1}\}} + \tilde{H}_{\{v_2+v_4\}} + \tilde{H}_{\{\text{Tetrad}\}} \\ &\quad + \tilde{H}_{\{\text{Dyad2}\}} + \tilde{H}_{\{\text{Pentad}\}} \end{aligned} \quad (1)$$

The terms of this expansion gather operators specific to the successive polyads in the representation of the irreducible tensor operators (ITO) [24,28,29]. They are written as a linear combination of rovibrational operators  $T_{\{n_s\}\{m_s\}}^{\Omega(K, n\Gamma_r)\Gamma_1\Gamma_2\Gamma_v}$  as

$$\tilde{H}_{\{P_k\}} = \sum_{\text{all indices}} t_{\{n_s\}\{m_s\}}^{\Omega(K, n\Gamma_r)\Gamma_1\Gamma_2\Gamma_v} T_{\{n_s\}\{m_s\}}^{\Omega(K, n\Gamma_r)\Gamma_1\Gamma_2\Gamma_v}. \quad (2)$$

In this equation, the  $t$ 's are the corresponding adjustable parameters to be determined. All the indices represent the intermediate quantum numbers and symmetries. Here  $T$  refers to a rovibrational operator expressed as a tensor product between rotational and vibrational operators as

$$T_{\{n_s\}\{m_s\}}^{\Omega(K, n\Gamma_r)\Gamma_1\Gamma_2\Gamma_v} = \beta \left( R^{\Omega(K, n\Gamma_r)} \otimes \varepsilon V_{\{n_s\}\{m_s\}}^{\Gamma_1\Gamma_2\Gamma_v} \right)^{(A_1)}. \quad (3)$$

Rotational operators  $R^{\Omega(K, n\Gamma_r)}$  are the usual symmetry-adapted irreducible tensor operators (ITO) for spherical top molecules [23,24,30–34]. The upper index  $\Omega$  is the rotational degree in the molecular frame components  $J_x, J_y$  and  $J_z$ ;  $K$  is the

tensor rank relative to the full rotation group  $O(3)$ ,  $n$  is a multiplicity index and  $\Gamma_r$  ( $=A_1, A_2, E, F_1$  or  $F_2$ ) is the rotational symmetry in the  $T_d$  point group. The vibrational ITOs are constructed by recursive coupling of elementary creation  $a^+$  and annihilation  $a$  operators for each of the four normal modes.  $\varepsilon$  is the parity in conjugate momenta with the condition  $\varepsilon = (-1)^\Omega$  and  $\Gamma_1$  and  $\Gamma_2$  are the symmetries of vibrational operators;  $n_s$  and  $m_s$  are the degrees in the  $a^+$  and  $a$  operators. Here,  $\beta$  is a numerical factor, equal to  $[\Gamma_1]^{1/2}(-\sqrt{3}/4)^{\Omega/2}$  if  $(K, n\Gamma_r) = (0, 0A_1)$ , and 1 otherwise. Since the Hamiltonian is totally symmetric, we have necessary  $\Gamma_r = \Gamma_v$ . The order of the individual terms in  $T$  is defined as  $\sum_s (n_s + m_s) + \Omega - 2$ . Orientation of the ITOs in the chain  $O(3) \supset T_d$  has been described in details in Ref. [35].

In order to perform intensity calculation, the dipole moment operator is built in the same way as the effective Hamiltonian operator. Its components  $\mu_\Theta$  ( $\Theta = X, Y$  or  $Z$ ) in the laboratory fixed frame (LFF) are linked to the components  $\mu_\theta$  ( $\theta = x, y$  or  $z$ ) in the molecule fixed frame (MFF) as

$$\mu_\Theta = \sum_\theta \lambda_{\Theta\theta} \mu_\theta, \quad (4)$$

where  $\lambda_{\Theta\theta}$  are the direction cosines. The transformed LFF components can be expressed in terms of the MFF transformed components as

$$\tilde{\mu}_\Theta = \frac{1}{2} \sum_\theta (\lambda_{\Theta\theta} \tilde{\mu}_\theta + \tilde{\mu}_\theta \lambda_{\Theta\theta}), \quad (5)$$

where the MFF effective dipole moment is of symmetry  $F_2$  with its three components  $\tilde{\mu}_\theta$  expanded as a series of rovibrational operators [23,36]

$$\tilde{\mu}_\theta^{(F_2)} = \sum_{\{i\}} \tilde{\mu}^{(i), \Gamma} \left( R^{\Omega(K, n\Gamma_r)} \otimes \varepsilon V_{\{n_s\}\{m_s\}}^{\Gamma_1\Gamma_2\Gamma_v} \right)^{(F_2)}. \quad (6)$$

Here  $\tilde{\mu}^{(i), \Gamma}$  are the effective dipole moment parameters to be determined. A way to determine these parameters will be described in the Section 3.2.2. Note that the order of the individual terms of the dipole moment is  $\sum_s (n_s + m_s) + \Omega - 1$ .

### 3.2. Analysis and discussion

#### 3.2.1. Line positions

Line-by-line analysis and assignments of experimental spectra are known to be tedious and difficult tasks. They are usually based on empirical models for the polyads of close states where line positions are generated from empirically-fitted effective Hamiltonians (EH). Using a vibrational extrapolation scheme, we simulate the spectrum by considering the polyad  $P_0$ , which contains the GS level and the polyad  $P_7$  containing the  $v_2+2v_4$ ,  $v_2 + v_3$ ,  $4v_2$ ,  $v_1+2v_2$  and  $2v_1$  levels as the upper polyad. To obtain a result with a precision comparable with the experimental accuracy, the effective Hamiltonian was expanded up to the sixth order. As a starting point, we needed a first estimation of the tensorial parameters. That permitted a correct description of the experimental spectrum at a given temperature. Starting with the best possible set of parameters permits to identify a large number of observed line positions. Here as a first guess, the *ab initio* based EH parameters obtained via the CT method were used and further refined after various iterations by increasing the number of assigned lines. This supplied realistic estimations for the rovibrational coupling parameters between various bands of a given polyad that are responsible for the intensity perturbations due to accidental resonances. This first set of parameters allowed a preliminary simulation. In what follows, parameters for  $v_2$ ,  $v_4$ ,  $(2v_2, v_1)$  dyad,  $v_2 + v_4$ ,  $(2v_4, v_3, 3v_2, v_1)$  tetrad and  $(2v_2 + v_4, v_1 + v_4)$  dyad states will be fixed to the “*ab initio*” values derived from CTs. That means that only

the parameters for the ground state and ( $\nu_2 + \nu_4$ ,  $\nu_2 + \nu_3$ ,  $4\nu_2$ ,  $\nu_1 + 2\nu_2$ ,  $2\nu_1$ ) are adjusted. Spectrum simulations, line assignments and the fitting process were carried out using the MIRS [37,38] and SpectrAssign softwares based on an iterative semi-automatized comparison of synthetic and experimental spectra. The  $T_d$  formalism is implemented in MIRS that thus provides all vibrational and rotational quantum numbers. The SpectrAssign program was used for assignments and visualization. It allows a simultaneous display of predicted lines from MIRS output and of an experimental spectrum. A limited number of observed lines were assigned by comparing the positions and intensities of the observed and simulated lines calculated from the initial values of the spectroscopic parameters. Afterwards, new assignments were made leading to a new set of fitted tensorial parameters. This process was performed iteratively until no more lines could be assigned.

Finally, among the 618 parameters included in the sixth order effective Hamiltonian, only 48 have been fitted. The root mean square (RMS) deviation of the IR data is very satisfactory. In Fig. 3, we present an overview of the pentad region. It compares the observed spectrum with the calculated one at the same experimental conditions by using our final set of parameters in the region 1600–1800  $\text{cm}^{-1}$  dominated by the  $\nu_2 + \nu_3$  combination band. As shown in this figure, calculated and observed spectra are in very good agreement. As can be seen in Fig. 4, different zooms reveal the precision obtained of calculated positions in the R and P-branches though weaker lines belonging to hot bands need to be still improved. Note that the Q-branch has an unresolved structure due to a tremendous number of overlapping lines making assignments very complicated. In order to keep the total number of adjusted parameters reasonable and to make the model “stable” when slightly changing parameters, an important part of the non-diagonal resonance parameters have been kept fixed to their theoretical values generated from the PES [19] in a manner similar to that described for the methane molecule [23]. Although the *ab initio* values for non-diagonal parameters are approximations, they describe the resonance couplings in a quite qualitatively correct way. On the other hand, an empirical “blind” fitting of a big set of effective polyad Hamiltonian parameters to rovibrational energies is known as a mathematically poorly defined inverse problem [26,27]. Purely empirical approach could improve a root-mean-square deviation on line positions, but it often results in

wrong values of interaction parameters and thus wrong intensity extrapolations. The accurate theoretical calculations of resonance coupling parameters from *ab initio* PES and of the corresponding intensities are difficult tasks.

According to this preliminary study, we were able to assign 1831 line positions, belonging mainly to the  $\nu_2 + \nu_3$  combination band, with a root mean square deviation of  $1.590 \times 10^{-3} \text{ cm}^{-1}$ . Nevertheless, a large part of the transitions remained unassigned. A complete modelling of all bands in this region will require more assignments of both hot and cold transitions. Number of assigned transitions per band depends on the inherent intensity of the band and on the degree of interaction with other states. Fig. 5 shows the residual defined by the difference between the observed and calculated wavenumbers versus  $J$ . It gives a global idea of the quality of the fit. We note that the data are well distributed on the scale of the  $J$  values. It allowed to make sure that the rovibrational energy levels are well sampled and that the results of the fit are reliable. Fig. 6 presents the resulting calculated reduced upper state energies vs  $J$  showing the mixtures between the vibrational sub-levels.

### 3.2.2. Line intensities

Tetrafluoromethane has the same tetrahedral geometry as methane but the main difference is that the  $\text{CF}_4$  molecule is heavier and consequently has much lower vibrational modes and much higher  $J$ -levels are populated. This leads to a considerable number of IR-active transitions belonging to hot bands that make the experimental spectra more complicated to analyse. Rey et al. [19] showed that the  $\text{CF}_4$  spectra could be described from (i) a set of strong lines responsible for the strongest absorption features and (ii) of a kind of quasi-continuum formed by an accumulation of very weak overlapping lines (see figure of [19]). This situation is quite similar to that observed in hot methane spectra at  $\approx 1000\text{--}1200 \text{ K}$  [39]. Because the infrared spectra of  $\text{CF}_4$  are very congested, no single line intensity measurements were made. Almost all the observed lines result from overlapping of several transitions. To overcome such a difficulty, we have adopted a strategy quite different from “usual” spectroscopic studies. It consists in fitting the effective dipole moment parameters directly on the variationally-determined *ab initio* intensities. In other words, the variational line list built in Ref. [19] was taken as an “observed” line list for the present work. To our knowledge, this procedure has never been applied before for such a complex

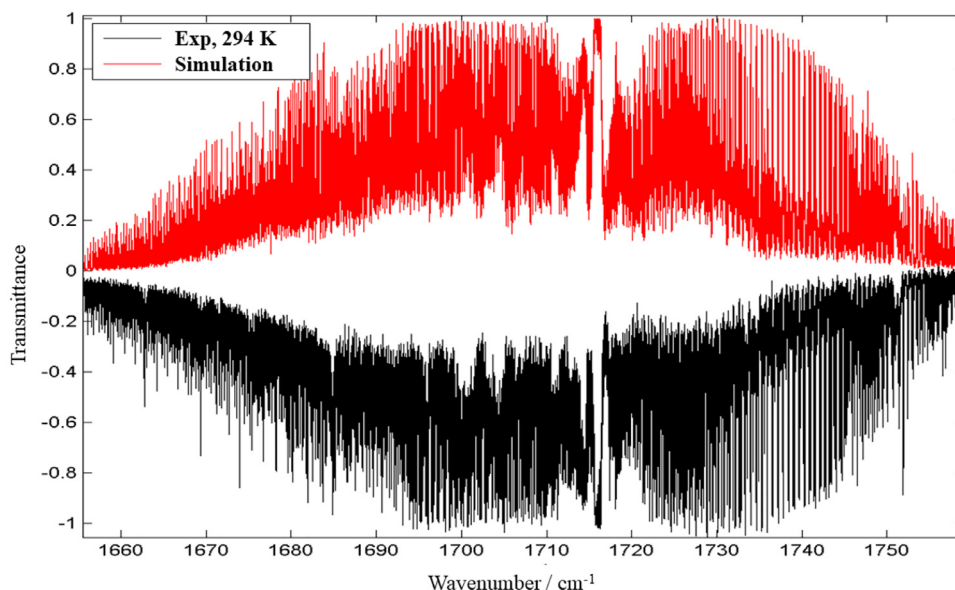
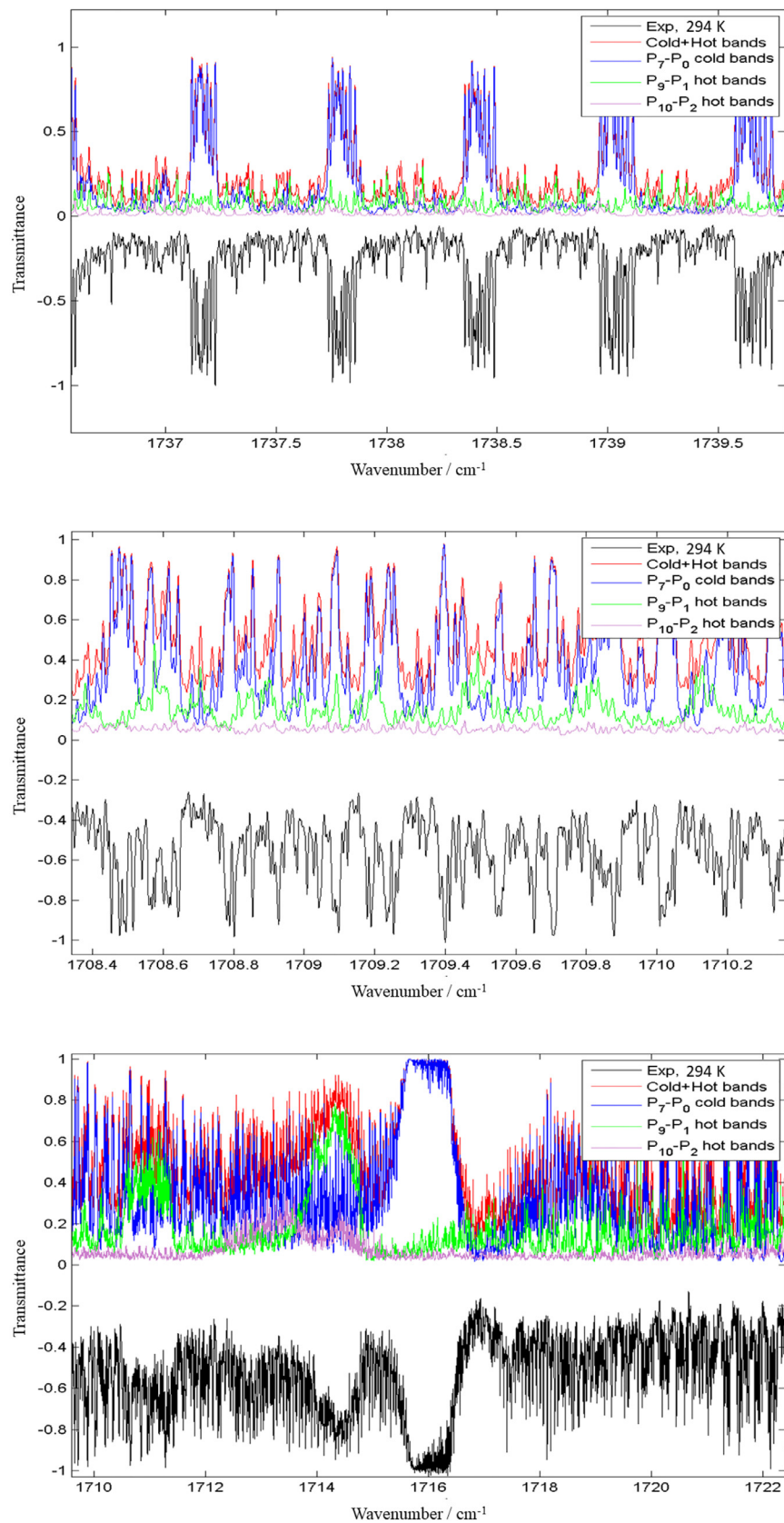


Fig. 3. Comparison between experimental and simulated spectra of  $\text{CF}_4$  at 294 K.

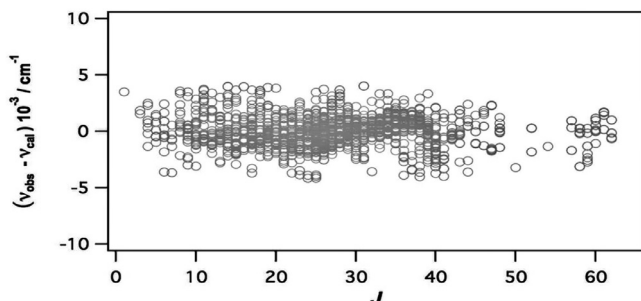


**Fig. 4.** Detailed comparison of the R, P and Q-branches between the observed spectrum and the simulation.

**Table 2**

Summary of the intensity fit statistics of hot and cold bands.

| Bands      |                                                                 | $J_{\max}$ | Parameters | Assignments | St. Deviation (%) |
|------------|-----------------------------------------------------------------|------------|------------|-------------|-------------------|
| Hot Bands  | $\{2\nu_2+2\nu_4 - \nu_2, 2\nu_2 + \nu_3 - \nu_2\}$             | 35         | 28         | 1871        | 8                 |
|            | $\{\nu_2+3\nu_4 - \nu_4, \nu_2 + \nu_3 + \nu_4 - \nu_4\}$       | 30         | 34         | 21,779      | 5                 |
| Cold Bands | $\{\nu_2+2\nu_4, \nu_2 + \nu_3, 4\nu_2, \nu_1+2\nu_2, 2\nu_1\}$ | 70         | 16         | 36,896      | 1.1               |



**Fig. 5.** Observed minus calculated wavenumbers versus  $J$  for the pentad obtained using the tensorial model. We have a lack of assignments between  $J = 50$  and  $J = 60$  because of the complexity of the  $\text{CF}_4$  spectrum in this region.

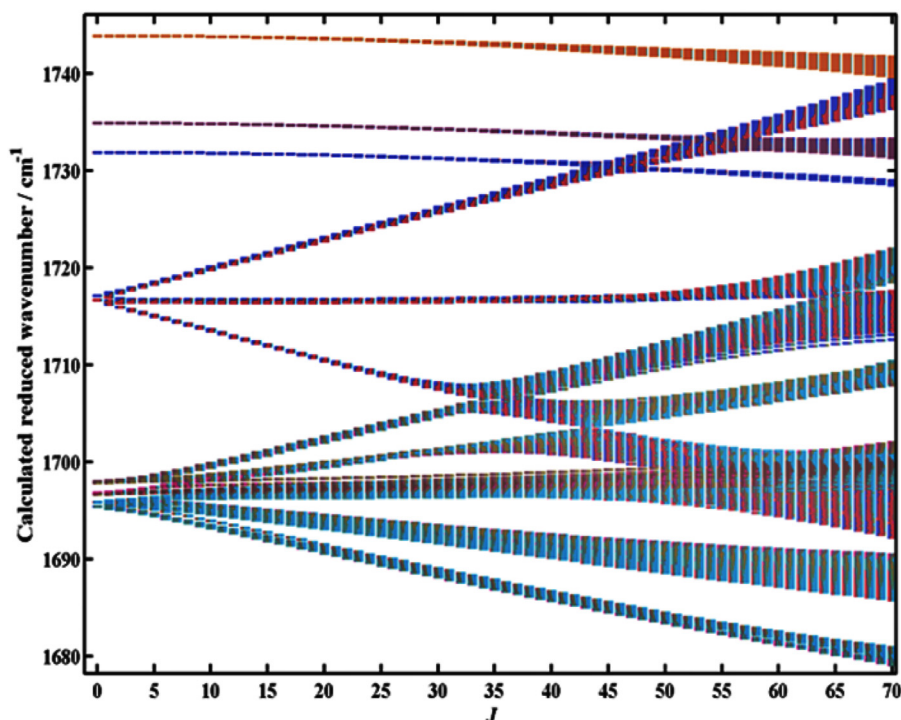
five-atomic molecular system. The improvement of hot bands intensity modelling will be necessary to achieve a more detailed assignment in the region between  $1600$  and  $1800\text{ cm}^{-1}$  because many other bands belonging to upper polyads e.g.  $\{2\nu_2+2\nu_4 - \nu_2, 2\nu_2 + \nu_3 - \nu_2, \nu_2+3\nu_4 - \nu_4, \nu_2 + \nu_3 + \nu_4 - \nu_4\}$  contribute to the absorption. Detailed statistics on this fit are listed in Table 2. The obtained effective dipole moment parameters for each band summarized in Tables 3–5 rely on tensorial formalism notations. Fig. 7 presents the resulting comparison between experiment and simulation, with a nice agreement.

**Table 3**Effective dipole moment parameters for the pentad  $\{\nu_2+2\nu_4, \nu_2 + \nu_3, 4\nu_2, \nu_1+2\nu_2, 2\nu_1\}$ .

|    | Tensorial nomenclature   |             | Parameter value   |
|----|--------------------------|-------------|-------------------|
|    | Rotational               | Vibrational |                   |
| 1  | R 0(0, 0A <sub>1</sub> ) | 0000 0102   | −3.1864538041E−04 |
| 2  | R 0(0, 0A <sub>1</sub> ) | 0000 0110   | 9.2315030571E−03  |
| 3  | R 1(1, 0F <sub>1</sub> ) | 0000 0102   | 1.2645029053E−06  |
| 4  | R 1(1, 0F <sub>1</sub> ) | 0000 0102   | −1.5390407345E−06 |
| 5  | R 1(1, 0F <sub>1</sub> ) | 0000 0102   | −1.0548647219E−06 |
| 6  | R 1(1, 0F <sub>1</sub> ) | 0000 0102   | 1.2126152550E−06  |
| 7  | R 1(1, 0F <sub>1</sub> ) | 0000 0102   | 1.5140232527E−06  |
| 8  | R 1(1, 0F <sub>1</sub> ) | 0000 0110   | −1.2529078490E−05 |
| 9  | R 1(1, 0F <sub>1</sub> ) | 0000 0110   | 2.2076230400E−05  |
| 10 | R 1(1, 0F <sub>1</sub> ) | 0000 1200   | 2.2577731112E−05  |
| 11 | R 2(2, 0E)               | 0000 0110   | 1.5222082012E−08  |
| 12 | R 2(2, 0F <sub>2</sub> ) | 0000 0110   | −1.6089425075E−08 |
| 13 | R 2(0, 0A <sub>1</sub> ) | 0000 0110   | −1.5434076560E−08 |
| 14 | R 2(2, 0E)               | 0000 0110   | −2.6796194562E−09 |
| 15 | R 2(2, 0F <sub>2</sub> ) | 0000 0110   | −4.8524933639E−09 |
| 16 | R 2(2, 0F <sub>2</sub> ) | 0000 2000   | −2.7018916009E−06 |

#### 4. Conclusion

In this work, we have built a tensorial model that permitted for the first time a detailed analysis of the  $(\nu_2+2\nu_4, \nu_2 + \nu_3, 4\nu_2, \nu_1+2\nu_2, 2\nu_1)$  band system of  $\text{CF}_4$  by combining non-empirical CT Hamiltonian, *ab initio* based variational predictions and experimental information. The present analysis supplies a first accurate and



**Fig. 6.** Calculated reduced rovibrational energy levels of  $\text{CF}_4$  transitions from  $1600$  to  $1800\text{ cm}^{-1}$  up to  $J = 70$ . Various colors correspond to mixing coefficients of normal mode wavefunctions due to resonance interactions.

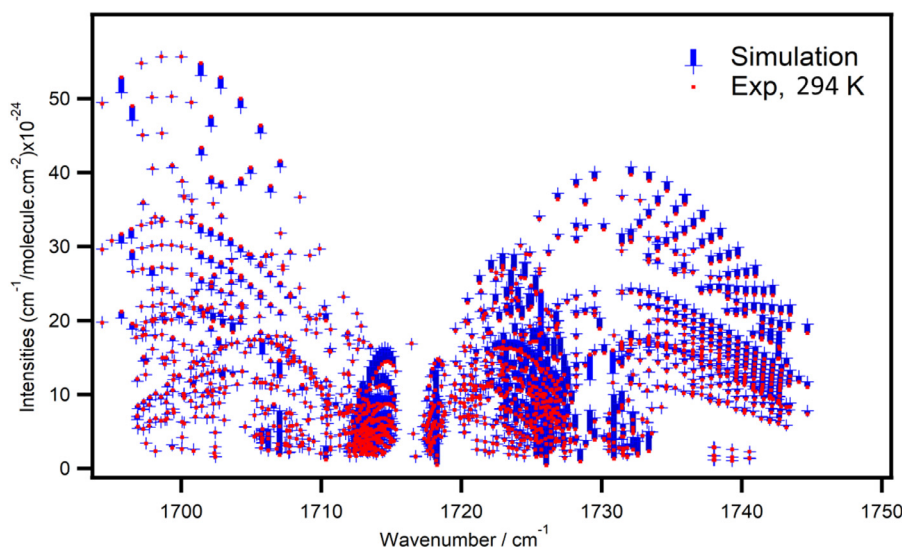


Fig. 7. Observed and calculated intensities (in  $\text{cm}^{-1}/\text{molecule.cm}^{-2}$ ) vs line positions (in  $\text{cm}^{-1}$ ) for the pentad transitions.

Table 4

Effective dipole moment parameters for the  $\{2\nu_2+2\nu_4 - \nu_2, 2\nu_2 + \nu_3 - \nu_2\}$  bands.

|    | Tensorial nomenclature   |             | Parameter value   |
|----|--------------------------|-------------|-------------------|
|    | Rotational               | Vibrational |                   |
| 1  | R 0(0, 0A <sub>1</sub> ) | 0100 0202   | 3.9825129095E-04  |
| 2  | R 0(0, 0A <sub>1</sub> ) | 0100 0210   | 8.5008049280E-03  |
| 3  | R 0(0, 0A <sub>1</sub> ) | 0100 0202   | 6.5544003405E-06  |
| 4  | R 0(0, 0A <sub>1</sub> ) | 0100 0210   | 7.9030718238E-05  |
| 5  | R 0(0, 0A <sub>1</sub> ) | 0100 0202   | -4.4623481744E-04 |
| 6  | R 0(0, 0A <sub>1</sub> ) | 0100 0210   | 1.2930872868E-02  |
| 7  | R 0(0, 0A <sub>1</sub> ) | 0100 1002   | -4.6917331225E-03 |
| 8  | R 0(0, 0A <sub>1</sub> ) | 0100 1010   | 2.4241217720E-03  |
| 9  | R 1(1, 0F <sub>1</sub> ) | 0100 0210   | -1.4446087275E-05 |
| 10 | R 1(1, 0F <sub>1</sub> ) | 0100 0210   | 2.0583793143E-05  |
| 11 | R 1(1, 0F <sub>1</sub> ) | 0100 0210   | -2.4295434890E-08 |
| 12 | R 1(1, 0F <sub>1</sub> ) | 0100 0210   | 6.1335719221E-07  |
| 13 | R 1(1, 0F <sub>1</sub> ) | 0100 0210   | -1.8777757468E-05 |
| 14 | R 1(1, 0F <sub>1</sub> ) | 0100 0210   | 3.0644918638E-05  |
| 15 | R 1(1, 0F <sub>1</sub> ) | 0100 1002   | 3.2123332254E-05  |
| 16 | R 1(1, 0F <sub>1</sub> ) | 0100 1002   | 3.4243832575E-05  |
| 17 | R 1(1, 0F <sub>1</sub> ) | 0100 1002   | -5.2783161222E-05 |
| 18 | R 1(1, 0F <sub>1</sub> ) | 0100 1002   | -2.0547145593E-05 |
| 19 | R 1(1, 0F <sub>1</sub> ) | 0100 1002   | 3.2786033893E-06  |
| 20 | R 1(1, 0F <sub>1</sub> ) | 0100 1010   | 4.8417050701E-04  |
| 21 | R 1(1, 0F <sub>1</sub> ) | 0100 1010   | 2.4135565610E-04  |
| 22 | R 1(1, 0F <sub>1</sub> ) | 0100 2100   | 6.9831133618E-05  |
| 23 | R 1(1, 0F <sub>1</sub> ) | 0100 2100   | 2.7944989174E-04  |
| 24 | R 2(2, 0E)               | 0100 1010   | 1.1388797951E-07  |
| 25 | R 2(2, 0F <sub>2</sub> ) | 0100 1010   | 1.6758890863E-07  |
| 26 | R 2(0, 0A <sub>1</sub> ) | 0100 1010   | -3.1635728036E-07 |
| 27 | R 2(2, 0E)               | 0100 1010   | 3.0201035785E-07  |
| 28 | R 2(2, 0F <sub>2</sub> ) | 0100 1010   | -6.7979486512E-09 |

Table 5

Effective dipole moment parameters for the  $\{\nu_2+3\nu_4 - \nu_4, \nu_2 + \nu_3 + \nu_4 - \nu_4\}$  bands.

|    | Tensorial nomenclature   |             | Parameter value   |
|----|--------------------------|-------------|-------------------|
|    | Rotational               | Vibrational |                   |
| 1  | R 0(0, 0A <sub>1</sub> ) | 0001 0103   | 4.2413135847E-04  |
| 2  | R 0(0, 0A <sub>1</sub> ) | 0001 0103   | 2.7477468633E-05  |
| 3  | R 0(0, 0A <sub>1</sub> ) | 0001 0103   | 5.0088447964E-04  |
| 4  | R 0(0, 0A <sub>1</sub> ) | 0001 0103   | -1.2620967754E-05 |
| 5  | R 0(0, 0A <sub>1</sub> ) | 0001 0103   | 4.7003806247E-04  |
| 6  | R 0(0, 0A <sub>1</sub> ) | 0001 0103   | 6.4939084452E-05  |
| 7  | R 0(0, 0A <sub>1</sub> ) | 0001 0103   | 1.7797635608E-06  |
| 8  | R 0(0, 0A <sub>1</sub> ) | 0001 0111   | -1.5473534628E-02 |
| 9  | R 0(0, 0A <sub>1</sub> ) | 0001 0111   | -3.0138555124E-04 |
| 10 | R 0(0, 0A <sub>1</sub> ) | 0001 0111   | -1.6398372407E-05 |
| 11 | R 0(0, 0A <sub>1</sub> ) | 0001 0111   | 2.5471384745E-04  |
| 12 | R 0(0, 0A <sub>1</sub> ) | 0001 0111   | 7.8534831544E-05  |
| 13 | R 0(0, 0A <sub>1</sub> ) | 0001 0111   | -2.0691429033E-05 |
| 14 | R 0(0, 0A <sub>1</sub> ) | 0001 0111   | -1.9854225634E-04 |
| 15 | R 1(1, 0F <sub>1</sub> ) | 0001 0111   | 2.2599131901E-05  |
| 16 | R 1(1, 0F <sub>1</sub> ) | 0001 0111   | -3.8669830582E-05 |
| 17 | R 1(1, 0F <sub>1</sub> ) | 0001 0111   | 9.1638928149E-07  |
| 18 | R 1(1, 0F <sub>1</sub> ) | 0001 0111   | -6.7549134470E-07 |
| 19 | R 1(1, 0F <sub>1</sub> ) | 0001 0111   | 2.9523535789E-07  |
| 20 | R 1(1, 0F <sub>1</sub> ) | 0001 0111   | 9.7740749402E-07  |
| 21 | R 1(1, 0F <sub>1</sub> ) | 0001 0111   | -5.7925405576E-07 |
| 22 | R 1(1, 0F <sub>1</sub> ) | 0001 0111   | 3.8835519379E-07  |
| 23 | R 1(1, 0F <sub>1</sub> ) | 0001 0111   | 1.2693515619E-06  |
| 24 | R 1(1, 0F <sub>1</sub> ) | 0001 0111   | 1.1521893674E-06  |
| 25 | R 1(1, 0F <sub>1</sub> ) | 0001 0111   | -5.5755866620E-08 |
| 26 | R 1(1, 0F <sub>1</sub> ) | 0001 0111   | -4.1284363860E-07 |
| 27 | R 1(1, 0F <sub>1</sub> ) | 0001 0111   | -4.6669243132E-07 |
| 28 | R 1(1, 0F <sub>1</sub> ) | 0001 0111   | 4.8599003861E-07  |
| 29 | R 1(1, 0F <sub>1</sub> ) | 0001 0111   | 5.2493460965E-08  |
| 30 | R 1(1, 0F <sub>1</sub> ) | 0001 0111   | -6.2965529693E-07 |
| 31 | R 1(1, 0F <sub>1</sub> ) | 0001 0111   | -8.3742085981E-07 |
| 32 | R 1(1, 0F <sub>1</sub> ) | 0001 0111   | -7.7632490926E-07 |
| 33 | R 1(1, 0F <sub>1</sub> ) | 0001 0111   | 1.2083069804E-06  |
| 34 | R 1(1, 0F <sub>1</sub> ) | 0001 0111   | -6.9909391873E-07 |

complete description of the infrared high-resolution spectrum of CF<sub>4</sub> in the region 1600–1800  $\text{cm}^{-1}$ . The final set of effective Hamiltonian/dipole moment parameters allows to reproduce simultaneously (i) quite accurately the cold bands and (ii) in a qualitative manner the main hot band transitions. The global fit of parameters and calculation of infrared spectrum was carried out with the MIRS and SpectrAssign softwares. Assignments was made up to  $J_{\text{max}} = 70$  and a set of 48 tensorial parameters has been adjusted. The accuracy of the final empirically optimized calculated line list is satisfactory. We have obtained a root mean square deviation RMS of about  $1.590 \times 10^{-3} \text{ cm}^{-1}$  for 1831 cold transitions. The standard deviation for line intensities for the cold bands  $\{\nu_2+2\nu_4, \nu_2 + \nu_3,$

$4\nu_2, \nu_1+2\nu_2, 2\nu_1\}$  is of 1.1%, with respect to the *ab initio* intensities, while, for the hot bands  $\{2\nu_2+2\nu_4 - \nu_2, 2\nu_2 + \nu_3 - \nu_2\}$  and  $\{\nu_2+3\nu_4 - \nu_4, \nu_2 + \nu_3 + \nu_4 - \nu_4\}$ , the standard deviation is of  $\approx 8\%$  and  $5\%$ , respectively. The main objective of this study was to construct a CF<sub>4</sub> line list for various applications. This final list will include new assignments in the forthcoming versions of the spectroscopic HITRAN and GEISA databases. The line list for the assigned transitions is available as supplementary material to this paper. By using cold spectra ( $< 200 \text{ K}$ ), future works will consist

to both improve the present region and extend analysis to other spectral ranges on the basis of the present results.

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## Supplementary material

Supplementary material associated with this article can be found, in the online version, at doi:[10.1016/j.jqsrt.2019.01.018](https://doi.org/10.1016/j.jqsrt.2019.01.018).

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Potential Energy Surface of SF<sub>6</sub>

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## SPECTROSCOPY OF AMBIENT MEDIUM

# Potential Energy Surface of SF<sub>6</sub>

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**Abstract**—For the first time, a 15-dimensional analytical form was obtained and the potential energy of the SF<sub>6</sub> molecule in the ground electronic state was found. An optimal grid of geometries was constructed, which, taking into account the full symmetry of the molecule, unambiguously determines the potential energy surface of the sixth order. Using the MP2 method with the cc-pVTZ base set, the potential energy surface of the fourth order was calculated.

**Keywords:** potential energy surface, SF<sub>6</sub>, octahedral group

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## INTRODUCTION

Sulfur hexafluoride (or elegas) SF<sub>6</sub> has being investigated over forty years [1–3]. It is widely used in industry and medicine. Due to the strong absorption bands in the atmospheric transparency windows, SF<sub>6</sub> is a potent greenhouse gas, and its global warming potential (GWP500) is 29 300 times higher than that of CO<sub>2</sub> [4]. The atmospheric concentration of SF<sub>6</sub> is small, but it is one of the most stable gaseous chemicals with an atmospheric lifetime of 3200 years [5]. According to Earth System Research Laboratory Global Monitoring Division data [6], the SF<sub>6</sub> concentration increased from 3.5 ppt (1995) to 9.5 ppt (2018). This rapid increase and long atmospheric lifetime can result in considerable accumulation of SF<sub>6</sub> in the troposphere and its important role in global climate change in the future.

The radiation properties of the long-lived heavy greenhouse gases, such as CF<sub>4</sub> or SF<sub>6</sub>, are little studied. A common feature of these molecules is the relatively low frequency of oscillations [7] and a significant population of excited levels. Hence, the “hot bands” corresponding to transitions between excited vibrational states affect the IR absorption even at room temperature, which leads to complex vibrational spectra, where it is difficult to resolve isolated lines. Only a few vibrational bands of SF<sub>6</sub> were analyzed within limited spectral ranges in [8, 9]. For those molecules, the spectroscopic databases (for example, GEISA or HITRAN) contain no high-accuracy data on the spectral line parameters.

To simulate IR spectra and energy levels from first principles, ab initio, it is necessary to obtain the

potential energy surface (PES) or the force field of a high order. These techniques are equivalent, but a technique based on an analytical PES is preferable for practical calculations. To derive an analytical expression for the PES, it is necessary to calculate the quantum-chemical energy for many geometries, the number of which is usually several orders of magnitude greater than the number of PES fitting parameters. In the case of a force field, the force constants are usually calculated by the finite difference method, and the validity of the resulting force field on a large geometric grid is rarely checked. To reduce the number of geometries used in the quantum-chemical calculations, it is necessary to take into account the symmetry of the molecule, which results in a more complex form of the PES. The SF<sub>6</sub> molecule has octahedral structure: the molecule is a cube with an S atom at the center and six F atoms at centers of each face (Fig. 1).

In this work, we discuss the calculation of the PES of the SF<sub>6</sub> molecule, which is to be used for calculations of the vibrational-rotation spectra with variational techniques. The equilibrium geometry of the SF<sub>6</sub> molecule is calculated, one-dimensional cross-sections are given, a grid of geometries is obtained, and a technique for PES construction is described.

## 1. EQUILIBRIUM GEOMETRY AND SYMMETRY COORDINATES

Several experimental and theoretical works were devoted to the study of the SF<sub>6</sub> equilibrium geometry [10, 11]. In our work, all ab initio calculations were carried out with MOLPRO software [12] using the

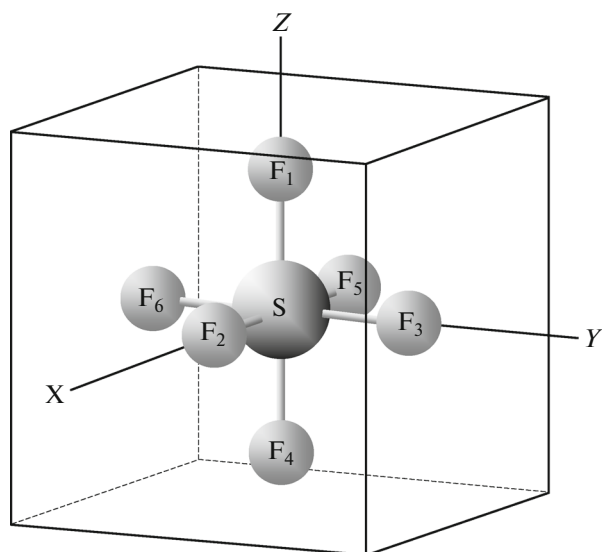


Fig. 1. Internal coordinates in the octahedral molecules.

coupled clusters (CCSD(T)) and the Møller–Plesset perturbation theory of the 2nd and 3rd orders (MP2 and MP3) [13, 14]. The CCSD(T) method successfully works for octahedral molecules [7]. This technique is based on the cluster expansion of the wave function in single and double electron excitations and takes into account a correction on the triple excited states according to perturbation theory. To calculate the equilibrium geometry, we used the correlation-consistent bases. The results obtained for optimized geometry and energy are compared in Table 1.

It is obvious that the CCSD (T)/VTZ technique overestimates the S–F bond length (5701 Å) and energy (–996.076 Hartree). Consideration of the correlation of internal electrons reduces the S–F bond length. The use of the VQZ basis improves the calculation results, which correlate with those from [16, 15], but considerably increases the PES computation time. Taking into account scalar relativistic effects, cc-pVTZ-DK does

**Table 1.** Energies and the equilibrium lengths of the S–F bond obtained with the optimized CCSD(T) technique using different basis sets

| Technique/basis set     | Energy, Hartree | Equilibrium length S–F, Å |
|-------------------------|-----------------|---------------------------|
| MP3/VTZ                 | –996.00822804   | 1.55952379                |
| MP2/VTZ                 | –996.02881762   | 1.57197994                |
| CCSD(T)/VTZ             | –996.07574167   | 1.57015425                |
| CCSD(T)/CVTZ            | –996.79947177   | 1.56003654                |
| CCSD(T)/AVTZ            | –996.14253699   | 1.57357635                |
| CCSD(T)/CVTZ–DK         | –998.23964874   | 1.56042314                |
| CCSD(T)/VQZ             | –996.30894816   | 1.56375885                |
| Ref. [15]               |                 | 1.5647                    |
| Experimental value [16] |                 | 1.5622                    |

not improve the precision of description of the SF<sub>6</sub> equilibrium geometry. We have selected the cc-pVTZ set for further calculations as a compromise between the precision and computation time. Because of the large number of electrons, quantum-chemical methods cannot be used to obtain the SF<sub>6</sub> equilibrium geometry with a precision comparable with experimental measurements.

Consideration of the symmetry with the use of the irreducible tensor representation lets one obtain an optimal set of atom configurations and to minimize the computation time. To calculate the 15D PES for a large grid of the equilibrium geometries, we use the geometry grid generation technique for the CH<sub>4</sub> molecule described in [17, 18]. This technique allows us to perform spectral-precision calculations ab initio in a wide range of energies. To parameterize all possible configurations of the SF<sub>6</sub> molecule, it is preferable to use polar coordinates because they are built into software packages ab initio, such as MOLPRO. The internal polar coordinates are defined in a standard way: six vectors connecting the S atom with the F atoms; six lengths of the S–F bonds  $\{r_1, r_2, r_3, r_4, r_5, r_6\}$ ; five angles between the bonds  $\{q_{12}, q_{13}, q_{14}, q_{15}, q_{16}\}$ , and four torsion angles  $\{t_{23}, t_{24}, t_{25}, t_{26}\}$  (see Fig. 1). The symmetry coordinates are recommended to be used for the analytical representation of PES, which are linear combinations of the internal coordinates.

A seven-atom molecule has 15 vibrational modes. However, different works report different sets of the symmetry coordinates, and most of those sets are incompatible with the method of the irreducible tensorial operators for the O<sub>h</sub> group [19]. In this work, we use the basis elements of the O<sub>h</sub> group (Table 2) for composing the following set of the symmetry coordinates (1)–(18).

For  $A_{1g}$  vibration:

$$S_{A_{1g}}(r) = \frac{1}{\sqrt{6}}(r_1 + r_2 + r_3 + r_4 + r_5 + r_6), \quad (1)$$

$$S_{A_{1g}}(q) = \frac{1}{\sqrt{8}}(q_{12} + q_{13} + q_{15} + q_{16} + q_{23} + q_{24} + q_{26} + q_{34} + q_{35} + q_{45} + q_{46} + q_{56}). \quad (2)$$

For  $E_g$  vibration:

$$S_{E_{ga}}(r) = \frac{1}{\sqrt{12}}(2r_1 + 2r_4 - r_2 - r_3 - r_4 - r_5), \quad (3)$$

$$S_{E_{gb}}(r) = \frac{1}{2}(r_2 - r_3 + r_5 - r_6), \quad (4)$$

$$S_{E_{ga}}(q) = \frac{1}{\sqrt{24}}(q_{12} + q_{13} + q_{15} + q_{16} + q_{24} + q_{34} + q_{45} + q_{46} - 2q_{23} - 2q_{26} - 2q_{35} - 2q_{56}), \quad (5)$$

$$S_{E_{gb}}(q) = \frac{1}{\sqrt{8}}(q_{13} + q_{16} + q_{34} + q_{46} - q_{12} - q_{15} - q_{24} - q_{45}). \quad (6)$$

**Table 2.** O<sub>h</sub> group elements

| Group element                             | $C_{31}^+$                                                                                              | $C_{4z}^+$                                                                                        | Inversion                                                                                              |
|-------------------------------------------|---------------------------------------------------------------------------------------------------------|---------------------------------------------------------------------------------------------------|--------------------------------------------------------------------------------------------------------|
| Transformation matrix                     | $\begin{pmatrix} 0 & 0 & 1 \\ 1 & 0 & 0 \\ 0 & 1 & 0 \end{pmatrix}$                                     | $\begin{pmatrix} 0 & -1 & 0 \\ 1 & 0 & 0 \\ 0 & 0 & 1 \end{pmatrix}$                              | $\begin{pmatrix} -1 & 0 & 0 \\ 0 & -1 & 0 \\ 0 & 0 & -1 \end{pmatrix}$                                 |
| Equivalent permutations of atoms (1 to 6) | $\begin{pmatrix} 1 & 2 & 3 & 4 & 5 & 6 \\ 3 & 1 & 2 & 6 & 4 & 5 \end{pmatrix}$<br>or<br>(1 2 3) (4 5 6) | $\begin{pmatrix} 1 & 2 & 3 & 4 & 5 & 6 \\ 1 & 6 & 2 & 4 & 3 & 5 \end{pmatrix}$<br>or<br>(2 6 5 3) | $\begin{pmatrix} 1 & 2 & 3 & 4 & 5 & 6 \\ 4 & 5 & 6 & 1 & 2 & 3 \end{pmatrix}$<br>or<br>(14) (25) (26) |

For  $F_{1u}$  vibration:

$$S_{F_{1ux}}(r) = \frac{1}{\sqrt{2}}(r_2 - r_5), \quad (7)$$

$$S_{F_{1uy}}(r) = \frac{1}{\sqrt{2}}(r_3 - r_6), \quad (8)$$

$$S_{F_{1uz}}(r) = \frac{1}{\sqrt{2}}(r_1 - r_4), \quad (9)$$

$$S_{F_{1ux}}(q) = \frac{1}{\sqrt{8}}(q_{12} + q_{23} + q_{24} + q_{26} - q_{15} - q_{35} - q_{45} - q_{56}), \quad (10)$$

$$S_{F_{1uy}}(q) = \frac{1}{\sqrt{8}}(q_{13} + q_{23} + q_{34} + q_{35} - q_{16} - q_{26} - q_{46} - q_{56}), \quad (11)$$

$$S_{F_{1uz}}(q) = \frac{1}{\sqrt{8}}(q_{12} + q_{13} + q_{15} + q_{16} - q_{24} - q_{34} - q_{45} - q_{46}). \quad (12)$$

For  $F_{2u}$  vibration:

$$S_{F_{2ux}}(q) = \frac{1}{\sqrt{8}}(q_{15} + q_{23} + q_{26} + q_{45} - q_{12} - q_{24} - q_{35} - q_{56}), \quad (13)$$

$$S_{F_{2uy}}(q) = \frac{1}{\sqrt{8}}(q_{13} + q_{26} + q_{34} + q_{56} - q_{16} - q_{23} - q_{35} - q_{46}), \quad (14)$$

$$S_{F_{2uz}}(q) = \frac{1}{\sqrt{8}}(q_{12} + q_{15} + q_{34} + q_{46} - q_{13} - q_{16} - q_{24} - q_{45}). \quad (15)$$

For  $F_{2g}$  vibration:

$$S_{F_{2gx}}(q) = \frac{1}{2}(q_{13} + q_{46} - q_{16} - q_{34}), \quad (16)$$

$$S_{F_{2gy}}(q) = \frac{1}{2}(q_{12} + q_{45} - q_{15} - q_{24}), \quad (17)$$

$$S_{F_{2gz}}(q) = \frac{1}{2}(q_{23} + q_{56} - q_{26} - q_{35}). \quad (18)$$

Radial coordinates (1), (3), (4), and (7)–(9) are compatible with those from [20]. A left-handed coordinate system is used in [21], and not all sets are

given. The atomic numeration used in [22] differs from our definition in permutation (23)(56). This permutation turns the  $F_{1u}$  coordinates [22] into (7)–(12). Let  $S^*$  denote the symmetry coordinates from [22], and  $S$ , Eqs. (7)–(18). In this case,  $S_x^* = S_y$ ,  $S_y^* = S_x$ , and  $S_z^* = S_z$  for  $F_{1u}$ . This corresponds to the permutation of the  $x$  and  $y$  components following the permutation (23)(56). The  $F_{2u}$  and  $F_{2g}$  coordinates from [22] are connected with Eqs. (12)–(18) in the following way:  $S_x^* = S_z$ ,  $S_y^* = S_x$ , and  $S_z^* = S_y$ . This transformation does not correspond to the permutation (23)(56). In this case, the following expressions are valid: for  $F_{2u}$  coordinates (13)–(16): (23)(56)  $S_x = -S_y$ , (23)(56)  $S_y = -S_x$ , (23)(56)  $S_z = -S_z$ ; and for  $F_{2g}$  coordinates (16)–(17): (23)(56)  $S_x = S_y$ , (23)(56)  $S_y = S_x$ , (23)(56)  $S_z = S_z$ . The above-considered permutations correspond to the geometric interpretation of the coordinate axis shown in Fig. 1. On the other hand, the  $F_{2u}$  coordinates from [22] can be transformed following the rules (23)(56)  $S_x^* = -S_x^*$ , (23)(56)  $S_y^* = S_z^*$ , which does not correspond to the molecular configuration shown in Fig. 1 from [22]. Our  $F_{2u}$  coordinates (13)–(15) differ from those in [23]. The coordinates in [22, 23] form a full set of coordinates, but the use of them in the irreducible tensorial operators method [19] does not result in suitable solutions for the PES of the third order or higher.

The technique [18] has been used to define a unique set of geometries that describe the PES. It is based on the force field constants and lets us obtain an optimal set of the atomic configuration geometries sufficient for modeling the force field up to a certain order. In the first step, one-dimensional axial cross-sections are constructed (Fig. 2), and then the symmetry coordinates are found where the energy takes the following values: 1000, 2500, 4000, 7000, 10000, 12500, and 15000 cm<sup>-1</sup>.

#### ANALYTICAL EXPRESSIONS FOR PES AND OPTIMAL GRID OF GEOMETRIES

The standard representation in terms of the irreducible tensorial operators [19] has been used to

derive analytical expressions for 15D PES. Three radial vibrational modes  $\nu_1$ – $\nu_4$  are transformed under the  $A_{1g}$  (774.5 cm<sup>-1</sup>),  $E_g$  (643.3 cm<sup>-1</sup>), and  $F_{1u}$  (948.1 cm<sup>-1</sup>) representations; three bending modes

$\nu_4$ – $\nu_6$  are transformed under the  $F_{1u}$  (615.0 cm<sup>-1</sup>),  $F_{2g}$  (524.0 cm<sup>-1</sup>), and  $F_{2u}$  (347.7 cm<sup>-1</sup>) representations. The potential function is written in terms of the irreducible tensorial operators:

$$V(R) = \sum_p F_p \Pi_{i\sigma} (S_{i\sigma}^{(\Gamma_i)})^{p_{i\sigma}} \quad (19)$$

$$= \sum_p F_p \times R_i^p(r_1, r_2, r_3, r_4, r_5, r_6, q_{12}, q_{13}, q_{14}, q_{15}, q_{16}, t_{23}, t_{24}, t_{25}, t_{26}),$$

where

$$R_i^p = \left( \left( [SR_{A_{1g}}]^{p_1} \times ([SR_{E_g}]^{p_2} \times [SR_{F_{1u}}]^{p_3})^{C_{23}} \right)^{C_{23}} \right)^{C_{23}} \times \left( [SR_{F_{1u}}]^{p_4} \times ([SR_{F_{2g}}]^{p_5} \times [SR_{F_{2u}}]^{p_6})^{C_{56}} \right)^{C_{56}} \quad (20)$$

Here  $p = p_1 + p_2 + \dots + p_6$  is the term degree;  $A_{1g}$ ,  $E_g$ ,  $F_{1u}$ ,  $F_{2u}$ , and  $F_{2g}$  are the nondegenerate, double and triple degenerate irreducible representations of the

$O_h$  group; and  $F_p$  is the set of parameters for the PES, ab initio.

The number of the irreducible tensors of the  $A_{1g}$  symmetry up to the 4th, 5th, 6th, 7th, and 8th (included) order are 123, 406, 1334, 3924, and 10 994, respectively. For building the grid of geometries, all tensors were divided into several groups, each group included tensors of the same degrees  $p_1$ ,  $p_2$ ,  $p_3$ ,  $p_4$ ,  $p_5$ , and  $p_6$ . For each tensor group, a set of the symmetry degrees was found, which uniquely identified that group. After that, we obtained a possible set of the force field constants that uniquely define the PES up to the 6th order. From that set, we removed all points that coincide after any of 48 permutations of the F atoms. As a result, a grid of 6412 geometries was constructed that uniquely defined the PES up to the sixth order. The study confirms that the quantum-chemical energies for these geometries are different.

## PES FITTING

The lengths  $r_i$  and the angles between bonds  $q_{ij}$  have been used as one-dimensional basis functions. As a rule, for parameterizing PES, to provide the best results using the least-squares technique, more complex one-dimensional basis functions are used. Morse functions are used for radial functions:

$$f_R(r_i) = 1 - \exp[-\alpha(r_i - r_e)], \quad (21)$$

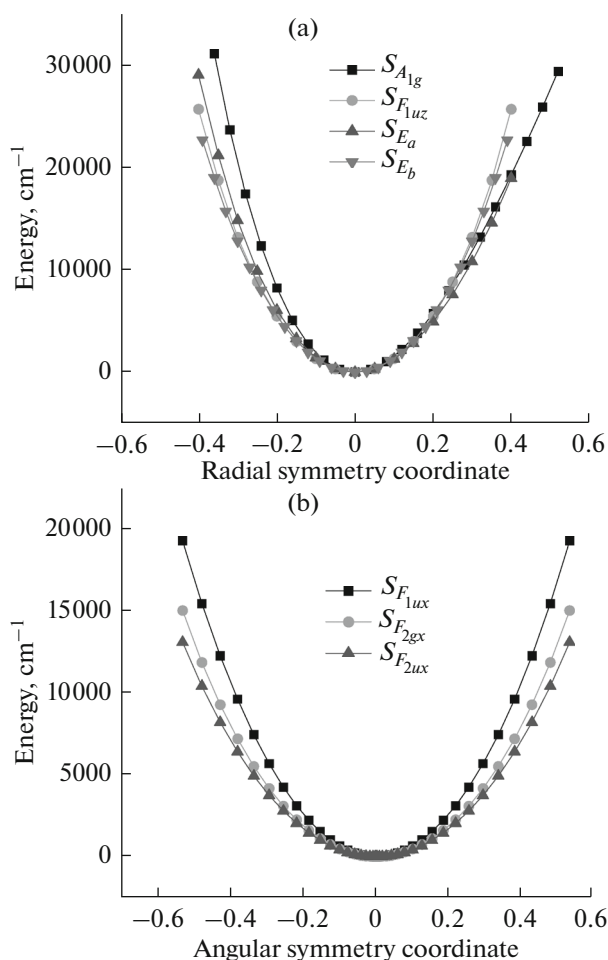
where the parameter  $\alpha$  depends on an approximate expression for one-dimensional radial potential

$$V_0 (1 - \exp[-\alpha(r_i - r_e)])^2; \quad (22)$$

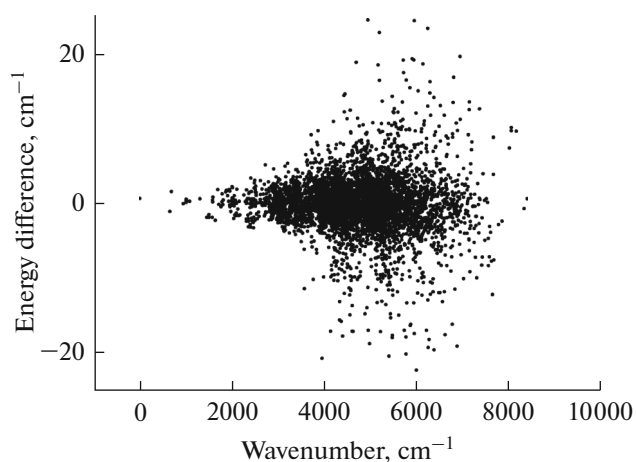
$\alpha = 1.9$  for SF<sub>6</sub>. The cosine function is often used for angular functions. However, by analogy with [17], the functions

$$f_A(q_{ij}) = (\cos q_{ij} - \cos q_e) + \alpha(\cos q_{ij} - \cos q_e)^2 \quad (23)$$

turned out preferable; here,  $\alpha = 0.3333$ . For example, the PES fitting with the hyperbolic tangent with the parameter equal to 7000 [17] as a weight function results in the following standard errors: 3.7 cm<sup>-1</sup> for the cosine function and 2.3 cm<sup>-1</sup> for Eq. (22). The high quality of the fourth order PES fitting proves a high rigidity of the SF<sub>6</sub> molecule. Figure 3 shows the dependence of the standard errors on the geometry



**Fig. 2.** One-dimensional axial cross-sections of SF<sub>6</sub> in the dimensionless symmetry coordinate system.



**Fig. 3.** Standard error of the PES fitting vs the geometry energy for SF<sub>6</sub>, MP2/VTZ technique.

energy. We have obtained 123 PES constants using the MP2 technique with the cc-VTZ basis set.

## CONCLUSIONS

In this paper, an analytical form of the SF<sub>6</sub> potential surface has been obtained using the VTZ basis set. This form includes 123 PES constants up to the fourth order. It is shown that the structure of one-dimensional angular functions considerably affects the PES fitting accuracy. In our further studies, we plan to calculate the PES of higher orders and improve the computation accuracy due to the use of the high-precision CCSD(T) technique with the VTZ basis set. We also plan to use the contact transformations technique [24] for modeling the PES-based effective Hamiltonian. This is to help us simplify simulation of the SF<sub>6</sub> spectra and include them in the information database [25].

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## CONFLICT OF INTEREST

The authors declare that they have no conflicts of interest.

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*Translated by O. Ponomareva*

First Full-Dimensional Potential Energy and Dipole Moment Surfaces of SF<sub>6</sub>

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# First Full-Dimensional Potential Energy and Dipole Moment Surfaces of SF<sub>6</sub>

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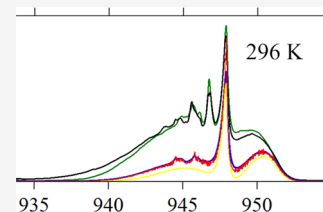


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Supporting Information

**ABSTRACT:** A 15-dimensional analytical form for the potential energy and dipole moment surfaces of the SF<sub>6</sub> molecule in the ground electronic state is obtained using *ab initio* methods. In order to calculate the equilibrium S–F distance, we applied the coupled cluster CCSD(T) method and several versions of the correlation-consistent basis sets from valence triple-zeta (VTZ) and augmented valence triple-zeta (AVTZ) to core-valence quadruple-zeta (CVQZ) with Douglas–Kroll (DK) relativistic corrections that provided good agreement with an empirical equilibrium value. *Ab initio* electronic energies on 15D grids of nuclear geometries are computed using the CCSD(T) method with VTZ and CVQZ-DK basis sets. The analytical representation of the potential energy surface is determined through an expansion in symmetry-adapted products of nonlinear coordinates up to the 5th order. The influence of additional redundant coordinates on the quality of the fit was investigated. Parameters of full-dimensional dipole moment surfaces are determined up to the 4th order expansion in normal mode coordinates. For validation of *ab initio* results, the fundamental vibration frequencies and absorption cross sections of the principal sulfur hexafluoride isotopologue are calculated, giving good agreement with cold (180 K) and room temperature (296 K) experimental spectra. Absorption cross sections calculated from our preliminary line list agree much better with these observations than the simulations using SF<sub>6</sub> line-by-line lists constructed from effective models included in currently available spectroscopic databases.



## 1. INTRODUCTION

Investigation of electronic structure and radiative properties of high octahedral symmetry sulfur hexafluoride (SF<sub>6</sub>) molecule is of interest both for the theory<sup>1–3</sup> and for applications in many domains of industry and medicine. Because of strong absorption bands in the atmospheric transparency windows, SF<sub>6</sub> is a powerful greenhouse gas with the global warming potential (GWP500) 29,300 times larger than that of CO<sub>2</sub>.<sup>4</sup> At present, the sulfur hexafluoride atmospheric concentration is yet small, but SF<sub>6</sub> is one of the most stable gaseous chemicals with an atmospheric life time of 3200 years.<sup>5</sup> According to the conclusions of Earth System Research Laboratory Global Monitoring Division,<sup>6</sup> the SF<sub>6</sub> concentration increased from 3.5 ppt (1995) to 9.5 ppt (2018). A rapid increase of the SF<sub>6</sub> concentration and its long atmospheric life time might result in a considerable accumulation of SF<sub>6</sub> in the troposphere. Thus, in future, sulfur hexafluoride may play an important role in the climate global changing. The radiative properties of SF<sub>6</sub> were insufficiently studied until now. Because of relatively low-frequency vibrations,<sup>7</sup> the Boltzmann population of the excited vibrational levels appears to be significant. Thus, the “hot bands” corresponded to transitions between excited vibrational states produce a considerable contribution to IR absorption even at room temperature. The ro-vibrational spectra are quite complex and crowded making it difficult to identify isolated lines in experimentally recorded absorption samples. Only few vibrational bands of SF<sub>6</sub> within limited spectral ranges have been studied<sup>8–12</sup> using the effective Hamiltonian approach. Despite many efforts, a simulation of the room-temperature

SF<sub>6</sub> absorbance using effective spectroscopic models was not sufficiently complete. For this reason, the most advanced spectroscopic databases (as HITRAN<sup>13</sup> or GEISA<sup>14</sup>) do not contain line parameters for all necessary SF<sub>6</sub> bands in order to describe the absorbance at room temperature with high accuracy.

An approach based on *ab initio* potential energy surfaces (PES) and dipole moment surfaces (DMS) using variational methods for the solution of the full-dimensional nuclear motion Schrödinger equation has been successfully applied to many 4–5 atomic<sup>7,15–22</sup> and also to six-atomic<sup>23–25</sup> molecules such as C<sub>2</sub>H<sub>4</sub>. In<sup>26</sup> some calculated levels of UF<sub>6</sub>, molecules having the same octahedral symmetry have been reported. However, high precision full-dimensional PESs and DMSs for seven-atomic molecules are still lacking in the literature.

Several works used alternative methods for the nuclear motion calculations for medium and large molecules using vibrational configuration-interaction approach or the perturbation theory (refs<sup>27–29</sup> and references herein). However, in most cases, the rotation states were computed for low *J* only and the contribution of hot bands have been neglected.

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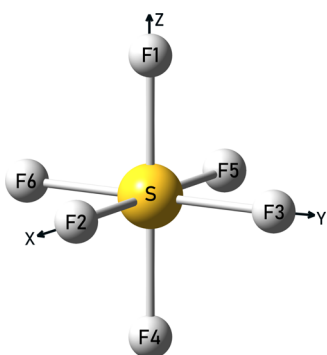
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To compute energy levels and IR spectra beyond the harmonic oscillator approximation using *ab initio* techniques, it is necessary to obtain either the PES on a sufficiently large grid of nuclear displacements or the force field with high-order derivatives. Both techniques are often used, but for high-resolution applications, the technique based on an analytical PES determined at an extended geometrical grid has proved to produce more accurate results.<sup>21,26,30</sup> In the full-grid approach, the computational difficulties are related to a simultaneous account of relatively a large number of electrons (70 electrons in case of SF<sub>6</sub>) and of a large number of degrees of freedom including representative nuclear geometries for PES and DMS parametrization. As a rule, the number of different geometries is several times bigger than the number of the fitted PES parameters. The force field technique often uses the finite difference method for calculating the force constants, and as a rule, a validity of the obtained force field could not be checked for sufficiently large nuclear displacements. To reduce the number of nuclear geometries in the electronic structure calculations, it is necessary to take into account the molecular symmetry.

This paper is devoted to first calculations of the full-dimensional PES and DMS of the SF<sub>6</sub> molecule and is structured as follows. Equilibrium structure at various basis sets, the coordinate choice, and the analytical model for PES and DMS surfaces are considered in Section 2, whereas *ab initio* calculations and fits are discussed in Sections 3 and 4. Calculation of the vibrational-rotation spectra using the variation techniques<sup>31–34</sup> shows a good agreement with experimental absorption cross sections in the range of the strongest bands of <sup>32</sup>SF<sub>6</sub> and <sup>34</sup>SF<sub>6</sub> isotopologues (Section 5), providing a preliminary validation of *ab initio* results.

## 2. EQUILIBRIUM GEOMETRY AND ANALYTICAL FORM OF PES AND DMS

The SF<sub>6</sub> molecule belongs to the octahedral O<sub>h</sub> point group: at the equilibrium geometry, the molecule is a cube with the S atom at the cube center and six F atoms at centers of each edge (Figure 1). Thus, only one geometrical parameter (S–F



**Figure 1.** Equilibrium structure and the axes choice for the SF<sub>6</sub> molecule.

distance) determines this equilibrium structure unambiguously. Experimental and theoretical studies of the SF<sub>6</sub> equilibrium geometry have been considered in.<sup>35,36</sup> In order to calculate the equilibrium S–F distance, we applied coupled cluster CCSD(T) method and several versions of the correlation consistent basis sets from valence triple-zeta (VTZ) and augmented valence triple-zeta (AVTZ) to core-

valence quadruple-zeta (CVQZ) with Douglas–Kroll (DK) relativistic corrections. All *ab initio* calculations were carried out using MOLPRO suite of programs as implemented in the 2019.1 version.<sup>37,38</sup> More details concerning *ab initio* methods will be given in Section 3. A comparison of results obtained for various basis sets with some empirical values is collected in Table 1.

**Table 1. Energies and the Equilibrium Lengths of the S–F Bond Obtained with the Optimization CCSD(T) Technique Using Various Basis Sets**

| method/basis set                   | energy, hartree | equilibrium length S–F, Å |
|------------------------------------|-----------------|---------------------------|
| CCSD(T)/VTZ                        | –996.07574167   | 1.57015425                |
| CCSD(T)/CVTZ                       | –996.79947177   | 1.56003654                |
| CCSD(T)/AVTZ                       | –996.14253699   | 1.57357635                |
| CCSD(T)/CVTZ-DK                    | –998.23964874   | 1.56042314                |
| CCSD(T)/VQZ                        | –996.30894816   | 1.56375885                |
| CCSD(T)/CVQZ                       | –997.10132168   | 1.55654004                |
| CCSD(T)/CVQZ-DK                    | –998.70581919   | 1.55702740                |
| empirical variational <sup>a</sup> |                 | 1.556                     |
| ref 39                             |                 | 1.5647                    |
| empirical value <sup>40</sup>      |                 | 1.5622                    |

<sup>a</sup>Preliminary result of empirical optimization of S–F distance using variational calculations of rotation energies.<sup>7,32</sup>

Two calculations with two basis sets CVQZ and CVQZ-DK give values close to the empirical ones. In any case without additional corrections, it is impossible to calculate the equilibrium geometry with a precision better than 0.001 Å°.

Various sets of the symmetry coordinates have been used in many previous works. Most of them are incompatible with our definitions of Clebsch–Gordan coefficients for the O<sub>h</sub> group.<sup>41,42</sup> Here, we use the symmetric coordinates and the molecule orientation from the work.<sup>43</sup> Initially, we begin with linear combinations of bond lengths and bond angles, which will be further adapted using non-linear elementary morse-type and cosine-type functions as described in Section 3. At the first step, six “radial” (stretching) symmetrized coordinates were used to accommodate three irreducible representations: A<sub>1g</sub>, E<sub>g</sub>, and F<sub>1u</sub>.

$$SR_{A_{1g}} = (r_1 + r_2 + r_3 + r_4 + r_5 + r_6 - 6r_e)/\sqrt{6} \quad (1)$$

$$SR_{E_g(a)} = (2r_1 + 2r_4 - r_2 - r_3 - r_5 - r_6)/\sqrt{12} \quad (2)$$

$$SR_{F_{1u}(x)} = (r_2 - r_5)/\sqrt{2} \quad (3)$$

Nine angular coordinates  $q_{ij}$  were symmetrized according to three irreducible representations: F<sub>1u</sub>, F<sub>2g</sub>, and F<sub>2u</sub>.

$$SA_{F_{1u}(x)} = (q_{12} + q_{23} + q_{24} + q_{26} - q_{15} - q_{35} - q_{45} - q_{56})/\sqrt{8} \quad (4)$$

$$SA_{F_{2g}(x)} = (q_{13} + q_{46} - q_{16} - q_{34})/2 \quad (5)$$

$$SA_{F_{2u}(x)} = (q_{15} + q_{23} + q_{26} + q_{45} - q_{12} - q_{24} - q_{35} - q_{56})/\sqrt{8} \quad (6)$$

Some of these  $q_{ij}$  angles in Figure 1, which take  $\pi/2$  values at the equilibrium, are called “adjacent” angles. Others, which take  $\pi$  values at the equilibrium are called “opposition” angles.

The number of possible angles is bigger than the number of symmetrized angular coordinates. Hence, two types of additional symmetrized angular coordinates can be added: three coordinates  $S''$  of symmetry  $A_{1g}$ ,  $E_g$  composed of “opposition” angles  $q_{14}$ ,  $q_{25}$ , and  $q_{36}$ , or three coordinates  $S'$  of the same symmetry  $A_{1g}$ ,  $E_g$  from 12 “adjacent” angles

$$\begin{aligned} S''_{E_a} &= (q_{25} + q_{36} - 2q_{14})/\sqrt{6} \\ S''_{E_b} &= (q_{36} - q_{25})/\sqrt{6} \\ S''_{A_{1g}} &= (q_{25} + q_{36} + q_{14} - 3q_e'')/\sqrt{3} \\ S'_{E_a} &= (q_{12} + q_{15} + q_{34} + q_{46} + q_{13} + q_{16} + q_{24} + q_{45} \\ &\quad - 2q_{23} - 2q_{35} - 2q_{56} - 2q_{26})/\sqrt{24} \\ S'_{E_b} &= (q_{12} + q_{24} + q_{15} + q_{45} - q_{13} - q_{46} - q_{34} - q_{16}) \\ &\quad / \sqrt{8} \\ S'_{A_{1g}} &= (q_{12} + q_{15} + q_{34} + q_{46} + q_{13} + q_{16} + q_{24} + q_{45} \\ &\quad + q_{23} + q_{35} + q_{56} + q_{26} - 12q_e')/\sqrt{12} \end{aligned} \quad (7)$$

Here,  $r_e$  is the equilibrium values for the S–F distance and  $q_e''$  and  $q_e'$  are equilibrium values for “adjacent” or “opposition” angles. The additional angular coordinates can significantly accommodate the fit of the shape of the surface near the bottom of the potential well that will affect the energy levels and wave functions. In the case of methane, this has been shown in the ref 21.

The complete set of coordinates and the  $O_h$  group together with the matrices of irreducible representations (irrep) were reported in ref 43. As explained in ref 43, our set of internal angular coordinates is different from that of refs.<sup>44–46</sup> The definition of the radial symmetrized coordinates (3) coincides with that of ref 47. Matrix representation of group elements<sup>43</sup> (see Table 1 of Appendix 1 for matrices) coincides with that of ref 48 as given in our Table 2.

**Table 2.**  $O_h$ : Point Group Elements

| group element                                | $C_{31}^+$                                                            | $C_{4z}^+$                                                           | inversion                                                              |
|----------------------------------------------|-----------------------------------------------------------------------|----------------------------------------------------------------------|------------------------------------------------------------------------|
| transformation matrix                        | $\begin{pmatrix} 0 & 0 & 1 \\ 1 & 0 & 0 \\ 0 & 1 & 0 \end{pmatrix}$   | $\begin{pmatrix} 0 & -1 & 0 \\ 1 & 0 & 0 \\ 0 & 0 & 1 \end{pmatrix}$ | $\begin{pmatrix} -1 & 0 & 0 \\ 0 & -1 & 0 \\ 0 & 0 & -1 \end{pmatrix}$ |
| equivalent permutations of F atoms (eqs 1–6) | $\begin{pmatrix} 123456 \\ 312645 \end{pmatrix}$<br>or (132)<br>(465) | $\begin{pmatrix} 123456 \\ 162435 \end{pmatrix}$<br>or (2653)        | $\begin{pmatrix} 123456 \\ 456123 \end{pmatrix}$<br>or (14)(25)(26)    |

An account of the symmetry and the use of the irreducible tensor representation permitted obtaining an optimal non-redundant grid of nuclear geometries. One grid adapted for the analytical PES development up to the 6-th order (6412 geometries) has been constructed in the work<sup>43</sup> using the geometry set generation technique already applied for other symmetric molecules.<sup>49,50</sup> Here, we use also extended grid of 78,739 geometries that unambiguously define the PES development up to 8-th order constructed by the same technique. These two grids of geometries allowed us to carry out *ab initio* electronic energy calculations for a wide range of

nuclear displacements. Initially, the grids were constructed in symmetrized coordinates (eqs 1–6), but for *ab initio* calculations both grids were recalculated in internal coordinates as in ref 43: six vectors connecting the S atom with the F atoms; six lengths of the S–F bonds  $\{r_1, r_2, r_3, r_4, r_5, r_6\}$ ; five angles between the bonds  $\{q_{12}, q_{13}, q_{14}, q_{15}, q_{16}\}$ , and four torsion angles  $\{t_{23}, t_{24}, t_{25}, t_{26}\}$  (see Figure 1).

In the analytical expressions for PES, it is recommended to use the symmetry coordinates that are linear combinations of the internal coordinates or of 1D elementary functions of these internal coordinates. A seven-atomic molecule has 15 vibrational degrees of freedom. The standard representation in terms of the irreducible tensor operators (ITO)<sup>41,42</sup> was used to obtain analytical expressions for 15-dimensional (15D) PES and DMS developments. Three stretching vibrational modes  $\nu_1$  (non-degenerate at 774.5  $\text{cm}^{-1}$ ),  $\nu_2$  (two-fold degenerate at 643.3  $\text{cm}^{-1}$ ), and  $\nu_3$  (three-fold degenerate at 948.1  $\text{cm}^{-1}$ ) transform according to the  $A_{1g}$ ,  $E_g$ , and  $F_{1u}$  irreducible representations, while three triply degenerate bending modes  $\nu_4$ ,  $\nu_5$ , and  $\nu_6$  transform under the  $F_{1u}$  (615.0  $\text{cm}^{-1}$ ),  $F_{2g}$  (524.0  $\text{cm}^{-1}$ ), and  $F_{2u}$  (347.7  $\text{cm}^{-1}$ ) representations. The analytic potential function is written in terms of the ITO

$$V(R) = \sum_{i,p} F_{ip} R_i^p(r_1, r_2, r_3, r_4, r_5, r_6, q_{12}, q_{13}, q_{14}, q_{15}, q_{16}, t_{23}, t_{24}, t_{25}, t_{26}) \quad (9)$$

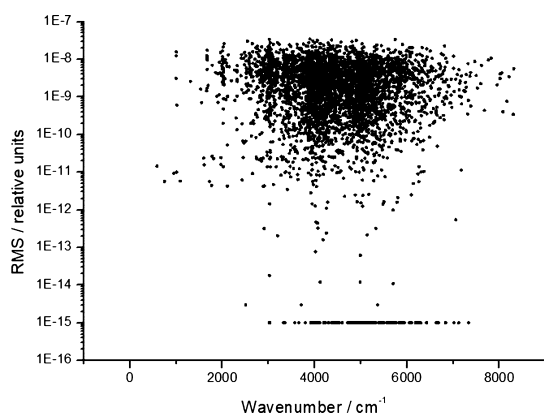
where each  $R_i^p$  tensor is obtained as an appropriate product of radial  $SR_\Gamma$  and angular  $SA_\Gamma$  non-linear symmetrized coordinates

$$R_i^p = ([SR_{A_{1g}}]^{p_1} \times [SR_{E_g}]^{p_2} \times [SR_{F_{1u}}]^{p_3})^{C_{23}} C' \times ([SA_{F_{1u}}]^{p_4} \times [SA_{F_{2g}}]^{p_5} \times [SA_{F_{2u}}]^{p_6})^{C_{56}} C'' \quad (10)$$

These non-linear radial  $SR_\Gamma$  and angular  $SA_\Gamma$  non-linear symmetrized coordinates involved in our PES model (eqs 9 and 10) were obtained by replacing the combinations (eqs 1–6) by the morse-type and cosine-type elementary functions in a similar way as was described in our works for the methane molecule.<sup>21,49,51</sup> We have tried several choices for an optimal fit to *ab initio* points. The final expressions of elementary functions used in the present work are given in the next Section 3. Here,  $p_i$  components of  $p = \{p_1, \dots, p_6\}$  correspond to powers of these coordinates, and  $p = p_1 + p_2 + \dots + p_6$  is the total power of the expansion term. The lower case index  $i$  corresponds to the set of symbols  $C_{23}$ ,  $C_{56}$ ,  $C$ ,  $C'$ , and  $C''$  in the summation (eq 9) that run over nondegenerate ( $A_{1g}$ ), doubly ( $E_g$ ), and triply degenerate ( $F_{1u}$ ,  $F_{2u}$ , and  $F_{2g}$ ) irreducible representations (irreps) of  $O_h$  point group. The  $C_{23}$  irrep results from the coupling of the symmetrized powers of the second and the third vibrational modes, whereas  $C'$  stands for the irrep corresponding to the coupling of the latter product with the symmetrized power of the first vibrational mode. In a similar manner, the irrep  $C_{56}$  corresponds to the coupling of the fifth and sixth modes coupled with the fourth mode via  $C''$ . For the PES expansion,  $C' = C''$ , so that only the final symmetry allowed type  $C$  of the  $R_i^p$  tensors is  $A_{1g}$ . Very similar expansion can be used for the DMS, but in this case, the final symmetry type  $C$  of the  $R_i^p$  tensors is  $F_{1u}$ . Each combination of the symmetrized coordinate powers  $p_1, \dots, p_6$  and of the coupling irreps  $C_{23}$ ,  $C_{56}$ ,  $C$ ,  $C'$ , and  $C''$  form a “tree”.<sup>41,52</sup> An ITO model for the PES or for the DMS

expansion at a given total order  $p$  corresponds to the summation over various symmetry allowed trees<sup>52</sup> with the  $F_{ip}$  parameters fitted to *ab initio* values. If the “redundant” symmetrized coordinates of the type  $S'$  (8) or  $S''$  (6) are included in the PES model,<sup>43</sup> then two more corresponding factors are added at the left-hand side of the ITO tree (eq 10).

The ITO formalism for the PES and DMS expansions can be also applied in terms of normal mode coordinates, as described in refs 31 and 32. This results in a similar expression to eq 10, where the symmetrized  $SR_j$  and angular  $SA_k$  coordinates are replaced by normal coordinates  $Q_i$ . In this work, we used the normal mode ITO representation for the dipole moment modeling and for intensity calculations because the matrix elements of the corresponding operators can be computed analytically. However, a sufficiently complete transformation from nonlinear symmetrized coordinates (eqs 1–6) to rectilinear normal coordinated  $Q_i$  in the case of  $SF_6$  appears to be quite complicated. In this work, we obtained the corresponding relations  $\{SR_j, SA_k\} \leftrightarrow Q_i$  numerically via the solution of the system of non-linear equations. A distribution of the RMS deviations of the fit of 15 normal coordinates to 15 internal symmetrized coordinates is shown in Figure 2. The



**Figure 2.** RMS distribution for non-linear fit of normal coordinates  $Q_i$  to internal symmetrized coordinates  $SR_j, SA_k$ .

error of the corresponding numerical solution  $Q_i = Q_i(SR_j, SA_k)$  was much smaller than the precision of the fit of the analytical form (eq 10) to *ab initio* points and thus did not affect the accuracy of the final calculation of spectra.

### 3. PES CALCULATION

To construct the molecular PES of  $SF_6$ , we first computed a set of *ab initio* electronic energies at various molecular configurations. Equilibrium geometries and vibrational frequencies were computed with the well-established coupled cluster approach CCSD(T) with all single and double substitutions from the Hartree–Fock reference determinant augmented by a perturbative treatment of connected triple excitations.<sup>53,54</sup> It was found to provide a good description for the equilibrium geometry and vibrational frequencies and then used in this work to calculate the first full-dimensional PES of  $SF_6$ . We used correlation-consistent polarized valence  $n$ -tuple-zeta basis sets of Dunning, Peterson and co-workers<sup>55–58</sup> for the calculations and specifically adapted DK basis for calculations including relativistic effects.<sup>59</sup> Energies for all electronic structure calculations were converged to MOLPRO default value  $10^{-10}$  a.u.

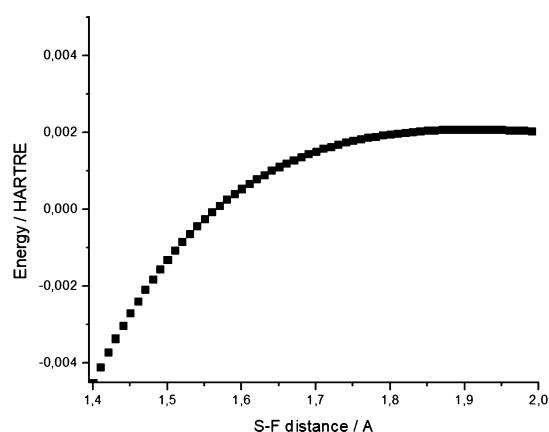
In order to obtain accurate calculations of vibrational and rotational energy levels from the theoretical PES useful for spectroscopic analysis, it is desirable to use high level *ab initio* methods with sufficiently large basis sets in electronic structure calculations. However, in the case of  $SF_6$ , the full-grid calculations with large basis sets (like those recently employed for methane<sup>21,30</sup>) are too demanding. In this work, an essential part of electronic energy calculations was performed with the CCSD(T)/VTZ ansatz at 6412 nuclear geometries of the grid optimized for the 6-th order PES. To refine the analytical PES form, the same method and the same basis set were additionally used at 2296 geometries of the 8-th order grid. In order to control the surface shape at larger nuclear displacements, we carried out supplementary MP2 calculations with VTZ atomic basis on extended grid of 85,136 geometries also optimized for the 8-th order PES.

Empirical PES corrections are often used to compensate the lack of accuracy of *ab initio* calculations. However, for an efficient application of an empirical optimization, these corrections have to be relatively small. Typically, simple empirical corrections are used to reduce calculated minus observed deviation for the fundamental vibration frequencies, but this could produce uncontrolled errors for highly excited states if the shape of the initial *ab initio* PES is not reliably established. This is an important point for a correct modeling of “hot” bands as they could produce a significant contribution to the atmospheric absorbance by heavy molecules with relatively low vibrational modes. In the case of  $SF_6$ , only a limited number of experimental vibrational frequencies are known that can be used for a validation of *ab initio* results. *Ab initio* equilibrium geometry obtained with CVQZ and CVQZ-DK basis sets was close to empirical ones (Table 1), but calculations using CCSD(T)/CVQZ-DK ansatz for a large grid of geometries are actually too demanding. For this reason, we have applied the algorithm of corrections similar to that of ref 49. First, we have computed the one-dimensional “radial” PES section using the CVQZ-DK basis set and electronic energies at 67 nuclear geometries sufficient to determine the force field up to the third order. Then, the difference between electronic energies in 67 geometries using CVQZ-DK and VTZ basis sets was fitted to an analytical function  $F_{O3}(S(r, q))$  of 15 symmetrized coordinates expanded up to 3rd order. One dimensional radial correction functions  $f_X(r)$  were constructed as a difference between two radial potentials calculated for X-basis and for VTZ basis

$$f_X(r) = V_X(r) - V_{VTZ}(r) - F_{O3}(S(r, r_e, r_e, \dots, q_e)) \quad (11)$$

where  $V_X(r)$  stands for electronic energies at the radial cut and  $X = CVQZ-DK$  or  $VQZ$ . The additional one-dimensional function  $F_{O3}(r, 0, 0, \dots, 0)$  was obtained from the third order correction. All other bond lengths were fixed to their equilibrium values for X and VTZ basis sets, whereas all angles were fixed to their equilibrium values. Functions  $f_X(r)$  for  $X = CVQZ-DK$  is shown in Figure 3. Note that the one-dimensional radial correction (eq 11) near the equilibrium geometries of VTZ and X basis sets quite significantly depends on the analytical representation.

The number  $N$  of irreducible tensors of the total  $A_{1g}$  representation involved in the PES expansion (eqs 9 and 10) increases quite rapidly with the order. Up to the orders 2, 3, 4, 5, 6, 7, and 8, this number  $N$  is equal to 9, 31, 123, 406, 1334, 3924, and 10,994, respectively. For PES parameterization, to provide the best results using the least square technique, more



**Figure 3.** One-dimensional radial correction  $E\{\text{CCSD(T)}\text{-DK}/\text{CVQZ-DK}\} - E\{\text{CCSD(T)}/\text{VTZ}\}$ .

complex one-dimensional basis functions are used in (eqs 1–6). Morse-type functions are used as primitive radial functions

$$f_r(r_i) = 1 - \exp(-\alpha(r_i - r_e)) \quad (12)$$

involved in the definitions of the symmetrized  $\text{SR}_F$  and angular  $\text{SA}_F$  coordinates which are used to build our final analytical PES model. Although the function  $f_r(r)^2$  for  $\alpha = 2.0$  reproduces very accurately the one-dimensional potential of SF, the best choice of  $\alpha$  in eq 12 for full-dimensional PES was  $\alpha = 1.9$ . Consequently, our PES model tends to the corresponding finite asymptote at large inter-nuclear distances contrary to a simple power series expansion in  $(r_i - r_e)$ .

The angular function was expressed in terms of the cosine function

$$f_a(q_{ij}) = \cos(q_{ij}) + \alpha \cos^2(q_{ij}) \quad (13)$$

Similar to methane PES,<sup>49</sup> it is preferable to use the value  $\alpha = 1/3$  for angular functions.

The relevance of various analytical representation of the PES was checked by the quality of the PES fit on the 8-th order grid of 85,136 geometries using MP2/VTZ calculated energies. Various orders of the PES expansion and various forms of the angular dependence were tested. The results of the least-squares fits using weight function  $w(E)$  similar to that previously employed for methane studies<sup>49,60</sup>

$$w(E) = \frac{\tanh(-0.0005(E - E_1) + 1.002002002)}{2.002002002} \quad (14)$$

where  $E_1 = 7000 \text{ cm}^{-1}$ , are collected in Table 3. It is seen that the parameter value  $\alpha = 1/3$  permitted a notable improvement of the standard deviation of the fit for 4-th and 5-th order of the PES expansion. The use of supplementary (redundant) coordinated, as reported in the case of methane,<sup>21,51</sup> could significantly change the PES near the equilibrium. In this work, we have also evaluated a possible impact of supplementary coordinates eqs 7 and 8 for  $\text{SF}_6$ . An introduction of  $S'$  coordinate in the PES expansion has somewhat larger effect on the standard deviation than that of  $S''$ , but in both cases, the improvement of the standard deviation was smaller with respect to the inclusion of  $\alpha = 1/3$  in the quadratic terms of eq 13. Finally, we consider that the fifth-order PES expansion provides an acceptable quality of the fit of *ab initio* energies up to  $7000 \text{ cm}^{-1}$ .

**Table 3.** Result of Various PES Fits Using 85136 Electronic MP2/VTZ Energies

| order | additional angular coordinates: order, type | #parameters | $\alpha$ | RMS $\text{cm}^{-1}$ | STD $\text{cm}^{-1}$ |
|-------|---------------------------------------------|-------------|----------|----------------------|----------------------|
| 4     |                                             | 123         | 0.       | 8.29                 | 3.71                 |
| 4     | O3, $S'$                                    | 151         | 0.       | 6.77                 | 3.16                 |
| 4     | O3, $S''$                                   | 151         | 0.       | 8.08                 | 3.62                 |
| 4     |                                             | 123         | 1/3      | 4.58                 | 2.13                 |
| 4     | O2, $S'$                                    | 128         | 1/3      | 3.98                 | 1.90                 |
| 4     | O2, $S''$                                   | 128         | 1/3      | 4.27                 | 2.05                 |
| 4     | O3, $S'$                                    | 151         | 1/3      | 3.48                 | 1.70                 |
| 4     | O3, $S''$                                   | 151         | 1/3      | 3.89                 | 1.91                 |
| 5     |                                             | 406         | 0.       | 1.96                 | 0.79                 |
| 5     | O3, $S'$                                    | 434         | 0.       | 1.57                 | 0.68                 |
| 5     | O3, $S''$                                   | 434         | 0.       | 1.66                 | 0.68                 |
| 5     |                                             | 406         | 1/3      | 0.94                 | 0.43                 |
| 5     | O3, $S'$                                    | 434         | 1/3      | 0.58                 | 0.28                 |
| 5     | O3, $S''$                                   | 434         | 1/3      | 0.75                 | 0.34                 |

The comparison of the observed fundamental band centers of  $^{32}\text{SF}_6$  with our variational nuclear motion calculations using *ab initio* CCSD(T)/VTZ and CCSD(T)/CVQZ-DK surfaces is given in Table 4. For the VTZ surface defined on a large grid, vibrational and rotational energies were computed using TENSOR variational code in normal mode ITO representation. Technical details can be found in the previous works devoted to five- and six-atomic molecules.<sup>25,32–34</sup> For the VQZ surface defined on a small grid, we used an incremental scheme accounting for the relative difference  $\omega_i(\text{VQZ}) - \omega_i(\text{VTZ})$  between the harmonic frequencies in these basis sets as described below.

The five of six fundamentals calculated with the CVQZ-DK basis set accounting for relativistic contributions show in average better agreement with observed frequencies than VTZ calculations. Note that the most significant contribution in the increase of the fundamental frequencies of angular modes  $\nu_4$ ,  $\nu_5$ , and  $\nu_6$  when using the CVQZ-DK PES in comparison with the VTZ PES is due to the decrease of the equilibrium S–F bond length. A rough estimation of the deviations between VQZ calculations and the observed fundamental band centers (as given in column 4 of Table 4) was obtained using approximate relation  $\nu_i(\text{VQZ}) = \nu_i(\text{VTZ}) + (\omega_i(\text{VQZ}) - \omega_i(\text{VTZ}))$ , where  $\omega_i$  are harmonic frequencies.

The S–F bond length computed in CVQZ-DK basis set approached toward the empirical value. This could simplify in future an empirical optimization of the PES parameters. However, in order to improve the RMS deviation of the 15-dimensional PES fit with a rather small number of parameters, the slightly increased value  $r_e = 1.5613$  (instead of 1.557 for CVQZ-DK) was used. The final *ab initio* PES of this work was constructed using the least-squares fit of the  $F_p$  parameters of the 5-th order expansion (eq 9) to CCSD(T)/VTZ electronic energies at 8708 geometrical nuclear configurations with the CCSD(T)/CVQZ-DK correction, as described above. In order to improve the fit accuracy near the bottom of the PES, the weighting function (eq 13) with the parameter  $E_1 = 5000 \text{ cm}^{-1}$  was applied. This permitted obtaining RMS deviation of  $0.5 \text{ cm}^{-1}$  for 4283 geometries in the range of  $5000 \text{ cm}^{-1}$  above the potential minimum. Because the order of the PES was not very high (in comparison with best methane calculations<sup>21</sup>), the MP2/VTZ electronic energies were not explicitly included in

**Table 4. Fundamental Band Centers ( $\text{cm}^{-1}$ ) for  $^{32}\text{SF}_6$ : Comparison of Purely *Ab Initio* Calculations Using the VTZ and CVQZ-DK PESs with Experimental Values**

|         | Obs      | Calc PES VTZ    | Obs-Calc VTZ | Obs-Calc PES VQZ <sup>a</sup> | Calc PES CVQZ-DK | Obs-Calc CVQZ-DK |
|---------|----------|-----------------|--------------|-------------------------------|------------------|------------------|
|         |          | $r_e = 1.57015$ |              | $r_e = 1.57015$               | $r_e = 1.5613$   |                  |
| $\nu_1$ | 774.5445 | 780.814         | 5.842        | 13.717                        | 773.178          | 1.365            |
| $\nu_2$ | 643.3736 | 662.137         | -12.113      | -1.070                        | 646.928          | -3.5546          |
| $\nu_3$ | 947.9763 | 982.123         | -25.218      | -5.373                        | 958.800          | -10.824          |
| $\nu_4$ | 614.9819 | 613.344         | 5.353        | 11.756                        | 620.836          | -5.854           |
| $\nu_5$ | 523.449  | 519.258         | 8.766        | 12.495                        | 525.897          | -2.448           |
| $\nu_6$ | 347.985  | 347.329         | 3.606        | 7.479                         | 351.724          | -3.739           |

<sup>a</sup>These VQZ values in the fourth column were obtained from full dimensional variational nuclear motion calculation at the VTZ level corrected by the difference  $\omega_i(\text{VQZ}) - \omega_i(\text{VTZ})$  between the harmonic frequencies. The values of  $r_e$  are given in Å.

the fit but were used to check the PES behavior for larger nuclear displacements.

#### 4. AB INITIO DMS FIT

Several studies<sup>8–12</sup> have been devoted to line-by-line analyses of experimental Raman and infrared  $\text{SF}_6$  spectra using effective Hamiltonian and effective transition moments models. An accurate determination of empirical line intensities of heavy molecules with low frequency modes is an extremely difficult task because of crowded spectra and blended lines. One of the most challenging issues is a separation of contributions of overlapping cold and hot bands with a significant impact on the uncertainty in the temperature dependence of the absorption or emission of radiation.

Advanced *ab initio* calculations of molecular PESs and DMS together with the development of variational methods for nuclear motion calculations (refs<sup>18,61–66</sup> and references therein) permitted constructing quite complete lines lists for triatomic,<sup>61,67–70</sup> four-atomic,<sup>17,31,71–73</sup> five-atomic,<sup>19,22,30,34,74,75</sup> and six-atomic<sup>23–25,76</sup> molecules. In some of these cases, the accuracy of *ab initio* line intensities was competitive<sup>77–80</sup> with best empirical determinations being more complete and reliable for the hot bands.<sup>7,68,73,81–83</sup> On the other hand, knowledge of the shape of *ab initio* surfaces<sup>84,85</sup> helped understanding issues related to molecular dynamics.

One more atom in a molecule produces three supplementary degrees of freedom in the nuclear motion resulting in a very substantial computational effort for *ab initio* predictions of high-resolution spectra for polyatomic molecules. In a brute-force approach (without symmetry considerations), there would be roughly a factor  $\sim N^6$  of dimensionality between global spectra calculations of five-atomic molecule (like methane) and seven-atomic molecule such as  $\text{SF}_6$ , where  $N$  is the number of *ab initio* grid points or of basis set function for each degree of freedom.

Here, we present the first results for the full 15-dimensional  $\text{SF}_6$  DMS components and for intensity calculations in the range of the strongest bands. A full account of symmetry permitted drastically reducing the number of redundant geometries and the number of independent DMS parameters. For the construction of the DMS, we used the corresponding irreducible tensor representation similar to the analytical expansions (eqs 9 and 10) where the symmetrized internal coordinates were replaced by normal mode coordinates. The corresponding procedure has been described in our previous works<sup>71,86</sup> devoted to four-atomic and five-atomic molecules. In these works, the *ab initio* DMSs were first built in molecular-fixes frames corresponding to the axes choice options of the MOLPRO code. Then, the axes were rotated to the Eckart

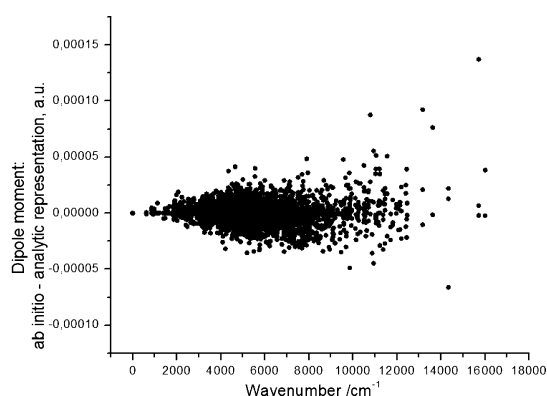
system. As in the case of  $\text{SF}_6$ , the relations between internal and normal coordinated  $Q_i$  were found via numerical transformation, and it was easier to proceed by direct DMS fits in normal coordinates.

A progress in methane *ab initio* calculations has taken a couple of decades.<sup>21,60,81,87,88</sup> The line intensity accuracy in infra-red of 1–5% has been recently reported<sup>80</sup> using the coupled-cluster method with ACVSZ one-particle basis sets. In the case of  $\text{SF}_6$ , such calculations are presently too demanding. In this work, we used CCSD(T) method with two basis sets for the dipole moment calculations. The calculations using VTZ basis were conducted on the first grid containing 1664 nuclear geometries, which was a sub-set of the 6-th order grid relative to PES computations described above. This was then supplemented by AVTZ calculations at 500 geometries. In both cases, the optimal samples of geometries near the bottom of the PES were determined. At each geometry, three projections of the dipole moment were obtained. Dipole moments were computed as the derivative of the energy with respect to the weak external uniform electric field using the finite difference scheme with a field variation of 0.0002 a.u. around at the zero field strength.

For the DMS construction, we used here the scheme, which has been employed in our previous works for PES and DMS of four-to-six atomic molecules.<sup>22,24,71,86</sup> After having computed the dipole moments with AVTZ and VTZ basis sets, we fitted the corresponding 15-dimensional function  $\text{diff}_{2,03}(Q_i)$ , which describes the corresponding difference  $\text{DMS}(\text{AVTZ}) - \text{DMS}(\text{VTZ})$ . This function was modeled by the third-order ITO expansion containing 61 adjustable parameters, of which 50 parameters were statistically well defined. Then, for supplementary geometries from  $N = 501$  to  $N = 1664$ , the AVTZ calculations were lacking, and the AVTZ dipole moments were evaluated as VTZ dipole moments plus  $\text{diff}_{2,03}(Q_i)$ . Finally, the forth-order parameters of the DMS power series expansion in normal coordinated were fitted to the above defined *ab initio* dipole moment values using weight function (eq 14) with  $E_1 = 10,000 \text{ cm}^{-1}$ . Of 260 fourth-order DMS parameters, 194 ones were statistically well determined. This procedure was conducted to high accuracy of the fit corresponding to the RMS deviation of  $1.4 \times 10^{-5}$  a.u. and the weighted standard deviation of  $9.9 \times 10^{-6}$ . The distribution of the fit discrepancies is shown in Figure 4.

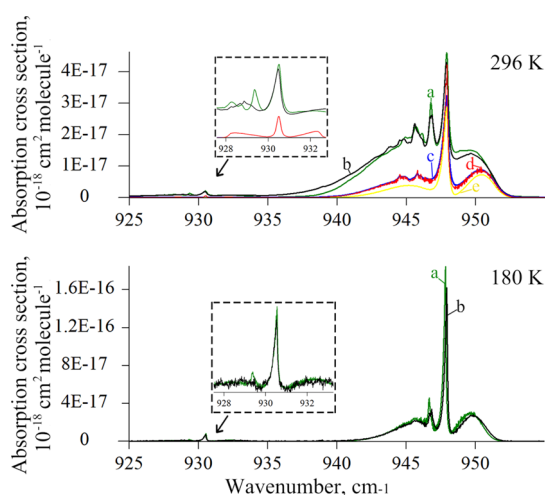
#### 5. RESULTS AND CONCLUSIONS

Preliminary evaluation of  $\text{SF}_6$  band intensities using algorithms of variational methods described in refs<sup>31–34</sup> permits to conclude that our PES and DMS provided realistic predictions of absorption spectra, at least in the range of strong bands. The



**Figure 4.** Deviations between fitted analytical representation of the DMS and *ab initio* dipole moment values versus energies at the corresponding nuclear geometries.

comparison of our calculated absorption cross sections with experimental  $\text{SF}_6$  cross section<sup>13,89</sup> shows a very good agreement at cold temperature of 180 K, see Figure 5, lower



**Figure 5.** Comparison of cross section at a temperature of 296 K (upper panel) and of 180 K (lower panel). Upper panel: our simulation from line list of TheoReTS (green—a), the experimentally observed PNNL spectrum available in HITRAN (black—b), simulation from line list of GEISA 2015 (blue—c), the theoretical calculation of Dijon (red—d), and simulation from line list of HITRAN 2016 database (yellow—e). Lower panel: our simulation from line list of TheoReTS (green—a) and the experimental observed spectrum from HITRAN (black—b).

panel. Note that calculated band centers of  $^{32}\text{SF}_6$  and  $^{34}\text{SF}_6$  isotopologues in Figure 5 were shifted to empirical values (964.196 and 932.534  $\text{cm}^{-1}$  correspondingly) available at <http://vamdc.icb.cnrs.fr/PHP/SF6.php>. An agreement of our calculations with experimental absorption cross sections of Sharpe et al.<sup>90</sup> at room temperature is also satisfactory (Figure 5, upper panel). The latter ones have been recorded at  $T = 298$  K with an  $\text{N}_2$  pressure of 760 Torr with a medium resolution of 0.112  $\text{cm}^{-1}$  and included in HITRAN-on-line cross-section compilation.<sup>13,90,91</sup> The absorption feature near 930  $\text{cm}^{-1}$  corresponds to the  $\nu_3$  band of the  $^{34}\text{SF}_6$  isotopologue at natural abundance. Our *ab initio* cross sections agree much better with these observations than the simulations using  $\text{SF}_6$  line-by-line lists constructed from effective models as available in HITRAN\_2016,<sup>13</sup> GEISA\_2015<sup>14</sup> and Dijon (SheCaS-

Da)<sup>92</sup> databases as shown in Figure 5, upper panel. However, *ab initio* room-temperature simulations appear to be slightly less accurate than for 180 K (Figure 5, lower panel).

Calculations at room temperature require an account of more significant contributions of hot bands involving higher vibrational states. The preliminary data obtained in this work will be made available through the TheoReTS<sup>66</sup> (<http://theorets.univ-reims.fr>, <http://theorets.tsu.ru>) information system. A further improvement could be possibly obtained using larger vibrational basis set that is computationally very demanding. We plan to extend the computation of full line lists in future studies..

## ■ ASSOCIATED CONTENT

### Supporting Information

The Supporting Information is available free of charge at <https://pubs.acs.org/doi/10.1021/acs.jpca.0c02733>.

PES in the internal coordinates and DMS in normal mode expansion in an easy-to-use explicit form for  $^{32}\text{SF}_6$  (ZIP)

Short description of C program for calculation potential energy values (PDF)

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### Notes

The authors declare no competing financial interest.

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Towards a complete elucidation of the ro-vibrational band structure in the SF<sub>6</sub> infrared spectrum  
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# Towards a complete elucidation of the ro-vibrational band structure in the SF<sub>6</sub> infrared spectrum from full quantum-mechanical calculations†

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The first accurate and complete theoretical room-temperature rotationally resolved spectra in the range 300–3000 cm<sup>−1</sup> are reported for the three most abundant isotopologues (<sup>32</sup>SF<sub>6</sub>, <sup>33</sup>SF<sub>6</sub> and <sup>34</sup>SF<sub>6</sub>) of the sulfur hexafluoride molecule. The literature reports that SF<sub>6</sub> is widely used as a prototype molecule for studying the multi-photon excitation processes with powerful lasers in the infrared range. On the other hand, SF<sub>6</sub> is an important greenhouse molecule with a very long lifetime in the atmosphere. Because of relatively low vibrational frequencies, the hot bands of this molecule contribute significantly to the absorption infrared spectra even at room temperature. This makes the calculation of complete rovibrational line lists required for fully converged opacity modeling extremely demanding. In order to reduce the computational costs, symmetry was exploited at all stages of the first global variational nuclear motion calculations by means of irreducible tensor operators. More than 2600 new vibrational band centers were predicted using our empirically refined *ab initio* potential energy surface. Highly excited rotational states were calculated up to  $J = 121$ , resulting in 6 billion transitions computed from an *ab initio* dipole moment surface and distributed over more than 500 cold and hot bands. The final line lists are made available through the TheoReTS information system (<http://theorets.univ-reims.fr>, <http://theorets.tsu.ru>). For the first time, the major (ro)vibrational band structures in the wavenumber range corresponding to the strongest absorption in the infra-red are completely elucidated for a seven-atom molecule, providing excellent agreement with the observed spectral patterns. It is shown that the obtained results are more complete than all available line lists, permitting reliable modelling of the temperature dependence of the molecular opacity.

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## 1 Introduction

SF<sub>6</sub> is a very important greenhouse gas with a global warming potential ~29 000 times larger than that of CO<sub>2</sub> and its atmospheric lifetime can last up to 3200 years<sup>1,2</sup> since it is chemically inert. SF<sub>6</sub> is an important prototype molecule for infrared laser chemistry and laser isotope separation and for ultrahigh resolution spectroscopy. In particular, various problems related to the laser excitation of SF<sub>6</sub> in the quasi-continuum band formed by congested spectral patterns have been discussed.<sup>3–5</sup> As was reviewed in ref. 5, a study of the literature indicated that,

in many cases, the experimental data related with multiple-photon excitation differed by factors of 2 or 3 in absolute value. This makes a more profound insight in the ro-vibrational patterns of levels and transitions of SF<sub>6</sub> in the infrared range. This molecule is mainly used as a gaseous dielectric insulating medium in the electrical industry<sup>6</sup> as well as an atmospheric<sup>7,8</sup> and ocean<sup>9</sup> tracer since its concentration is rapidly increasing in the atmosphere. Vertical profiles have been retrieved by Rinsland *et al.*<sup>10–12</sup> from a series of spectroscopic measurements. Though detected as a trace species, this gas plays an important role in the climate change – as recognized by the Kyoto protocol – making the need to have consistent theoretical models to monitor emissions.

SF<sub>6</sub> belongs to the O<sub>h</sub> point group at the equilibrium structure and possesses 15 vibrational degrees of freedom. Due to its high octahedral symmetry, the molecule has doubly and triply degenerate vibrational modes of E and F-types. Among the 6 fundamental vibrations, three stretching modes ( $\nu_1, \nu_2, \nu_3$ ) transform according

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**Table 1** *Ab initio* (CCSD(T)/CVQZ(DK))<sup>17</sup> and optimized (this work, TW) symmetrized force constants  $F_{ij} = \partial^2 V / \partial S_{ij}^{(I)} \partial S_{ij}^{(J)}$  (in mdyne, 1 dyn = 10<sup>-5</sup> N) and equilibrium geometry  $r_e$  (in Å, 1 Å = 100 pm) for SF<sub>6</sub>. The harmonic wavenumbers  $\omega_i$  (in cm<sup>-1</sup>) are given for <sup>32</sup>SF<sub>6</sub>. The values in the column “best” correspond to a final empirical optimization (see text) and were used to compute all energy levels. Here  $\omega^{\text{opt}}$  and  $\omega^{\text{best}}$  are the same

|              | <i>Ab initio</i> <sup>17</sup> | Opt. (TW) | “Best” (TW) |
|--------------|--------------------------------|-----------|-------------|
| $F_{11}$     | 0.224671                       | 0.220313  | 0.220314    |
| $F_{22}$     | 0.151637                       | 0.151728  | 0.151731    |
| $F_{33}$     | 0.172397                       | 0.169213  | 0.169213    |
| $F_{44}$     | 0.306367                       | 0.295676  | 0.295672    |
| $F_{34}$     | 0.174623                       | 0.163428  | 0.163423    |
| $F_{55}$     | 0.226574                       | 0.217858  | 0.217858    |
| $F_{66}$     | 0.200283                       | 0.192309  | 0.192034    |
| $r_e$        | 1.557027                       | 1.555953  | 1.555938    |
| $\omega_1$   | 794.8578                       | 787.1117  | —           |
| $\omega_2$   | 653.0081                       | 653.2106  | —           |
| $\omega_3^a$ | 964.7491                       | 963.7003  | —           |
| $\omega_4^a$ | 633.3353                       | 621.0293  | —           |
| $\omega_5$   | 539.9829                       | 529.5244  | —           |
| $\omega_6$   | 358.9892                       | 351.5388  | —           |

<sup>a</sup> ( $\omega_3^{\text{opt}}$ ,  $\omega_4^{\text{opt}}$ ) = (954.3495, 619.6474) and (945.5481, 618.2594) cm<sup>-1</sup> for <sup>33</sup>SF<sub>6</sub> and <sup>34</sup>SF<sub>6</sub>, respectively. The other  $\omega$ 's remain unchanged during isotopic substitutions  $^x\text{SF}_6 \rightarrow ^y\text{SF}_6$ .

to the A<sub>1g</sub>, E<sub>g</sub> and F<sub>1u</sub> irreducible representations and three bending modes ( $\nu_4, \nu_5, \nu_6$ ) according to F<sub>1u</sub>, F<sub>2g</sub> and F<sub>2u</sub>. The g type modes are Raman active while  $\nu_3$  and  $\nu_4$  are infrared (IR) active; only  $\nu_6$  is inactive. A brief inspection of the harmonic vibrational modes (see Table 1) shows no simple relationships between  $\omega_i$ . Consequently, the SF<sub>6</sub> spectrum does not exhibit a clear polyad structure, contrary to many other spherical top molecules like methane.<sup>13</sup> Moreover, the presence of relatively low degenerate vibrational modes, combined with a small rotational constant ( $B_e \approx 0.09$  cm<sup>-1</sup>), makes the rotationally resolved IR spectrum very dense and crowded, even at room temperature (RT). This particular feature gives rise to a tremendous number of degenerate and quasi-degenerate states which strongly complicate analysis of high-resolution spectra. Indeed, it can be shown that the density of lines in SF<sub>6</sub> spectra at  $T = 296$  K is similar to that of small hydrocarbons like CH<sub>4</sub>, C<sub>2</sub>H<sub>2</sub>, and C<sub>2</sub>H<sub>4</sub><sup>14–16</sup> at much higher temperatures ( $T \approx 700$ –1500 K). The strongest absorption region corresponds to the stretching fundamental IR band  $\nu_3$ (F<sub>1u</sub>) detectable by remote-sensing techniques,<sup>7,10</sup> which is located in the atmospheric windows around 10.5  $\mu\text{m}$ . However, the presence of low-lying vibrational energy levels makes the Boltzmann population of the excited states very important at RT leading to a substantial number of hot band (HB) transitions that significantly contribute to the overall opacity. For example, the HB transitions  $\nu_k + \sum n_i \nu_i - \sum n_i \nu_i$  ( $k = 3$  or  $i \neq k$ ) falling in the two most congested  $\nu_3$  and  $\nu_4$  regions represent 66 and 70%, respectively, of the integrated intensity at RT. As expected, the “forbidden” mode  $\nu_6$  of lowest frequency is involved in most of the HB transitions ( $n_6 = 1, 2, 3, \dots$ ) allowed by symmetry.

This paper is devoted to the construction of the first comprehensive, accurate and complete theoretical spectrum in the range 300–3000 cm<sup>-1</sup> from variational nuclear motion calculations. To achieve this end, our recent full-dimensional *ab initio* potential

energy (PES) and dipole moment (DMS) surfaces<sup>17</sup> were used. The quadratic force constants of the PES were empirically optimized for this work. The final aim is to fill the gap between what is available in the current spectroscopic databanks and what is really needed to properly model spectra, in particular in the 10  $\mu\text{m}$  windows of atmospheric interest. The paper is structured as follows. The state-of-the art of SF<sub>6</sub> infrared spectroscopy as well as the main motivation for using global variational calculations is presented in Section 2. Section 3 is devoted to the construction of the Hamiltonian model, the choice of basis sets and the computation of ro-vibrational energy levels. The full calculated line lists for the three most abundant species of the sulfur hexafluoride molecule, namely <sup>32</sup>SF<sub>6</sub>, <sup>33</sup>SF<sub>6</sub> and <sup>34</sup>SF<sub>6</sub>, as well as validation on raw experimental spectra are presented in Section 4.

## 2 Effective versus global variational approach

Undoubtedly, the biggest challenge concerns the modelling of all HBs, in particular those falling in the spectral windows of atmospheric interest. To this end, highly excited ro-vibrational states have to be properly modelled to predict all necessary transitions belonging to cold and hot bands, even under RT conditions. In the past four decades, a series of experimental and theoretical studies were devoted to this task to improve analysis and spectral predictions (see *e.g.* ref. 18–21). The intensities of the infrared-active fundamental vibrations have been also studied.<sup>22,23</sup> Recently, much effort has been made to analyse high-resolution SF<sub>6</sub> spectra using phenomenological effective Hamiltonians whose parameters were empirically adjusted to experimental data.<sup>24–29</sup> These data have been included in line list compilations like HITRAN<sup>30</sup> or GEISA.<sup>31</sup> The most extensive empirical model to date has been constructed for the main isotopologue <sup>32</sup>SF<sub>6</sub> by the Dijon group for modelling the band system<sup>28</sup> ( $\nu_1, \nu_2, \nu_3, \nu_1 + \nu_3, \nu_2 + \nu_3, 2\nu_3, \nu_1 + \nu_2 + \nu_3, 3\nu_3, \nu_3 - \nu_1, \nu_3 - \nu_2$ ), recently updated with the set<sup>32</sup> ( $\nu_1 + \nu_4, \nu_2 + \nu_6, \nu_5 + \nu_6$ ). All these combination bands and overtones are of key importance for computing some HBs falling in the strongly absorbing  $\nu_3$  and  $\nu_4$  regions. However, a purely empirical approach based on effective Hamiltonians faces several conceptual difficulties with the increasing number of atoms. First, a new analysis has to be conducted for each isotopologue, unlike global variational calculations for which all species can be predicted almost simultaneously (see *e.g.* ref. 34). Second, effective spectroscopic models are based on a line-by-line assignment procedure that suffers from the completeness issues in spectral analyses of heavy atmospheric molecules like SF<sub>6</sub> because of the presence of very few isolated lines in the highly congested spectra. The lack of assignments as well as ambiguities in the fits of the effective models could often lead to a poor determination of some fitted resonance coupling and empirical dipole transition moment parameters. In such cases, the ro-vibrational effective wavefunctions are not reliably determined and do not permit a correct description of the intensity borrowing effects induced by accidental resonances. In order to

partly get around the problem, the most recent analyses have been realized in a pragmatic way: all experimental spectra have been cooled down ( $\sim 150$  K). This permitted diminishing the density of lines by removing high- $J$  transitions as well as many HBs. In turn, a large amount of information concerning HBs is clearly missing at 296 K. For this reason, the recent line list in the SHECaSDa SF<sub>6</sub> database<sup>32,33</sup> as well as the closely related HITRAN<sup>30</sup> and GEISA<sup>31</sup> compilations does not contain enough line parameters to reproduce and explain the vibrational band structure at 296 K, in particular around  $\nu_3$ . Completeness of rovibrational bands will be thus the key requirement for a successful modelling of the radiative transfer.

To analyse rotationally resolved SF<sub>6</sub> absorption spectra in the infrared, it is imperative to elucidate the vibrational structure first. As explained above, all the previous studies have been done essentially by using the empirical approach with *ad hoc* effective models which clearly showed their limits. An important drawback of a purely empirical approach was that no systematic line intensity measurements have been carried out so far for this molecule because of the difficulty in extracting non-blended lines among the highly congested spectral patterns. In this context, the present paper reports for the very first time full-dimensional, accurate and complete variational calculations in the range 300–3000 cm<sup>−1</sup> – both for line positions and line intensities – for the three most abundant species of the sulfur hexafluoride molecule, namely <sup>32</sup>SF<sub>6</sub> ( $\sim 95\%$ ), <sup>34</sup>SF<sub>6</sub> ( $\sim 4\%$ ) and <sup>33</sup>SF<sub>6</sub> ( $\sim 0.7\%$ ). Though they may suffer from convergence issues concerning line positions, variational global calculations proved to be very efficient in accurately predicting line intensities for both cold and hot bands.<sup>14,16,35–39</sup> The importance of anharmonicity in the vibrational spectra of polyatomic molecules has been studied in many previous works (see *e.g.* ref. 40–45). The focus of our study is different. Here, the aim was to produce complete line lists by (i) including line-by-line intensities for a variationally converged set of spectral transitions at medium or high spectral resolution, (ii) including all hot bands and (iii) providing the temperature dependence of the global opacity. As discussed in a recent review on theoretical rotationally resolved spectral predictions,<sup>38</sup> only a few groups<sup>46–48</sup> (and see references therein) have recently developed specific tools for such applications, but for smaller molecules. In this paper, we focus on the main results permitting one to explain the complicated features in the observed SF<sub>6</sub> spectra. Technical details about the extension of the methods previously developed for smaller molecules<sup>35,49–52</sup> are beyond the scope of this paper.

### 3 Quantum-mechanical nuclear motion calculations

#### 3.1 Refined potential energy surface

As a starting point, the full-dimensional *ab initio* potential and dipole moment surfaces have been computed at the CCSD(T)/CVQZ level of theory (including Douglas-Kroll relativistic corrections), as reported in ref. 17. However, the harmonic frequencies computed from the pure *ab initio* force constant matrix  $F^{ab}$  of the PES,

together with the geometry-based  $G$  Wilson matrix, were off by about  $\sim 1$  to 10 cm<sup>−1</sup>. The *ab initio* equilibrium geometry  $r_e^{ab}$  was not accurate enough to reproduce the rotational structure. In order to better match experimental fundamental band centers, the approach proposed in ref. 35 was adopted to refine the initial solutions  $\{F^{ab}, r_e^{ab}\}$ . A set of optimized quadratic force constants  $F^{\text{opt}}$  for each symmetry block as well as a refined value  $r_e^{\text{opt}}$  were thus obtained (see Table 1) while the high-order *ab initio* force constants were kept unchanged and fixed at the values of ref. 17. The axis convention, the atom numbering as well as the definition of the symmetry coordinates can be found in ref. 17 and 53.

#### 3.2 Hamiltonian model and symmetry

Since SF<sub>6</sub> is a semirigid molecule without large amplitude vibrations, the normal mode coordinates  $q_{k\sigma_k}^{(\Gamma_k)}$  were used to build the Hamiltonian in the Watson–Eckart formalism<sup>54</sup> implemented in the TENSOR computer code. Here,  $\Gamma_k$  denotes an irreducible representation and  $\sigma_k$  a label to distinguish the degenerate components of a vibrational mode  $k$ . Simple analytically solvable potential representations for independent vibrational modes do not permit approaching experimental accuracy in high-resolution spectra. The aim of this work was to account as fully as possible for the high-order anharmonicity of the PES and for the coupling among ro-vibrational states. To this end, we do not use perturbation theory, as often employed in the literature for polyatomic molecules, but the symmetry-adapted variational approach in 18 dimensions: 15 vibrational and 3 rotational degrees of freedom. For a full account of symmetry and to build linearly independent polynomials in  $q$  at a given order of the expansion of the complete nuclear motion Hamiltonian, the use of irreducible tensor operators (ITOs) is an optimal choice. The ITO formalism in the version proposed by Nikitin *et al.*<sup>55</sup> turned out to be very appropriate to describe the coupling of degenerate vibrations. Note that the ITO formalism, initially designed for effective models,<sup>13</sup> is extended here for the first time for global variational nuclear motion calculations of XY<sub>6</sub>-type molecules. Concerning the construction of rotational basis functions adapted to the non-Abelian O<sub>h</sub> point group for very high  $J$  values, the symmetrization procedure proposed by Rey *et al.*<sup>56</sup> was employed. In summary, all the (3+15) rotation-vibrational degrees of freedom were simultaneously taken into account to build the sum-of-product Hamiltonian in the ITO representation – including all intra-molecular coupling terms and anharmonicities at a given order – as well as the direct-product symmetry-adapted basis functions.

In the next step, a “brute-force” 18-dimensional variational calculation of complete rotationally resolved spectra using so many terms in the Hamiltonian up to  $J = 120$  with standard tools and direct eigensolvers was not feasible. To overcome the first obstacle related to the large number of Hamiltonian terms, the reduction technique described in ref. 49, 51 and 52 was applied in a slightly modified form: instead of expressing the reduced tensor Hamiltonian in terms of harmonic oscillator creation and annihilation operators, this latter one was expressed directly as a function of  $q_i^{(\Gamma)}$  and of conjugate momenta  $p_i^{(\Gamma)} = -i\hbar/dq_i^{(\Gamma)}$ . This allowed us to drastically reduce the number of

terms and thus to accelerate the calculation of all matrix elements *via* the Wigner–Eckart theorem. Briefly, the initial  $m$ -th order polynomial expansion of the Eckart–Watson Hamiltonian in  $(q_i^{(r)}, p_i^{(r)})$  was reduced to a transformed polynomial of smaller order  $m' < m$  without loss of accuracy of the eigenvalues. It was then converted to the ITO representation.<sup>35</sup> For this work, both the normal-mode, refined PES and kinetic energy part have been expanded at order  $m = 9$  and reduced at order  $m' = 6$ . A final ITO Hamiltonian  $H^{(9 \rightarrow 6)}$  composed of  $\sim 5600$  non-zero ro-vibrational parameters was built. In the case of SF<sub>6</sub>, such normal mode representation proved to be very efficient and allowed a fast convergence of the Hamiltonian expansion. This type of calculation inherently and systematically accounts for all intra-molecular coupling terms and for anharmonicities through a power series expansion in normal coordinates  $q$  in both the kinetic and potential parts.

### 3.3 Computation of vibrational energy levels

In the next step, we have solved the eigenvalue/eigenfunction problem for the full anharmonic potential with  $J = 0$ . The vibrational energies were obtained by the exact numerical diagonalization of the Hamiltonian matrix built in a direct-product “pruned” basis defined by a set of primitive harmonic oscillator vibrational functions such that<sup>35</sup>

$$F_k(n) \Leftrightarrow \sum_{i=1}^6 \kappa_i v_i \leq n, \quad (1)$$

with  $v_i = 0, \dots, n$ . Here,  $\kappa_i (\geq 1)$  are weight coefficients chosen to select an appropriate number of stretching and bending basis functions. The maximum number of quanta  $v_i^{\max}$  for each mode is given by the integer part  $[n/\kappa_i]$ . For this work,  $n = 10$  and  $\kappa = \{1.05, 1, 1.05, 1, 1, 1\}$  were chosen, resulting in 32 579, 31 263, 63 824, 91 728, 93 029, 29 960, 30 640, 60 582, 94 470 and

93 775 functions for the symmetry blocks  $A_{1g}, A_{2g}, E_g, F_{1g}, F_{2g}$  ( $\tau = g$  and  $u$ ), respectively. For a given block of symmetry  $\Gamma$ , we have thus included in the calculations the set of eigensolutions up to  $\{E_v^{\Gamma(n=10)}, \Psi_v^{\Gamma(n=10)}\}$ . This compressed anharmonic vibrational basis will be used in the next step for rovibrational calculations. A comparison of our variationally computed energies with the empirical levels derived from effective models is given in Table 2 for <sup>32</sup>SF<sub>6</sub>. The error in the fundamental bands is below  $10^{-3} \text{ cm}^{-1}$ . A very good agreement proves the high accuracy of our refined PES, on the one hand, and validates all the theoretical “machinery” for solving the nuclear Schrödinger equation, on the other hand. However, there remain relatively large discrepancies (e.g.  $-0.4 \text{ cm}^{-1}$  for  $2\nu_1 + \nu_3$ ) that can be partly explained by the fact that only the quadratic force constants of the PES have been empirically refined using fundamental band origins. By doing this, we assumed that a large part of the errors was due to uncertainty in the quadratic terms (“harmonic” contribution to the PES). Higher-order terms in the PES also play a substantial role in for the modeling highly excited states but they were fixed at their *ab initio* values in this study. Consequently, though they were improved some overtones and combination bands did not reach the same accuracy as the fundamental bands.

In order to check the basis set convergence of  $J = 0$  energies, we have also computed the  $\Gamma = A_{1g}$  levels by using a larger set of primitive functions with the cut-off  $n = 11$  in eqn (1). In comparison with  $n = 10$ , the convergence errors for 1-, 2-, 3-, 4-, 5- and 6-quanta levels turned out to be 0.0001, 0.0003, 0.006, 0.02, 0.15,  $0.9 \text{ cm}^{-1}$  on average. Clearly, though their contributions to the absorption spectrum are almost negligible, very highly excited vibrational levels are not fully converged as expected. Unfortunately, the use of the  $F_k(11)$  basis set – composed of about twice as many basis functions than  $F_k(10)$  – goes beyond our computer resources to treat all symmetry blocks. Finally, the most extensive list composed of 2686 (in comparison with

**Table 2** Selected vibrational band centers (in  $\text{cm}^{-1}$ ) for <sup>32</sup>SF<sub>6</sub> computed from variational calculations using our refined PES (this work, TW) and from empirical effective models.<sup>20,26–29,57</sup> Calculated band centers for <sup>33</sup>SF<sub>6</sub> and <sup>34</sup>SF<sub>6</sub> as well as the isotopic shifts  $\Delta$  (in  $\text{cm}^{-1}$ ) with respect to <sup>32</sup>SF<sub>6</sub> are also given (see text for explanation)

| Band                                   | <sup>32</sup> SF <sub>6</sub> |                  |           | <sup>33</sup> SF <sub>6</sub> |                             | <sup>34</sup> SF <sub>6</sub> |                             |
|----------------------------------------|-------------------------------|------------------|-----------|-------------------------------|-----------------------------|-------------------------------|-----------------------------|
|                                        | Empirical                     | Variational (TW) | Emp.-Var. | Variational (TW)              | $\Delta(32 \rightarrow 33)$ | Variational (TW)              | $\Delta(32 \rightarrow 34)$ |
| $\nu_6(\text{F}_{2u})$                 | 347.7367                      | 347.7366         | 0.0001    | 347.7504                      | −0.0137                     | 347.7635                      | −0.0269                     |
| $\nu_5(\text{F}_{2g})$                 | 523.4490                      | 523.4494         | −0.0004   | 523.4705                      | −0.0215                     | 523.4906                      | −0.0412                     |
| $\nu_4(\text{F}_{1u})$                 | 614.9819                      | 614.9821         | −0.0002   | 613.6757                      | 1.3062                      | 612.3611                      | 2.6210                      |
| $\nu_2(\text{E}_g)$                    | 643.3737                      | 643.3734         | 0.0003    | 643.3908                      | −0.0171                     | 643.4073                      | −0.0339                     |
| $\nu_1(\text{A}_{1g})$                 | 774.5453                      | 774.5454         | −0.0001   | 774.5721                      | −0.0268                     | 774.5968                      | −0.0514                     |
| $\nu_3(\text{F}_{1u})$                 | 948.1025                      | 948.1024         | 0.0001    | 939.0272                      | 9.0753                      | 930.4646                      | 17.6378                     |
| $\nu_2 + \nu_6(\text{F}_{2u})$         | 989.1190                      | 989.1653         | −0.0463   | 989.2029                      | −0.0839                     | 989.2324                      | −0.0671                     |
| $\nu_2 + \nu_6(\text{F}_{1u})$         | 991.2740                      | 991.3071         | −0.0331   | 991.2641                      | 0.0099                      | 991.2401                      | 0.0670                      |
| $\nu_2 + \nu_4(\text{F}_{1u})$         | 1257.0849                     | 1257.0190        | 0.0659    | 1255.7262                     | 1.3587                      | 1254.4199                     | 2.5991                      |
| $\nu_1 + \nu_4(\text{F}_{1u})$         | 1388.3820                     | 1388.3192        | 0.0628    | 1387.0563                     | 1.3257                      | 1385.7712                     | 2.5480                      |
| $\nu_2 + \nu_3(\text{F}_{1u})$         | 1587.7173                     | 1587.9089        | −0.1916   | 1578.9063                     | 8.8110                      | 1570.3867                     | 17.5222                     |
| $\nu_2 + \nu_3(\text{F}_{2u})$         | 1591.1923                     | 1591.0909        | 0.1014    | 1582.1352                     | 9.0571                      | 1573.6517                     | 17.4392                     |
| $\nu_1 + \nu_3(\text{F}_{1u})$         | 1719.7390                     | 1719.8226        | −0.0836   | 1710.8106                     | 8.9284                      | 1702.2811                     | 17.5415                     |
| $2\nu_3(\text{A}_{1g})$                | 1889.0500                     | 1889.3553        | −0.3053   | 1871.3404                     | 17.7096                     | 1854.3259                     | 35.0294                     |
| $2\nu_3(\text{E}_g)$                   | 1891.6000                     | 1892.0052        | −0.4052   | 1873.9120                     | 17.6880                     | 1856.8083                     | 35.1969                     |
| $2\nu_3(\text{F}_{2g})$                | 1896.5300                     | 1896.5375        | −0.0075   | 1878.4443                     | 18.0857                     | 1861.3746                     | 35.1629                     |
| $\nu_1 + \nu_2 + \nu_4(\text{F}_{1u})$ | 2028.1549                     | 2027.9459        | 0.2090    | 2026.7878                     | 1.3671                      | 2025.5056                     | 2.4403                      |
| $2\nu_1 + \nu_4(\text{F}_{1u})$        | 2160.0257                     | 2159.9263        | 0.0994    | 2158.3789                     | 1.6468                      | 2158.2626                     | 1.6637                      |
| $2\nu_1 + \nu_3(\text{F}_{1u})$        | 2489.4682                     | 2489.8765        | −0.4083   | 2480.5044                     | 8.9638                      | 2473.5020                     | 16.3745                     |

~25 ones available in the literature) selected vibrational energy (sub)levels up to  $3000\text{ cm}^{-1}$  is given in the ESI† S<sub>1</sub>. Selected vibrational band centers for  $^{33}\text{SF}_6$  and  $^{34}\text{SF}_6$  as well as the isotopic shifts  $\Delta(32 \rightarrow 33)$  and  $\Delta(32 \rightarrow 34)$  are also given in Table 2. As already noted in ref. 21, “anomalous” negative frequency

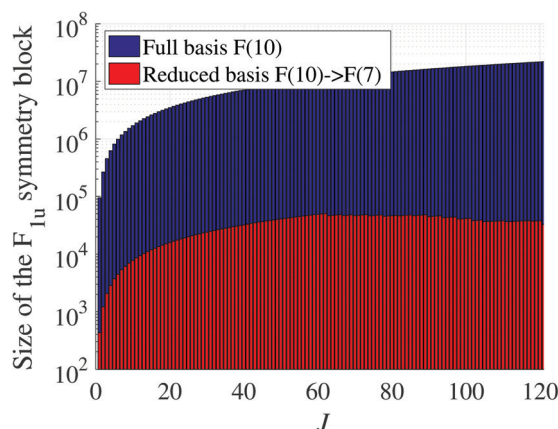


Fig. 1 Dimension of the Hamiltonian matrices to be diagonalized for the symmetry block  $F_{1u}$  as a function of  $J$  when using the full and reduced vibrational eigenfunctions  $\psi^{(10)}$  and  $\psi^{(10 \rightarrow 7)}$ , respectively.

shifts are observed. Such a feature is generally possible when some harmonic frequencies  $\omega_i$  remain unchanged during isotopic substitutions. For  $\text{SF}_6$ , anomalous shifts are thus expected for  $\nu_1$ ,  $\nu_2$ ,  $\nu_5$  and  $\nu_6$ . As stated in ref. 21, “the change in the mixed-mode anharmonicity constants can lower the ground-state energy more than the excited-state energy, resulting in an increase in  $\nu$ ”. The negative shift in  $\nu_1$  between  $^{32}\text{SF}_6$  and  $^{34}\text{SF}_6$  was predicted to be  $-0.044\text{ cm}^{-1}$  in ref. 21 while Volkov *et al.*<sup>58</sup> have observed a shift of  $-0.0574\text{ cm}^{-1}$ . In this work, the shift is given by  $-0.0514\text{ cm}^{-1}$  which is close to the observation. Similar shifts were predicted for  $\nu_2$ ,  $\nu_5$  and  $\nu_6$ , increasing the band origins from 0.025 to  $0.04\text{ cm}^{-1}$ . In this work, the band origins increase by 0.0269 to  $0.0412\text{ cm}^{-1}$ . Note that a similar situation was reported in ref. 59 between  $^{12}\text{CH}_4$  and  $^{13}\text{CH}_4$  for  $\nu_2$  because  $\omega_2$  is unchanged during the substitution  $^{12}\text{C} \rightarrow ^{13}\text{C}$ . Note that recently,  $J = 0$  energy levels were reported<sup>60</sup> for the  $\text{UF}_6$  molecule that belongs to the same  $O_h$  point group, by using a rectangular collocation approach and the product of harmonic basis functions but without exploiting the full symmetry.

### 3.4 Rovibrational energy levels for high- $J$ values

For  $J > 0$ , “brute-force” calculations from the direct-product basis set (1) coupled with symmetric-top wave functions are

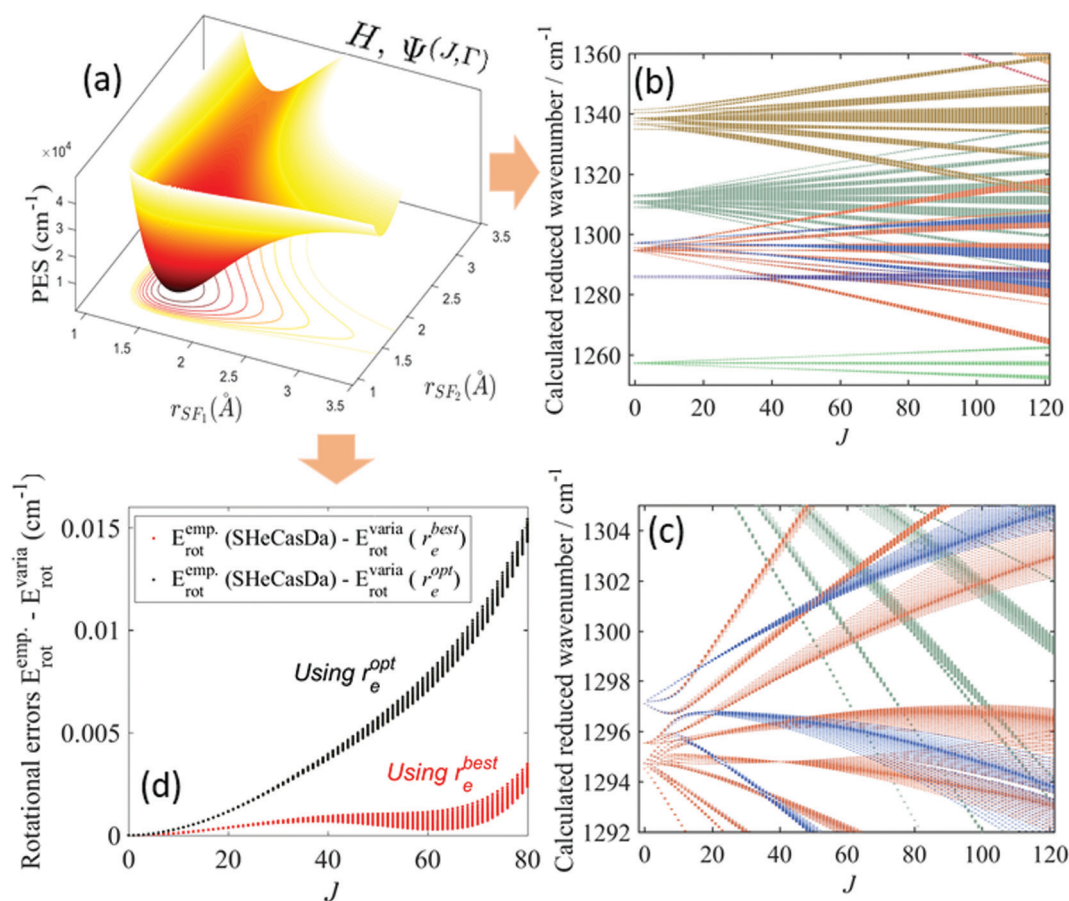


Fig. 2 (a) Cut of the  $\text{SF}_6$  PES<sup>17</sup> along the S–F<sub>1</sub> and S–F<sub>2</sub> bonds. (b) Reduced upper state energy levels for  $\nu_2 + \nu_4$ ,  $2\nu_2$ ,  $\nu_3 + \nu_6$ ,  $\nu_1 + \nu_5$ ,  $\nu_4 + 2\nu_6$  and  $\nu_2 + \nu_6$ , listed in increasing energy order. (c) Portion of reduced upper state energy levels for  $\nu_3 + \nu_6$  and  $\nu_1 + \nu_5$ . (d) Rotational errors using our optimized ( $r_e^{\text{opt}}$ ) and “best” ( $r_e^{\text{best}}$ , see text) equilibrium geometry values.

clearly proscribed using the available direct eigensolvers. Instead of considering iterative eigensolvers, as often recommended,<sup>61</sup> we have followed the method developed in our previous work for smaller molecules.<sup>50</sup> To this end, a set of reduced eigenfunctions  $\Psi_v^{I(n \rightarrow m)}$  associated with a basis  $F_{k'}(m)$  were obtained using the projection technique onto an appropriately chosen smaller subspace with  $m < n$  for which  $J > 0$  calculations are feasible. For this work, rovibrational energy levels were computed up to  $J_{\max} = 121$  using  $m = 7$ . In the last step a final compression of the basis set was applied by selecting  $p\%$  of eigenvectors where  $p = 100$  up to  $J = 61$  and then gradually decreases to  $p = 33$  at  $J = 121$ . This strategy allows one to keep approximately the same memory requirement as  $J$  increases, bearing in mind that accurate calculations are generally not necessary for  $J > 80$  since the corresponding transitions are very weak at room temperature. Note that the  $J_{\max}$  value was chosen to ensure convergence of the partition function  $Q(T)$  at  $T = 296$  K. This gives  $Q(296) = 1607253.7$  for  $^{32}\text{SF}_6$ , within 0.05% of error. To better illustrate the use of the reduction procedure, the maximum dimension of the eigenvector matrix for  $J = 121$  is  $106784 \times 35311$ ; without reduction it would have been  $(23 \times 23)$  millions. Fig. 1 displays the size of the symmetry blocks  $F_{1u}$  as a function of  $J$  when using or not the reduced vibrational eigenfunctions  $\Psi_v^{(10 \rightarrow 7)}$ . In this figure (see the red part), we clearly see that the number of rovibrational basis functions decreases beyond  $J = 61$  because of reduction of the vibrational subspace which varies from 100% to 33%.

In order to validate the optimized value  $r_e^{\text{opt}}$  for the equilibrium structure (see Table 1), let us first focus on the accuracy of our variationally predicted rotational energy levels  $E_{\text{rot}}^{\text{varia}}$ . To this end, we did a direct comparison with energies  $E_{\text{rot}}^{\text{emp}}$  obtained from the empirical effective Hamiltonian.<sup>24</sup> Fig. 2(d) displays the rotational errors  $\Delta_{\text{rot}} = E_{\text{rot}}^{\text{emp}} - E_{\text{rot}}^{\text{varia}}$  up to  $J = 80$ . We clearly see in this figure that the discrepancy increases with  $J$  because  $r_e^{\text{opt}}$  was optimized only for low- $J$  values by using a small basis set during our refinement procedure. A way to improve the accuracy was to introduce an empirical correction to  $E_{\text{rot}}^{\text{varia}}$  by the term  $\delta_{\text{rot}} = 1.8 \times 10^{-6}(J+1) \text{ cm}^{-1}$ . This correction is equivalent to a shift in the equilibrium geometry by  $\delta_{\text{re}} = -1.5 \times 10^{-5} \text{ \AA}$  to give the best estimated value  $r_e^{\text{best}} = r_e^{\text{opt}} + \delta_{\text{re}} = 1.555938 \text{ \AA}$ . As a direct consequence,  $F^{\text{opt}}$  had to be slightly modified ( $\equiv F^{\text{best}}$ ) to

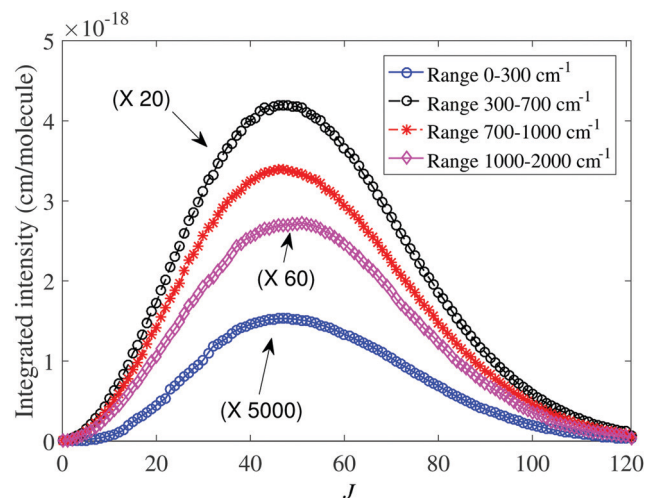


Fig. 4 Integrated intensities (in cm per molecule) for different spectral ranges at different values of the rotational angular momentum  $J$ .

preserve the accuracy of the vibrational levels. Finally, both the Hamiltonian model and energy levels were recalculated using the newly optimized couple  $\{F^{\text{best}}, r_e^{\text{best}}\}$  given in Table 1. This results in the improved set of rotational levels (see Fig. 2(d)) whose uncertainty does not exceed  $0.0035 \text{ cm}^{-1}$  at  $J = 80$ . Above this value, the extrapolation error of the effective model beyond the range of observed data also contributes to  $\Delta_{\text{rot}}$ , which increases up to  $0.1 \text{ cm}^{-1}$  for  $J = 120$ .

Selected sets of computed rovibrational energy levels are displayed in Fig. 2(b and c) in the region of  $\nu_2 + \nu_4$ ,  $2\nu_2$ ,  $\nu_3 + \nu_6$ ,  $\nu_1 + \nu_5$ ,  $\nu_4 + 2\nu_6$  and  $\nu_2 + \nu_6$ , listed in increasing energies. The sequence of quasi-degenerate groups of quantum levels called clusters is well known for  $\text{SF}_6$  and was explained by semi-classical analysis.<sup>62</sup> The different colours give information on how the energy levels are distributed and on how the states are mixed altogether. Except for strong local resonances (e.g. between  $\nu_3 + \nu_6$  and  $\nu_1 + \nu_5$  at  $J = 7$ , see Fig. 2(c)), the inter-band couplings appear to be relatively weak for  $\text{SF}_6$ . As an indicator of the degree of the coupling, the average value of the absolute maximum contribution in the wavefunction decomposition for the first 100 vibrational levels is 0.922, which is close to 1. This may partly explain why the convergence of the variational

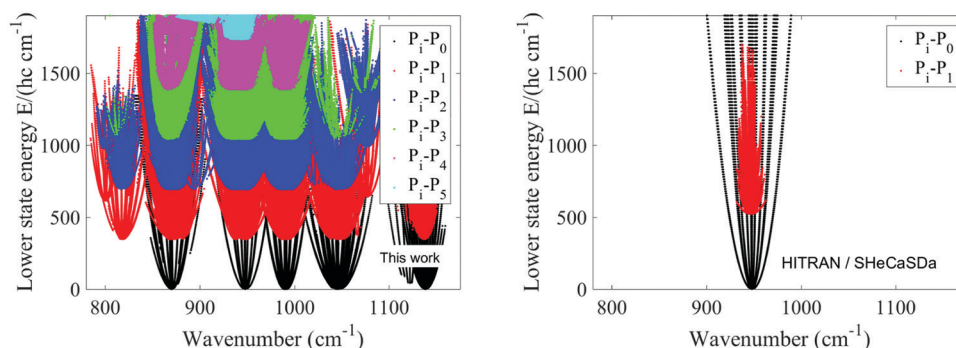


Fig. 3 Comparison of the lower state energies for transitions involved in the TheoReTS database<sup>48</sup> (left panel, this work) and in HITRAN<sup>30</sup>/SHeCaSDa<sup>32</sup> (right panel). The intensity cut-off was fixed at  $10^{-25} \text{ cm per molecule}$ . For the definition of  $P_i - P_j$ , see text.

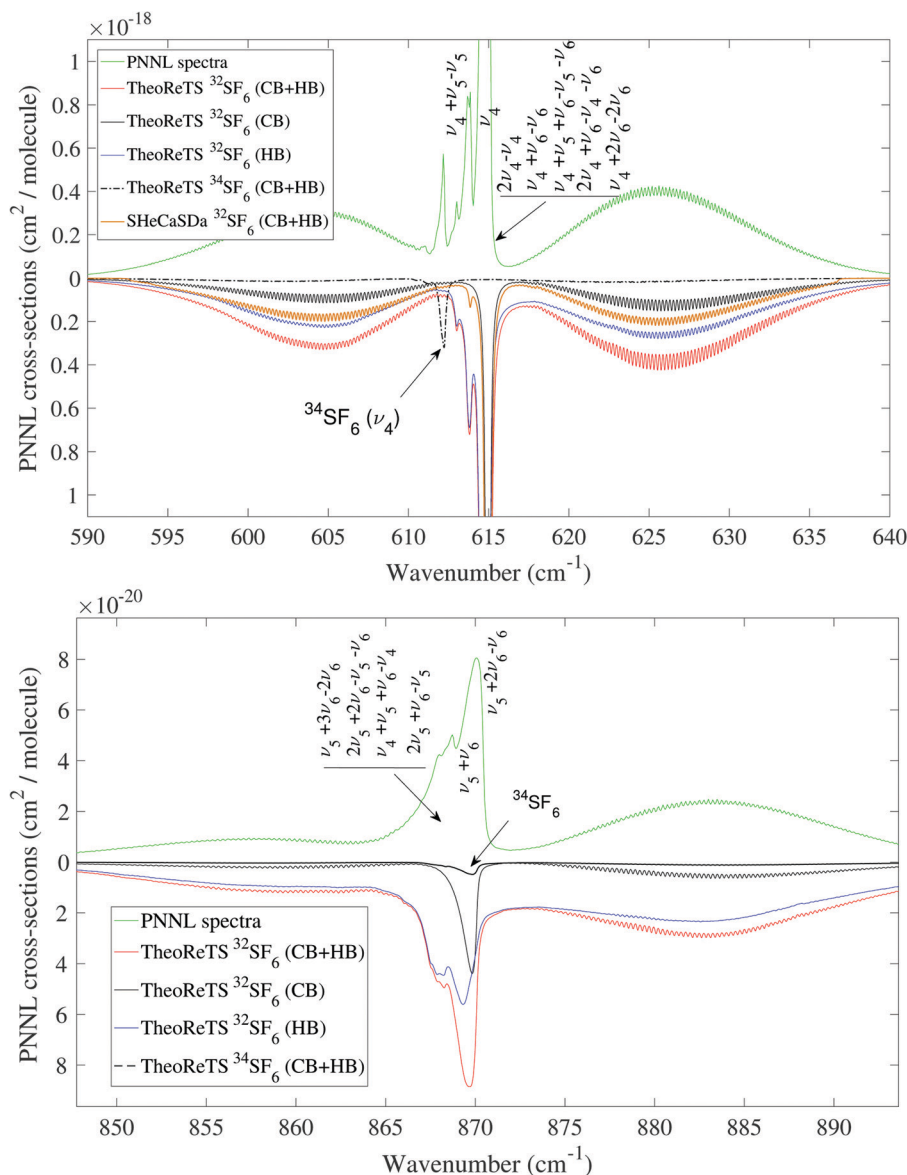
calculation for SF<sub>6</sub> is rather fast compared to other molecules like methane or ethylene.

## 4 Construction of the line lists – validation

### 4.1 Methodology

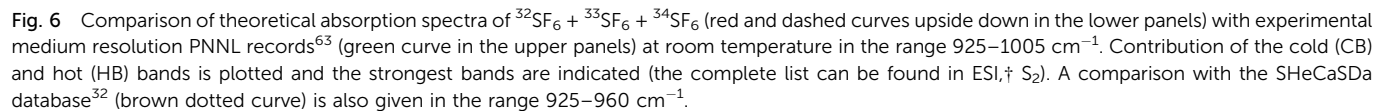
The ultimate aim of this work was the construction of the room-temperature *ab initio* based line lists for <sup>32</sup>SF<sub>6</sub>, <sup>33</sup>SF<sub>6</sub> and <sup>34</sup>SF<sub>6</sub> in the range 300–3000 cm<sup>−1</sup>. Here, we briefly recall the methodology obtained by the extension of the approach developed in our previous works<sup>35,49–52</sup> for smaller molecules.

– *Dipole moment*: the line intensities were computed using the home-made TENSOR code from the fourth order normal-mode DMS reported in ref. 17. In order to be consistent with the Hamiltonian treatment, the dipole moment component  $\mu_z(q)$  which transforms as the irreducible representation F<sub>1u</sub> in the molecular-fixed frame was also converted to ITOs. A set of 213 *ab initio* dipole moment parameters in the ITO representation were thus deduced algebraically and used to compute line intensities. The explicit use of symmetry considerations when building the dipole moment turns out to be of great help to deduce all the symmetry-allowed transitions ( $J', C'$ ) ← ( $J'', C''$ ) via the selection rules  $\Delta J = 0, \pm 1$  and  $C' \times C'' \supset A_{1u}$ . For intensity calculations, the nuclear spin statistical weights



**Fig. 5** Comparison of theoretical absorption spectra of <sup>32</sup>SF<sub>6</sub> + <sup>34</sup>SF<sub>6</sub> (red and dotted curves upside down in the lower panels) with experimental medium resolution PNNL records<sup>63</sup> (green curve in the upper panels) at room temperature in the range 590–895 cm<sup>−1</sup>. Contribution of the cold (CB) and hot (HB) bands is plotted and the strongest bands are indicated (the complete list can be found in ESI,† S<sub>2</sub>). A comparison with the SHECaSDa database<sup>32</sup> (brown curve) is also given in the range 590–640 cm<sup>−1</sup>.

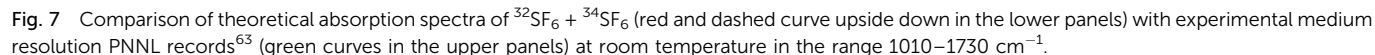
In order to properly converge the integrated intensities using an optimal number of lines, we chose to split the whole wavenumber range  $W$  in three intervals. For each interval of  $^{32}\text{SF}_6$ , a specific intensity cut-off  $I_{\text{cut}}^{\text{max}}$  was defined as follows: for  $300 < W \leq 910 \text{ cm}^{-1}$ ,  $I_{\text{cut}}^{\text{max}} = 5 \times 10^{-29} \text{ cm per molecule}$ ; for  $910 < W \leq 965 \text{ cm}^{-1}$ ,  $I_{\text{cut}}^{\text{max}} = 7 \times 10^{-28} \text{ cm per molecule}$  and  $I_{\text{cut}}^{\text{max}} = 1 \times 10^{-29} \text{ cm per molecule}$  otherwise. It is shown in Fig. 4 that the integrated intensities are nearly converged at  $J_{\text{max}} = 121$



was chosen for  $W \leq 300 \text{ cm}^{-1}$  and  $\alpha = 5000$  otherwise. The remaining set is formed by an accumulation of several million (and even billions in our case) very weak overlapping lines, resulting in a smooth quasi-continuum (QC). Though the contribution of each individual weak line could be very small, the whole set plays a substantial role in the modelling of the overall opacity.

## 4.2 Validation and distribution

Fig. 5–7 give examples of comparison between theoretical (including the CB, HB, and total CB + HB contributions shown in different colors) and experimental PNNL raw spectra<sup>63</sup> which have been recently included in HITRAN as absorption cross-sections without line-by-line assignments. All simulations were made at a resolution of  $0.1\text{ cm}^{-1}$  and a pressure of 760 Torr using a path length of 20 cm and a Voigt profile. The main absorption bands are indicated in these figures and the integrated



**Table 3** Integrated intensities  $\sum S_{ij}$  ( $\times 10^{17}$  cm per molecule) for the strongest absorption bands in the  $\nu_4$  and  $\nu_3$  regions of  $^{32}\text{SF}_6$  and comparison with the SHeCaSDa database<sup>32</sup>

| $\nu_4$ region |                                         |                    |                          | $\nu_3$ region |                                         |                    |                          |
|----------------|-----------------------------------------|--------------------|--------------------------|----------------|-----------------------------------------|--------------------|--------------------------|
| Center         | Band                                    | $\sum S_{ij}$ (TW) | $\sum S_{ij}$ (SHeCaSDa) | Center         | Band                                    | $\sum S_{ij}$ (TW) | $\sum S_{ij}$ (SHeCaSDa) |
| 614.982        | $\nu_4$                                 | 0.369              | 0.394                    | 948.102        | $\nu_3$                                 | 6.16               | 5.53                     |
| 615.727        | $\nu_4 + \nu_6 - \nu_6$                 | 0.203              | 0.214                    | 946.761        | $\nu_3 + \nu_6 - \nu_6$                 | 2.99               | —                        |
| 614.623        | $\nu_4 + \nu_5 - \nu_5$                 | 0.086              | —                        | 945.343        | $\nu_3 + \nu_5 - \nu_5$                 | 1.40               | 1.18                     |
| 614.790        | $\nu_4 + 2\nu_6 - 2\nu_6$               | 0.073              | —                        | 946.088        | $\nu_3 + \nu_4 - \nu_4$                 | 0.901              | —                        |
| 614.658        | $2\nu_4 - \nu_4$                        | 0.073              | —                        | 946.033        | $\nu_3 + 2\nu_6 - 2\nu_6$               | 0.869              | —                        |
| 616.162        | $\nu_4 + \nu_5 + \nu_6 - \nu_5 + \nu_6$ | 0.045              | —                        | 944.536        | $\nu_2 + \nu_3 - \nu_2$                 | 0.517              | 0.452                    |
| 615.694        | $2\nu_4 + \nu_6 - \nu_4 + \nu_6$        | 0.039              | —                        | 944.060        | $\nu_3 + \nu_5 + \nu_6 - \nu_5 + \nu_6$ | 0.500              | —                        |
| 613.939        | $\nu_2 + \nu_4 - \nu_2$                 | 0.032              | —                        | 946.370        | $\nu_3 + \nu_4 + \nu_6 - \nu_4 + \nu_6$ | 0.393              | —                        |
| 614.894        | $\nu_4 + 3\nu_6 - 3\nu_6$               | 0.020              | —                        | 948.435        | $2\nu_3 - \nu_3$                        | 0.231              | —                        |
| 613.633        | $\nu_2 + \nu_4 + \nu_6 - \nu_2 + \nu_6$ | 0.017              | —                        | 948.393        | $\nu_2 + \nu_3 + \nu_6 - \nu_2 + \nu_6$ | 0.188              | —                        |
| 614.355        | $2\nu_4 + \nu_5 - \nu_4 + \nu_5$        | 0.016              | —                        | 942.521        | $\nu_3 + 3\nu_6 - 3\nu_6$               | 0.166              | —                        |

**Table 4** Sum of experimental PNNL cross-sections ( $\text{cm}^2/\text{molecule}$ ) for  $\text{SF}_6$  at room temperature and comparison with simulations using the TheoReTS line list (this work) and the SHeCaSDa database<sup>32</sup>

| Region ( $\text{cm}^{-1}$ ) | TheoReTS (TW)          | PNNL                   | Error (%) |
|-----------------------------|------------------------|------------------------|-----------|
| 560–650                     | $1.98 \times 10^{-16}$ | $2.23 \times 10^{-16}$ | 11.2      |
| 820–920                     | $1.55 \times 10^{-17}$ | $1.63 \times 10^{-17}$ | 4.9       |
| 920–965                     | $3.12 \times 10^{-15}$ | $3.20 \times 10^{-15}$ | 2.3       |
| 965–1005                    | $4.15 \times 10^{-17}$ | $4.01 \times 10^{-17}$ | −3.4      |
| 1005–1500                   | $7.44 \times 10^{-18}$ | $7.62 \times 10^{-18}$ | 2.4       |
| 1550–1730                   | $3.23 \times 10^{-17}$ | $3.47 \times 10^{-17}$ | 6.9       |
| Region ( $\text{cm}^{-1}$ ) | SHeCaSDa <sup>32</sup> | PNNL                   | Error (%) |
| 560–650                     | $1.01 \times 10^{-16}$ | $2.23 \times 10^{-16}$ | 54.7      |
| 910–965                     | $1.19 \times 10^{-15}$ | $3.20 \times 10^{-15}$ | 62.8      |

intensities for the strongest absorption bands in the  $\nu_4$  and  $\nu_3$  regions of  $^{32}\text{SF}_6$  are given in Table 3. Though a good agreement can be seen in Fig. 5–7, plotting residuals between the PNNL and the TheoReTS data is quite a tricky task. On the one hand, the line positions, in particular for hot bands, may be shifted, and on the other hand, a study of PNNL line profile shapes was beyond the scope of this work, thus making a proper comparison difficult. Instead, the accuracy of the calculated patterns can be established by a direct comparison of the absorption cross-sections of  $\text{SF}_6$  in the regions displayed in Fig. 5–7. The results are summarized in Table 4. The agreement between theory and the observed PNNL

spectra in the strongest absorption region around  $947 \text{ cm}^{-1}$  is clearly confirmed as the deviation in the integrated cross-sections is only 2.3%. Assuming the value of the integrated intensity for  $\nu_4$  is consistent in SHeCaSDa (see Table 3), the first half of the 11% of error between PNNL and TheoReTS around  $615 \text{ cm}^{-1}$  can be possibly explained by a deviation (about 2.8%) of our dipole moment first derivative  $\partial\mu/\partial q_4$  and the other half by line strength errors for some hot bands. The experimental uncertainty in weaker absorption intervals due to baseline issues could also significantly contribute to the deviation between PNNL and TheoReTS in Table 4. Overall, to our knowledge, such an agreement in rovibrational absorption patterns of seven-atomic molecules has not been obtained so far in previous theoretical calculations.

Spectra simulations using the SHeCaSDa line lists are also shown in Fig. 5 and 6 for the regions of  $\nu_4$  (around  $600 \text{ cm}^{-1}$ ) and  $\nu_3$  (around  $950 \text{ cm}^{-1}$ ). The latter ones do not reproduce the overall absorption shape of the room-temperature observed spectra due to the lack of HBs, as indicated by Tables 3 and 4 and corroborated by Table 5. Among the 500 cold and hot bands contained in the “strong” subset of our calculated line list of  $^{32}\text{SF}_6$ , the integrated intensities for the 260 strongest ones can be found in ESI† S2. For example, the most important HB  $\nu_3 + \nu_6 - \nu_6$  that represents half of the  $\nu_3$  contribution to the absorption is clearly lacking in the SHeCaSDa and HITRAN line lists (see Table 3).

A detailed description of the final line lists for  $^{32}\text{SF}_6$  is given in Table 5. The number of transitions in the “strong lines

**Table 5** Summary of the final line lists for  $^{32}\text{SF}_6$  and comparison with the SHeCaSDa database.<sup>32</sup> The number of strong and quasi-continuum (QC) lines (#lines) as well as the integrated intensities ( $\sum$ lines) are indicated. Here, CB = cold bands, HB = hot bands

| Range/type      | Strong lines  |                        | QC                  |                        | SHeCaSDa <sup>32</sup> |                        | Deviations <sup>a</sup> |
|-----------------|---------------|------------------------|---------------------|------------------------|------------------------|------------------------|-------------------------|
|                 | #Lines        | $\sum$ lines           | #Lines              | $\sum$ Lines           | #Lines                 | $\sum$ Lines           |                         |
|                 | $\times 10^6$ | $\times 10^{-16}$      | $\times 10^9$       | $\times 10^{-16}$      | $\times 10^5$          | $\times 10^{-16}$      |                         |
| 300–910 CB      | 0.0426        | $3.828 \times 10^{-2}$ | 0.0005              | $4.273 \times 10^{-5}$ | 1.3165                 | $3.942 \times 10^{-2}$ | −2.8%                   |
| HB              | 2.7963        | $7.337 \times 10^{-2}$ | 1.6862              | $1.718 \times 10^{-2}$ | 0.3851                 | $2.201 \times 10^{-2}$ | 75.6%                   |
| 910–965 CB      | 0.0231        | 0.615                  | 0.0002              | $3.170 \times 10^{-4}$ | 0.5310                 | 0.553                  | 10.1%                   |
| HB              | 2.9370        | 1.008                  | 0.9543              | 0.253                  | 1.6568                 | 0.175                  | 86.1%                   |
| 965–3000 CB     | 0.6444        | $1.443 \times 10^{-2}$ | 0.0217              | $2.650 \times 10^{-4}$ | —                      | —                      | —                       |
| HB              | 7.8830        | $2.254 \times 10^{-2}$ | 2.9239              | $1.213 \times 10^{-2}$ | —                      | —                      | —                       |
| Total (CB + HB) | 15.470        | 1.771                  | 5.9991 <sup>b</sup> | 0.283                  | 3.3584                 | 0.789                  | —                       |

<sup>a</sup> Relative deviations between our TheoReTs and SHeCaSDa integrated intensities computed as  $\sum(\text{Strong} + \text{QC} - \text{SHeCaSDa}) / \sum(\text{Strong} + \text{QC}) \times 100$ .

<sup>b</sup> The QC line list was compressed to 1.81 millions of lines using the super-line approach.

subset” represents 0.25% of the full line list file for  $^{32}\text{SF}_6$ , providing 86% of the opacity. The remaining 5.999 billion lines of the QC were compressed using our super-line approach,<sup>14</sup> permitting reduction of the number of lines by 3 orders of magnitude with a minimal loss of accuracy. The final compressed QC line list is composed of 1.81 million super-lines. For theoretical simulation of absorption or emission cross-sections at a given temperature, the super-lines can be treated in the same manner as the true lines contained in the “strong line subset”. A comparison with the recent SHeCaSDa database is also given in this table. As already seen in Fig. 3 and confirmed in Table 4, the lack of opacity in the  $\nu_4$  and  $\nu_3$  regions in SHeCaSDa is due to the lack of 80% of HBs. Both the strong and QC line lists are freely downloadable from the TheoReTS information system<sup>48</sup> (<http://theorets.univ-reims.fr>, <http://theorets.tsu.ru>).

## 5 Conclusion

To summarize, we present in this work the first global theoretical predictions of infrared rotationally resolved spectra for the sulfur hexafluoride molecule. The rotation-vibration energy levels are determined from variational calculations using our refined PES obtained by the optimisation of the *ab initio* one of ref. 17. Most importantly, absolute line intensities are calculated and extensive theoretical molecular line lists are constructed for the three most abundant species  $^{32}\text{SF}_6$ ,  $^{33}\text{SF}_6$  and  $^{34}\text{SF}_6$  using *ab initio* dipole moment functions. Room-temperature calculated spectra match the observed ones with unprecedented accuracy for ro-vibrational patterns of seven-atomic molecules. This is, in particular, the case of the most important range in the atmospheric windows around 10  $\mu\text{m}$  where the error between the sum of experimental absorption cross-sections of PNNL and TheoReTS is 2.3%, which can be compared to the deviation of 63% between PNNL and the line lists obtained from the most recent line-by-line analyses. This work permitted one to elucidate for the first time all complicated cold and hot band structures in the rotationally resolved room-temperature  $\text{SF}_6$  spectra in the strongly absorbing ranges. For applications requiring accurate opacity calculations and its temperature dependence, our complete theoretical line lists should thus be preferred to partial lists included in the HITRAN or SHeCaSDa databases.

## Conflicts of interest

There are no conflicts to declare.

## Acknowledgements

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